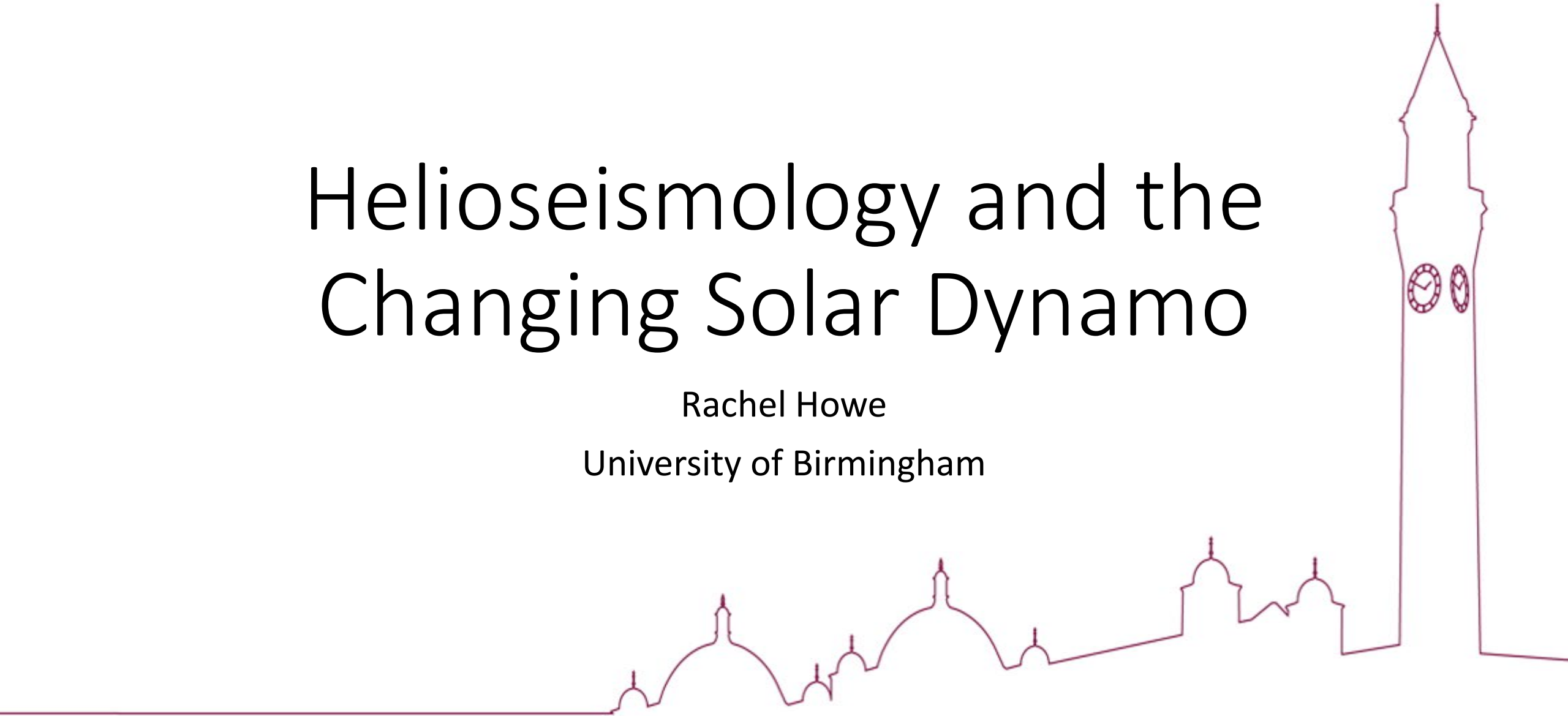


# Helioseismology and the Changing Solar Dynamo

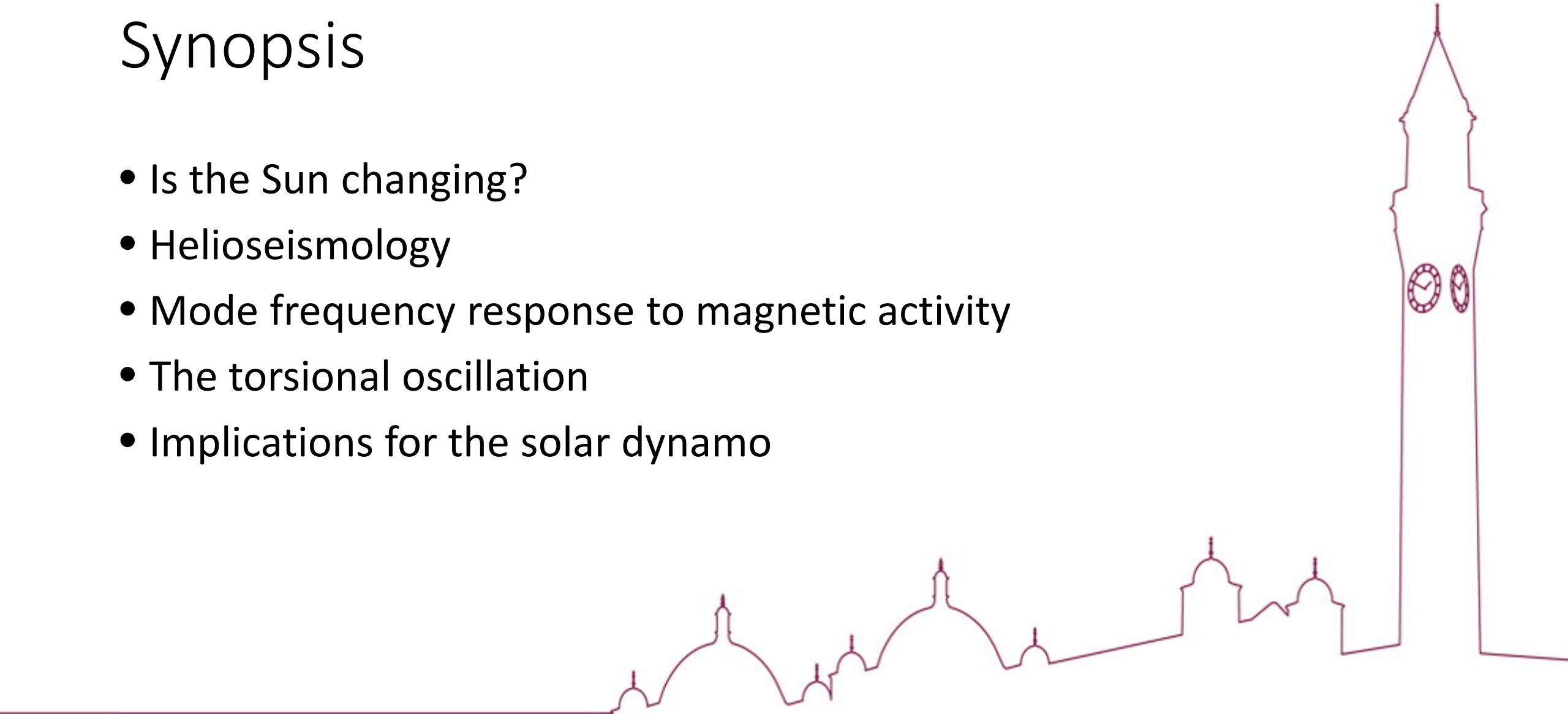
Rachel Howe

University of Birmingham



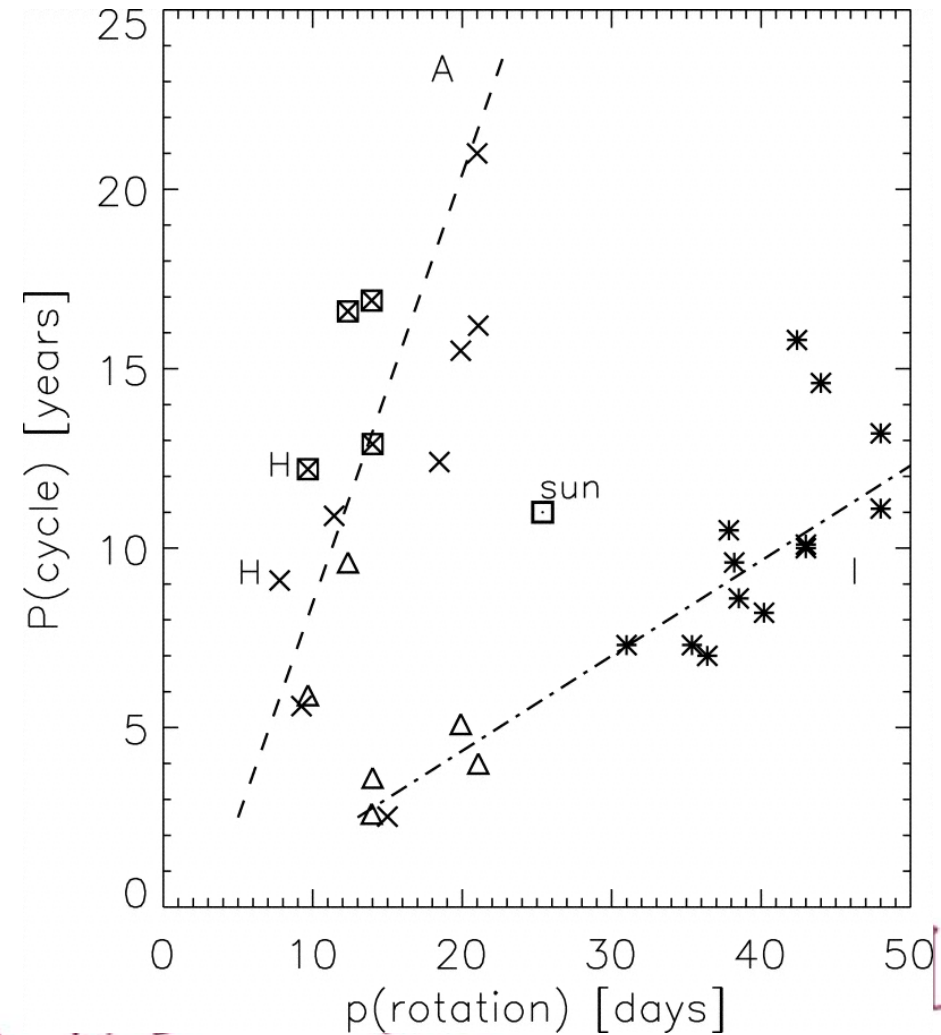
# Synopsis

- Is the Sun changing?
- Helioseismology
- Mode frequency response to magnetic activity
- The torsional oscillation
- Implications for the solar dynamo



# Is the Sun changing?

Bohm-Vitense (2007 *Astrophys. J.* 657, 486) – Stellar populations seem to lie along two branches on the rotation/cycle period plot, but the Sun lies somewhere in the middle.



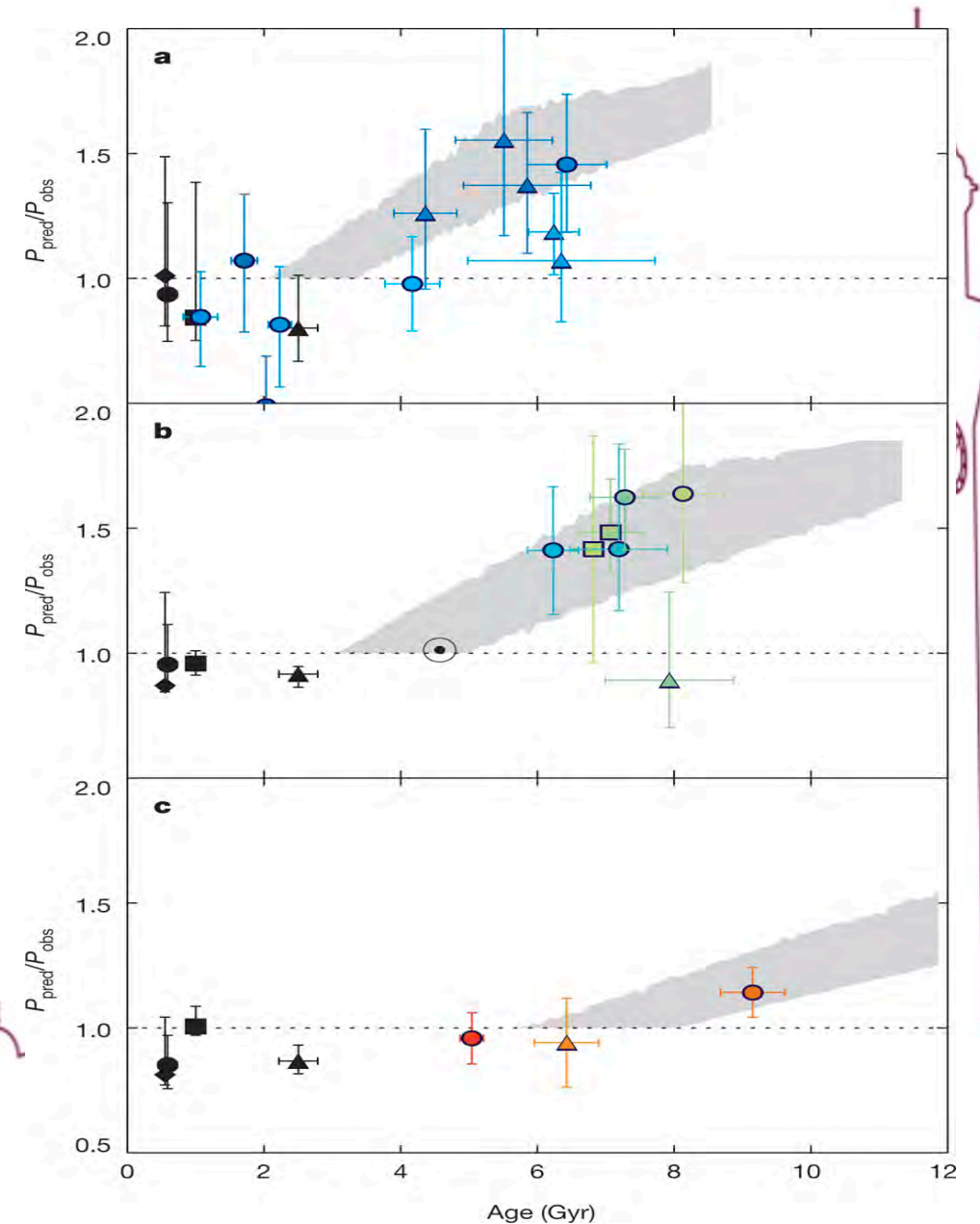
Ratios of the predicted  
rotation period to the  
observed period.

J L van Saders *et al. Nature* 1-4 (2016)  
doi:10.1038/nature16168

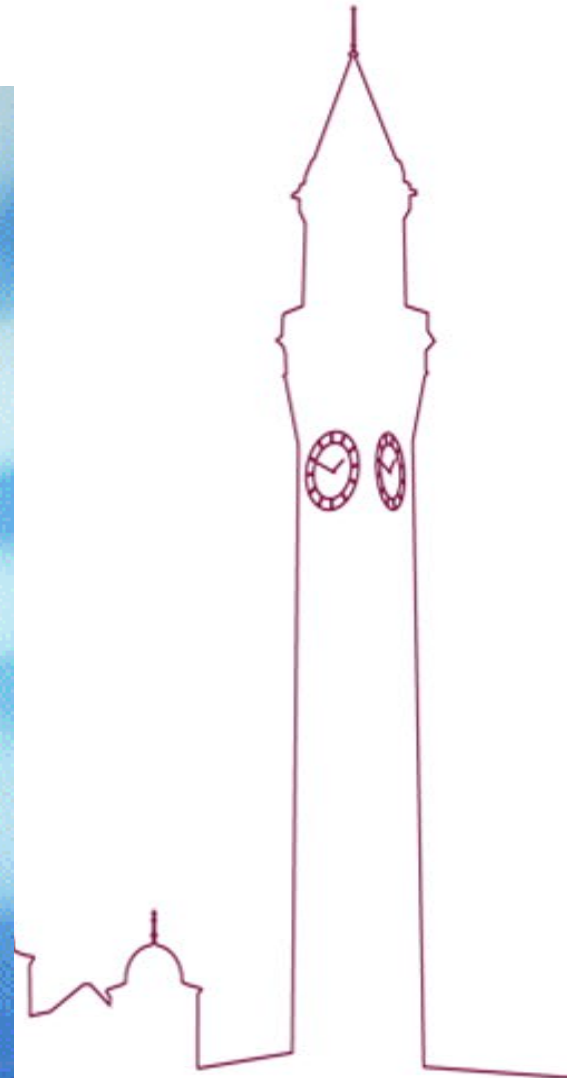
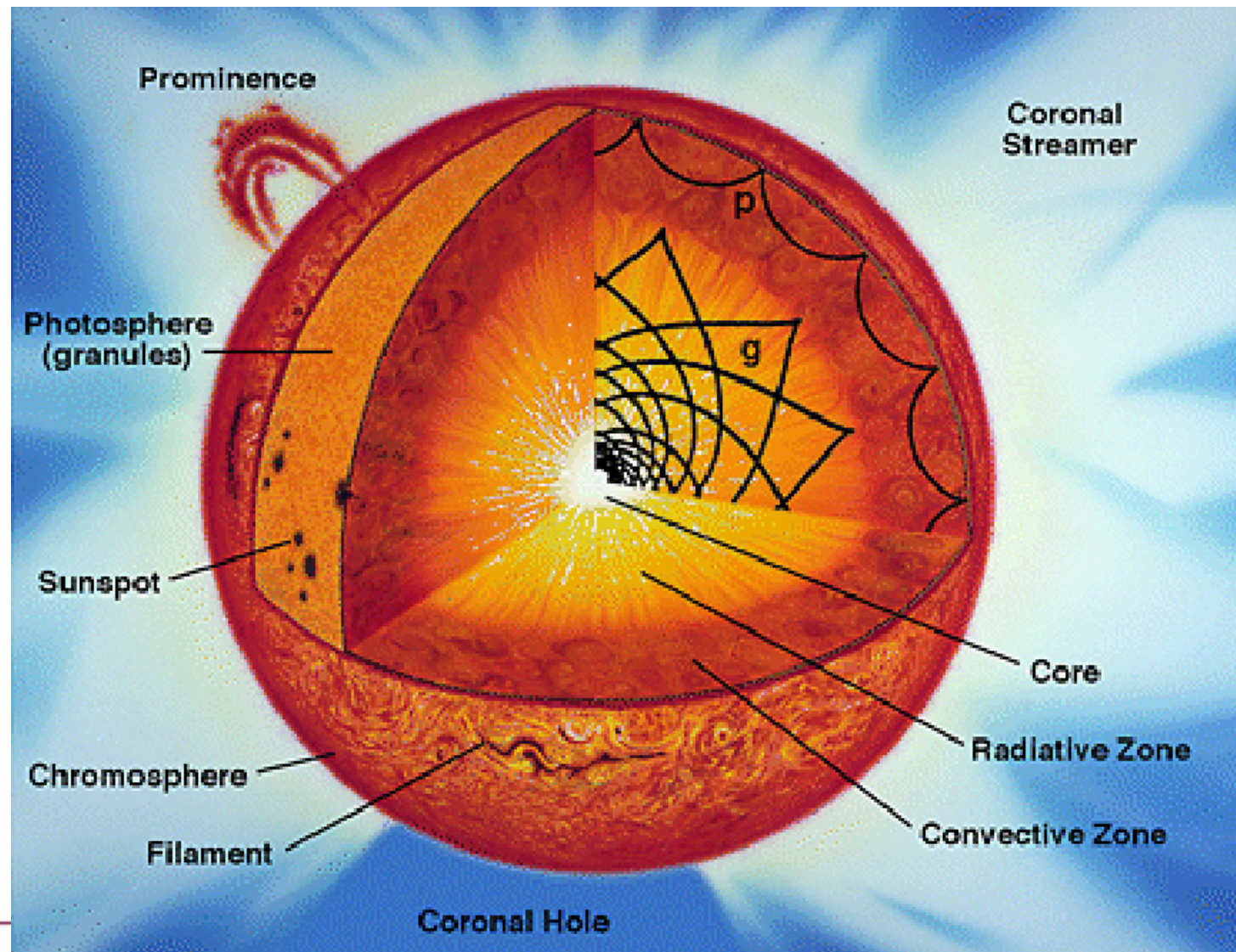
Older stars rotate faster than expected –  
reduced magnetic braking?

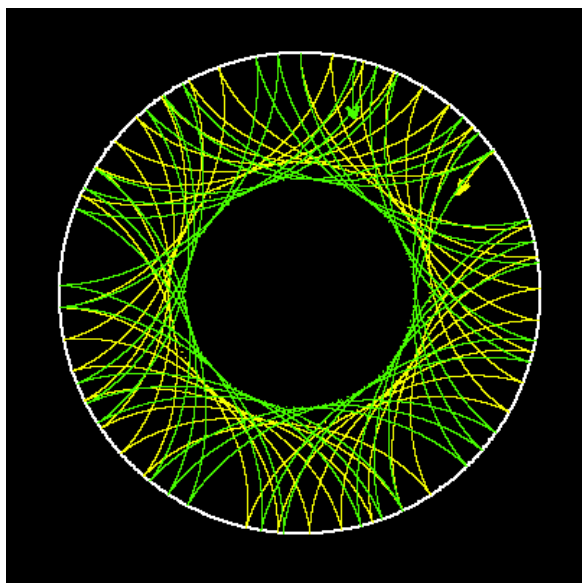
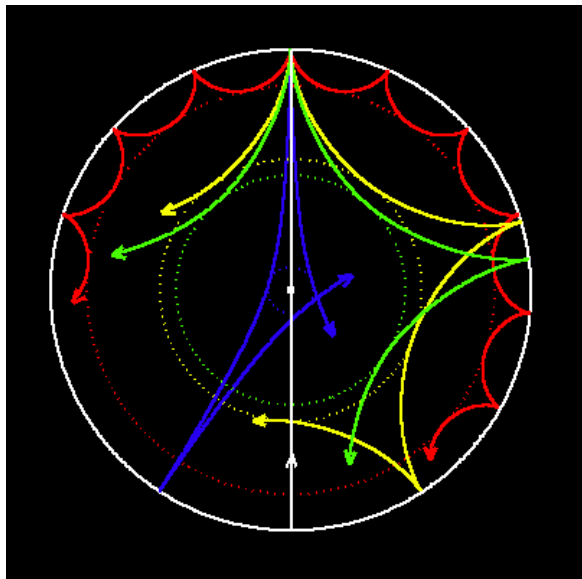
**“Our findings might suggest a  
fundamental change in the nature  
of ageing stellar dynamos, with the  
Sun being close to the critical  
transition to much weaker  
magnetized winds.”**

nature



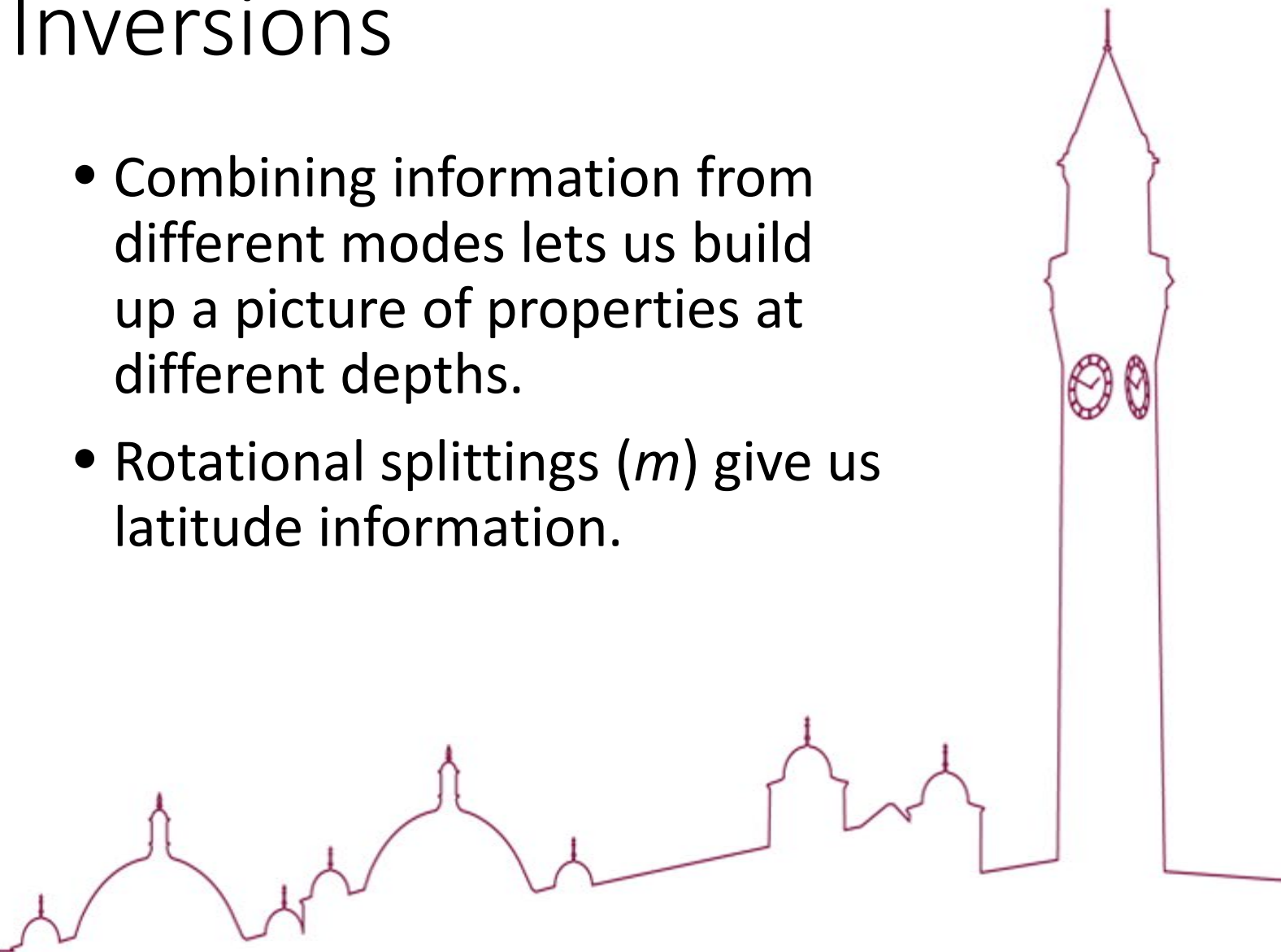
# Inside the Sun





# Inversions

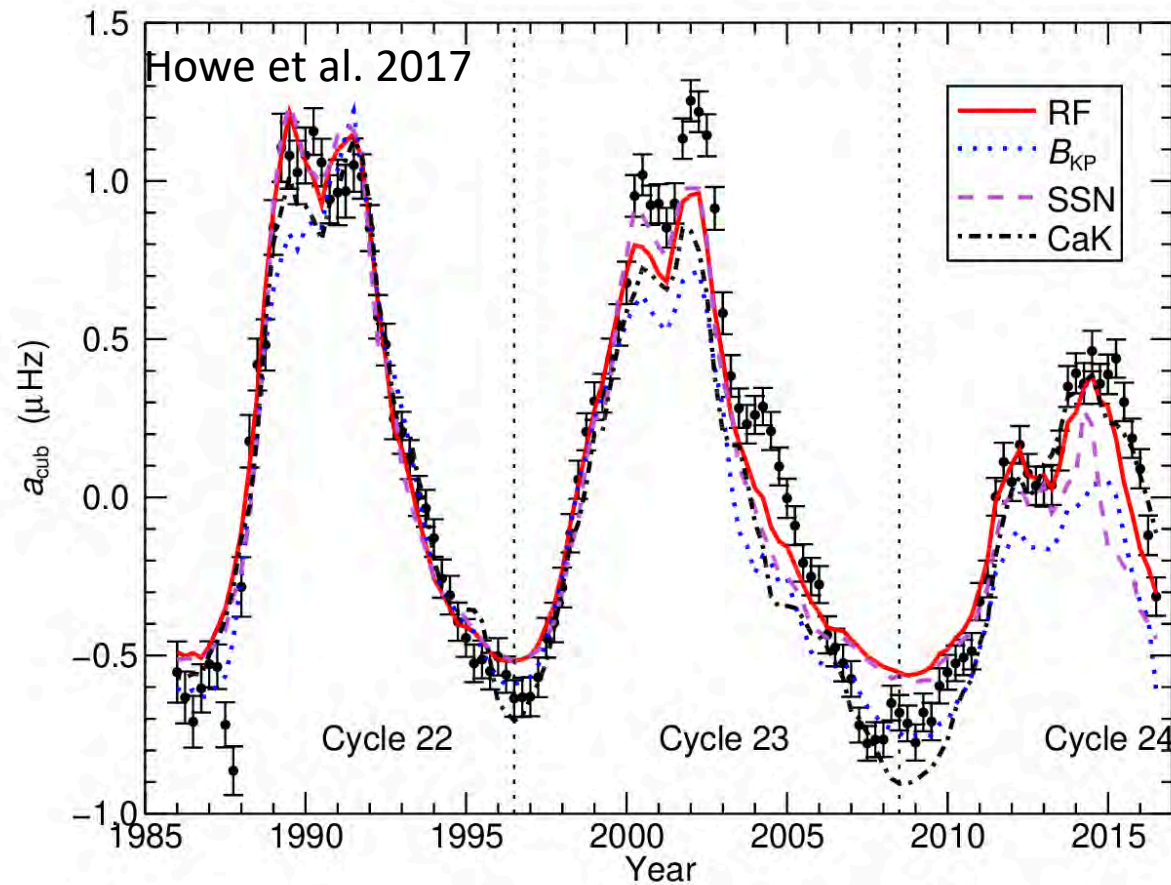
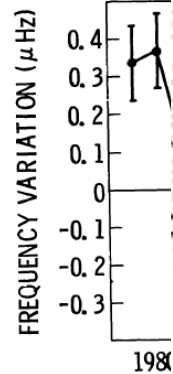
- Combining information from different modes lets us build up a picture of properties at different depths.
- Rotational splittings ( $m$ ) give us latitude information.



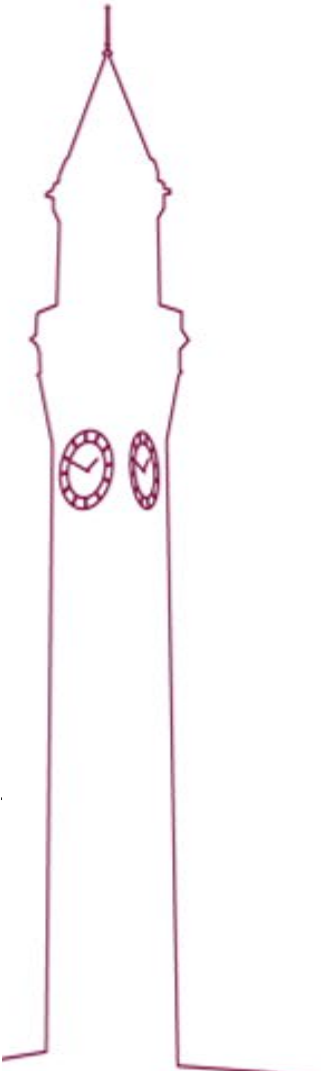
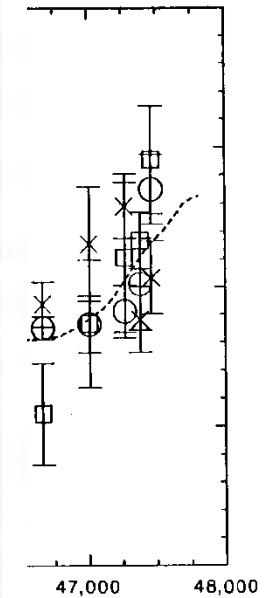


# Frequency shifts with solar cycle

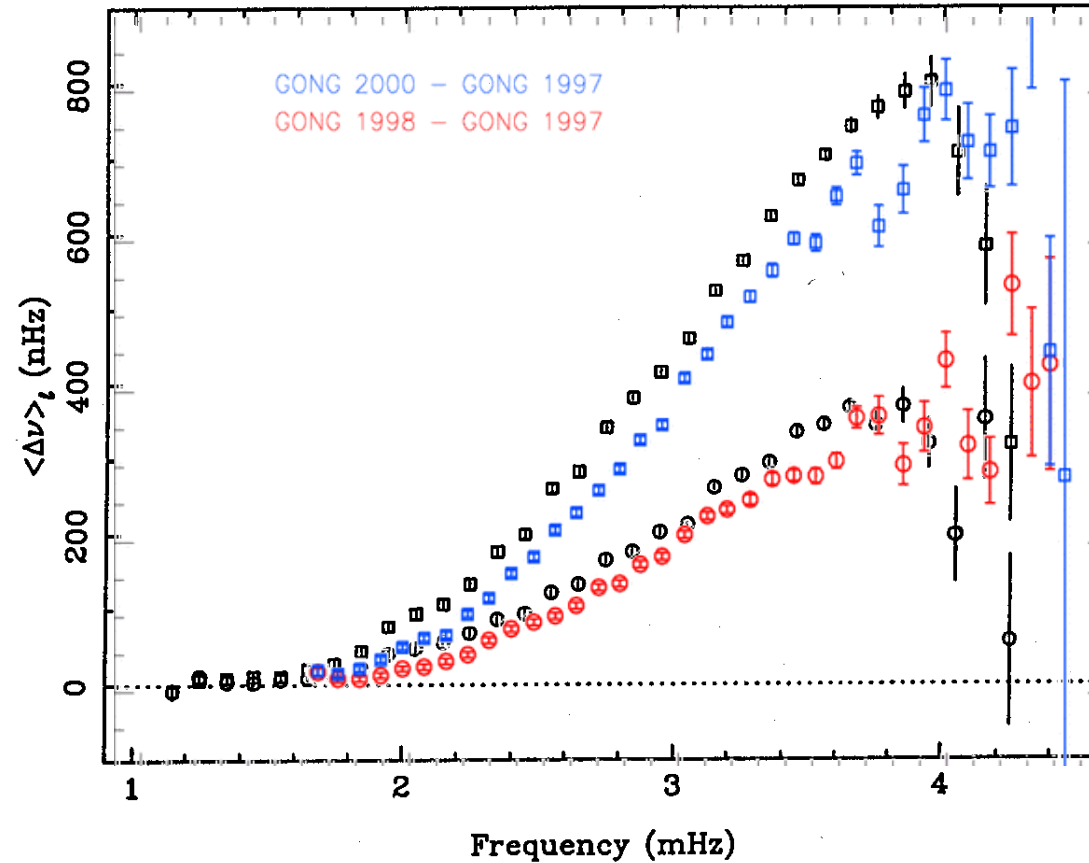
- AC  
19  
Gr



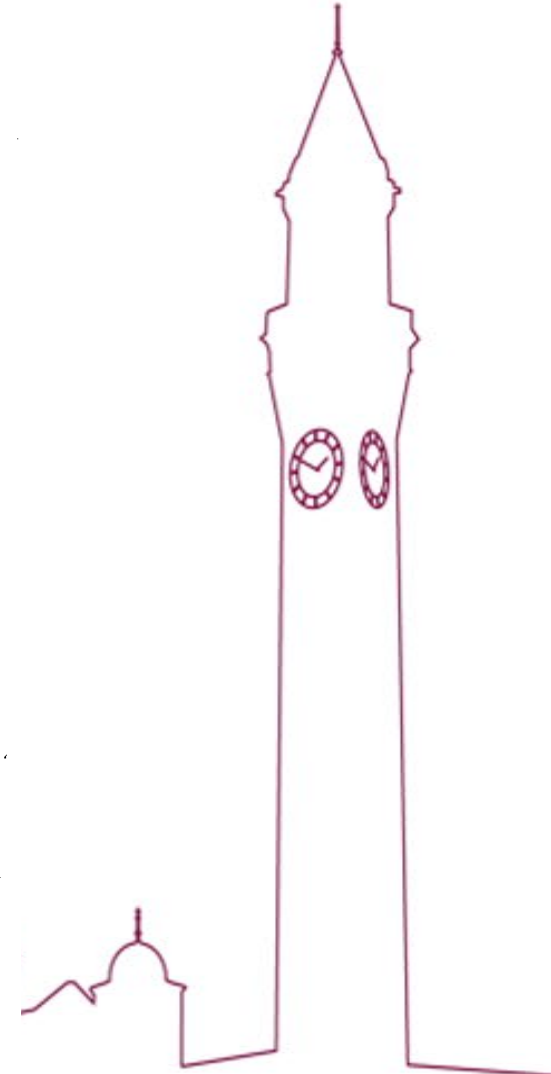
t al.  
(1990)



# Frequency shifts with solar cycle



**Figure 18.** Frequency shift as a function of frequency, using frequencies from 1986 as a reference. The frequency dependence was obtained by averaging over modes in the range  $4 \leq l \leq 140$  in degree. Data from 1988 are denoted by circles, data from 1989 by squares. From Woodard and Libbrecht (1991).





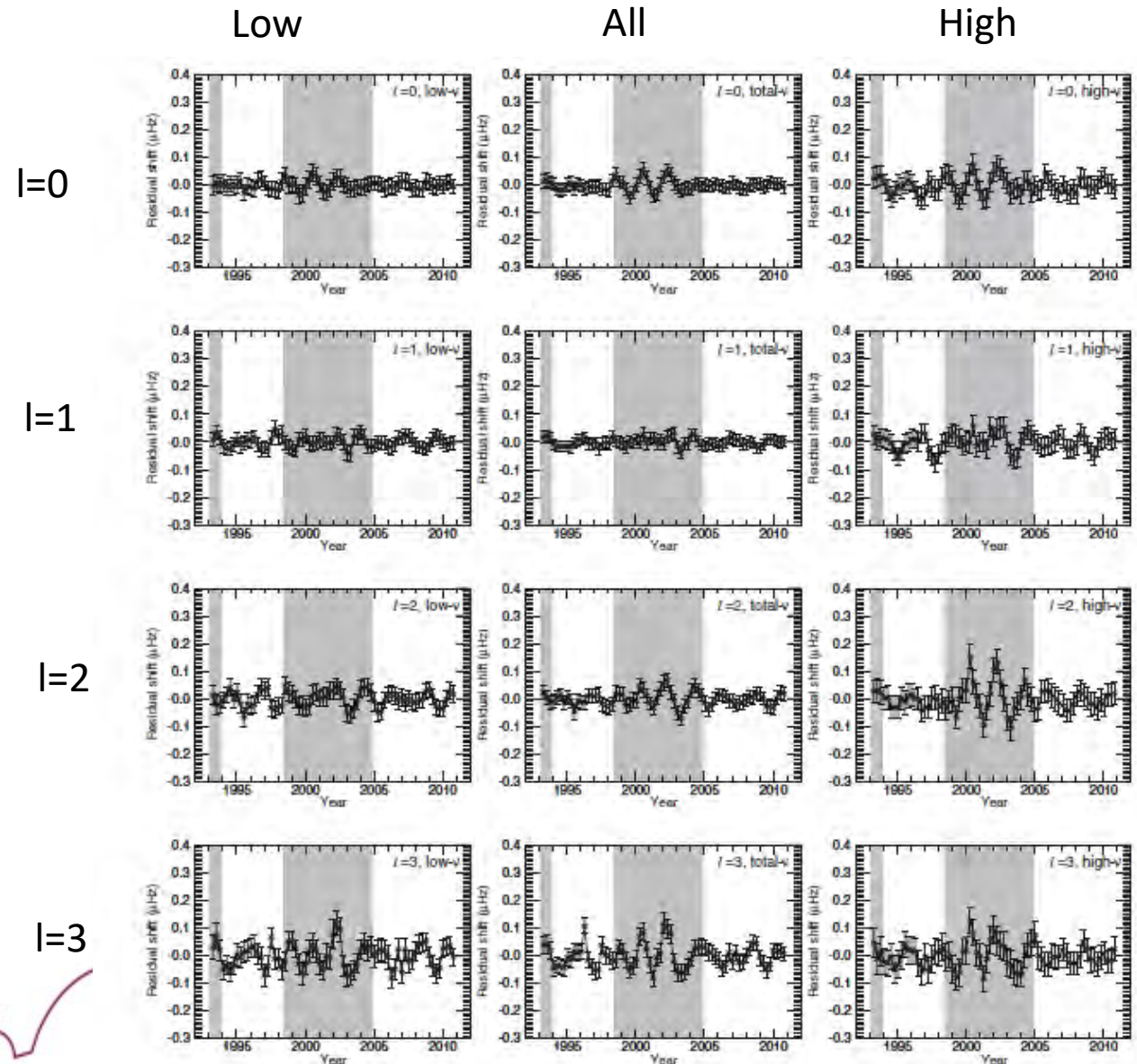
# Quasi-biennial oscillation

(Broomhall et al. 2014, etc.)

Short-term variations in activity index and corresponding frequency change after main 11-year cycle removed.

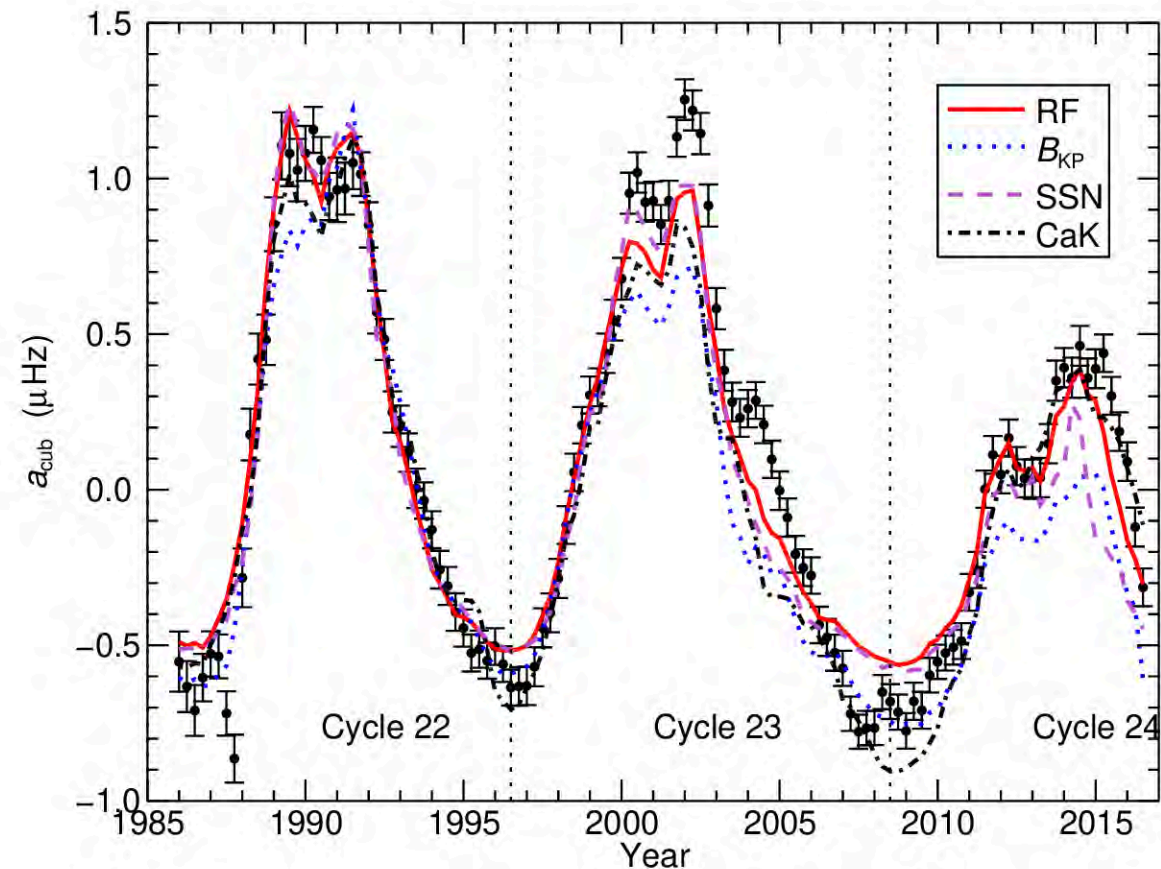
Modulated by the solar cycle

Frequency response varies less steeply with frequency than for main cycle – deeper fields?



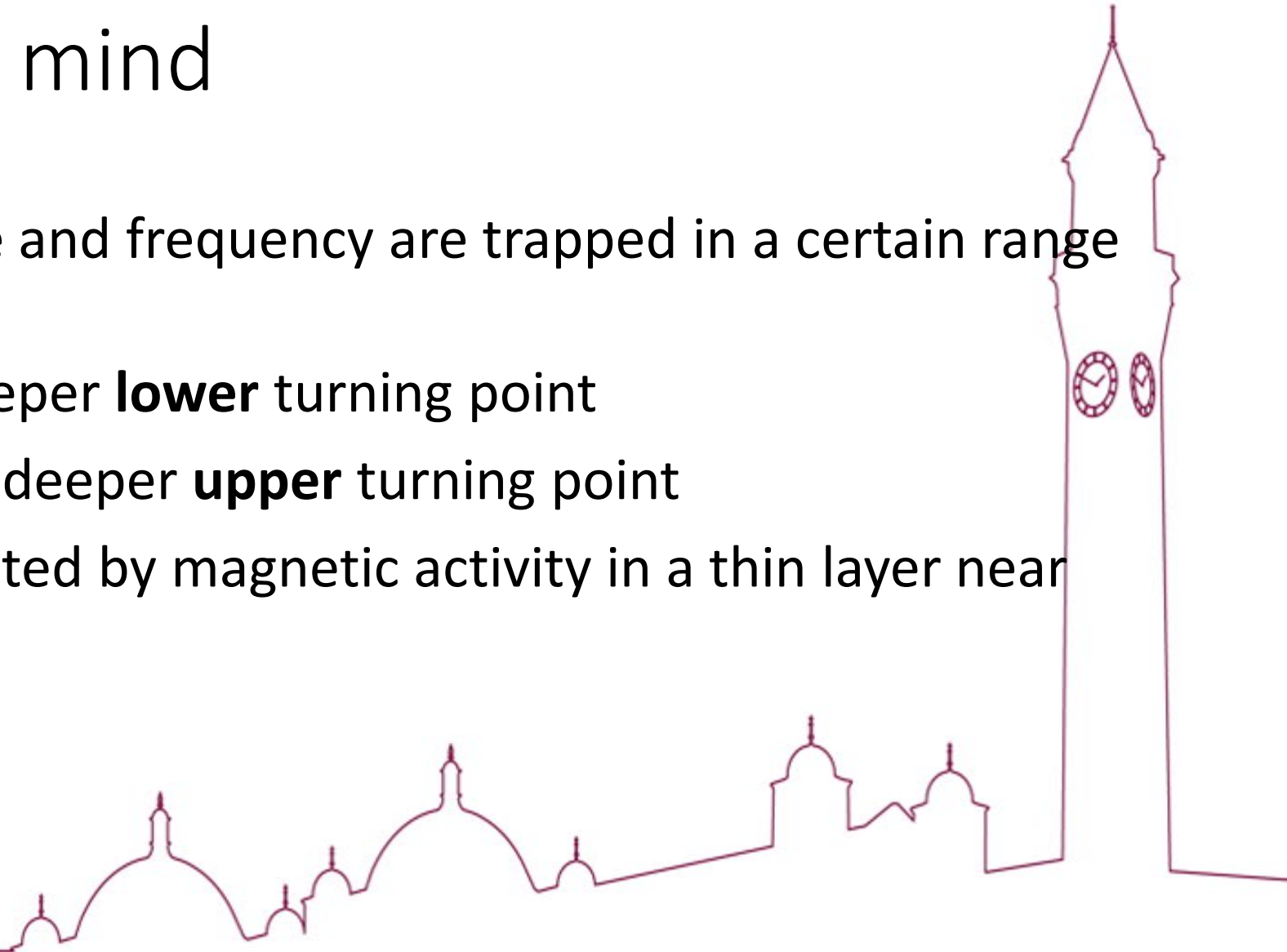
# BiSON frequency shifts over three cycles

Proxies scaled to Cycle 22 – not such a good fit for Cycles 23 and 24



# Things to bear in mind

- Modes of a given degree and frequency are trapped in a certain range of radii.
- Lower degree means deeper **lower** turning point
- Lower frequency means deeper **upper** turning point
- Frequency shifts dominated by magnetic activity in a thin layer near the surface.

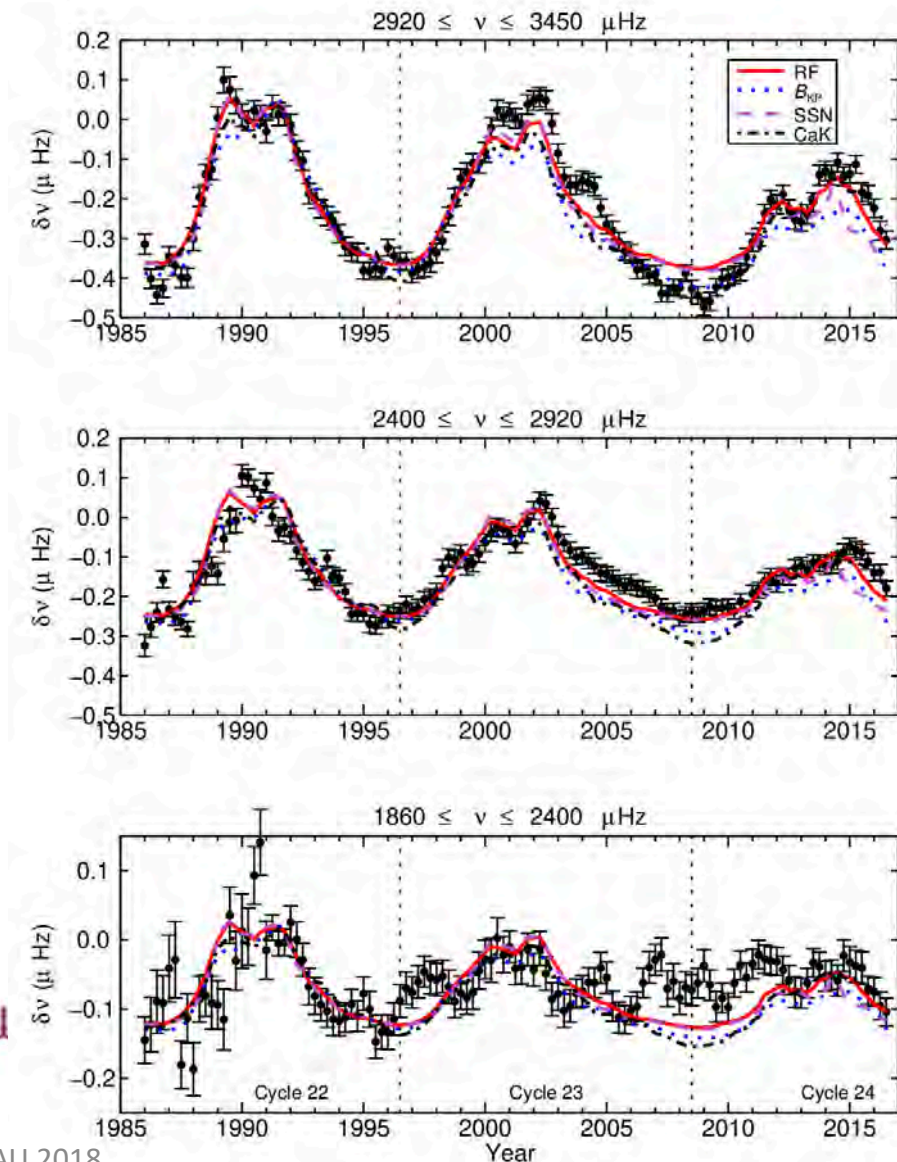


# BiSON frequency shifts over three cycles

In three frequency bands

Howe et al., 2018, *MNRAS* **480L**, 79

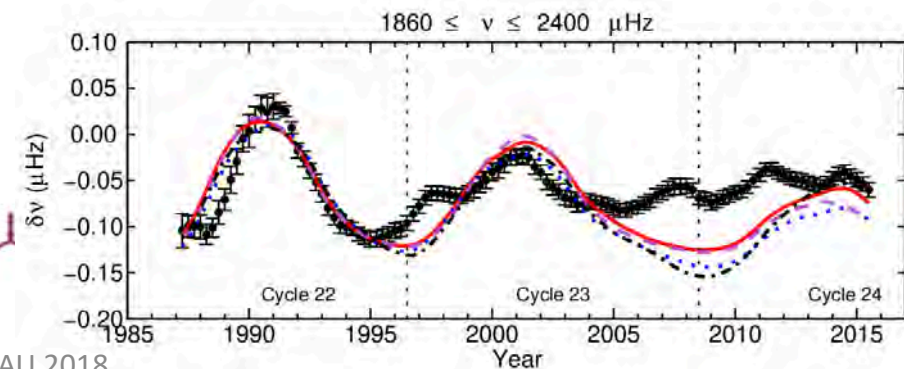
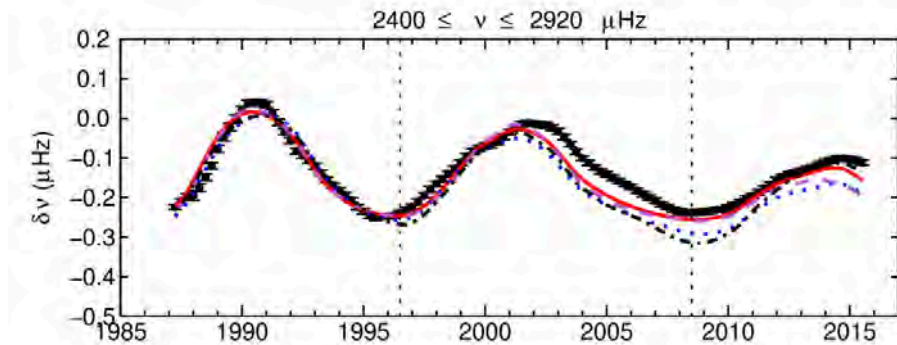
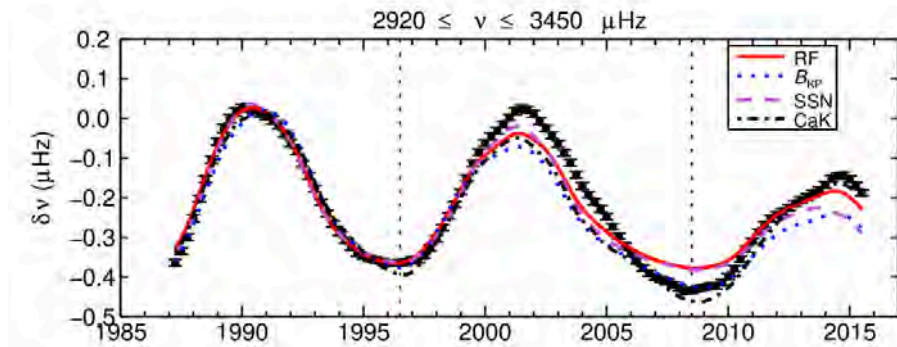
c.f. Basu et al. 2012, *ApJ* **758**, 43



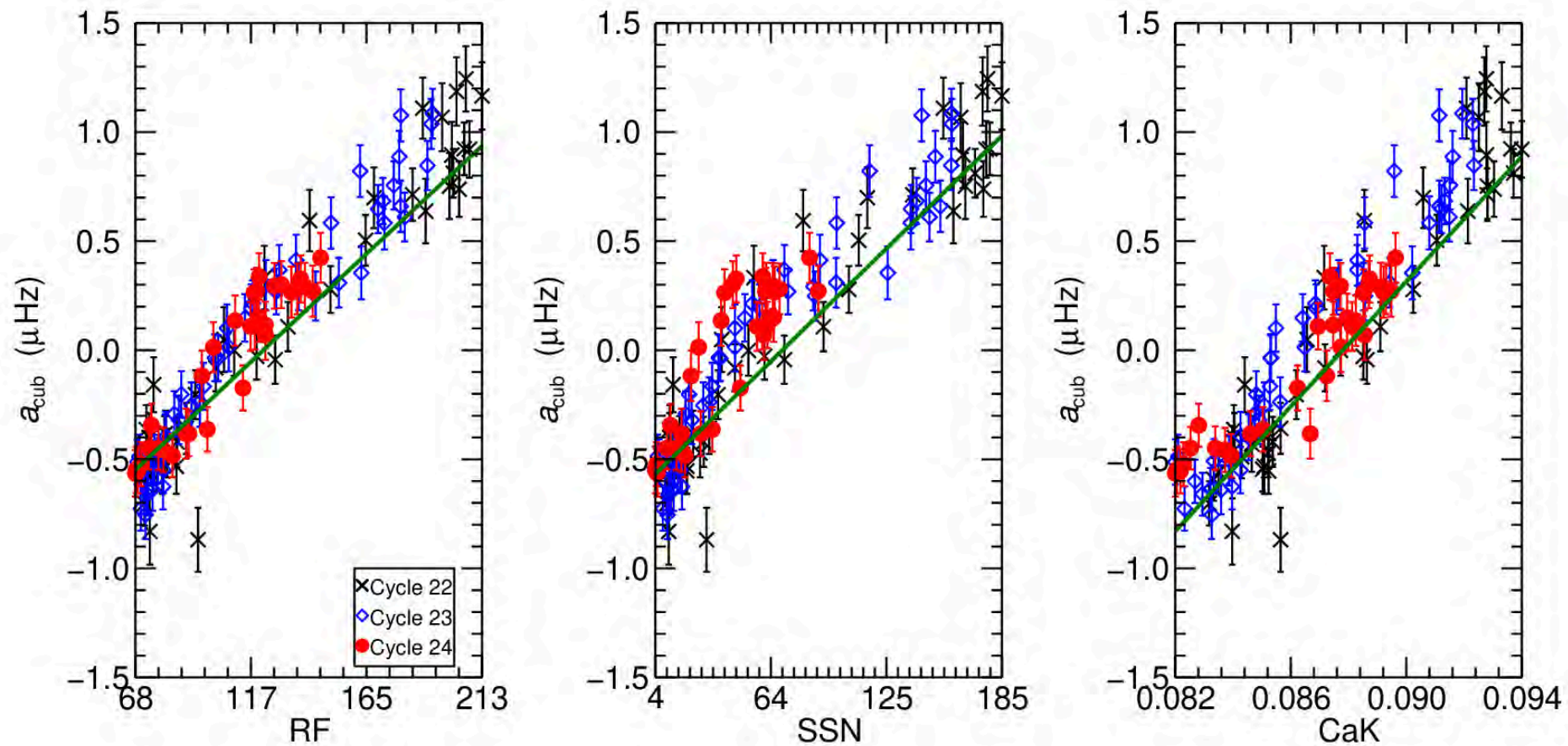


# BiSON frequency shifts over three cycles

In three frequency bands, smoothed



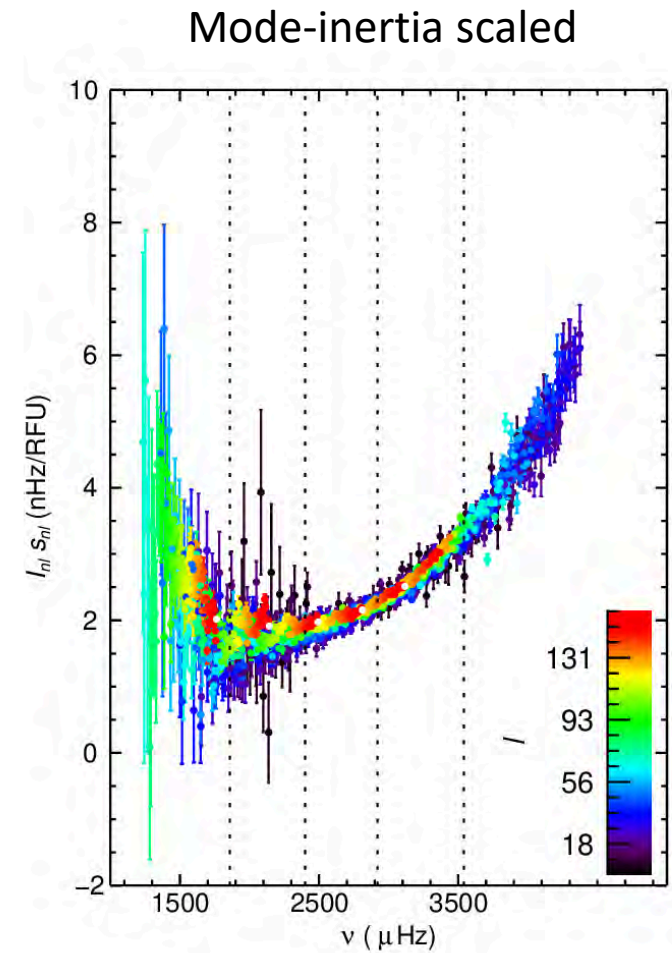
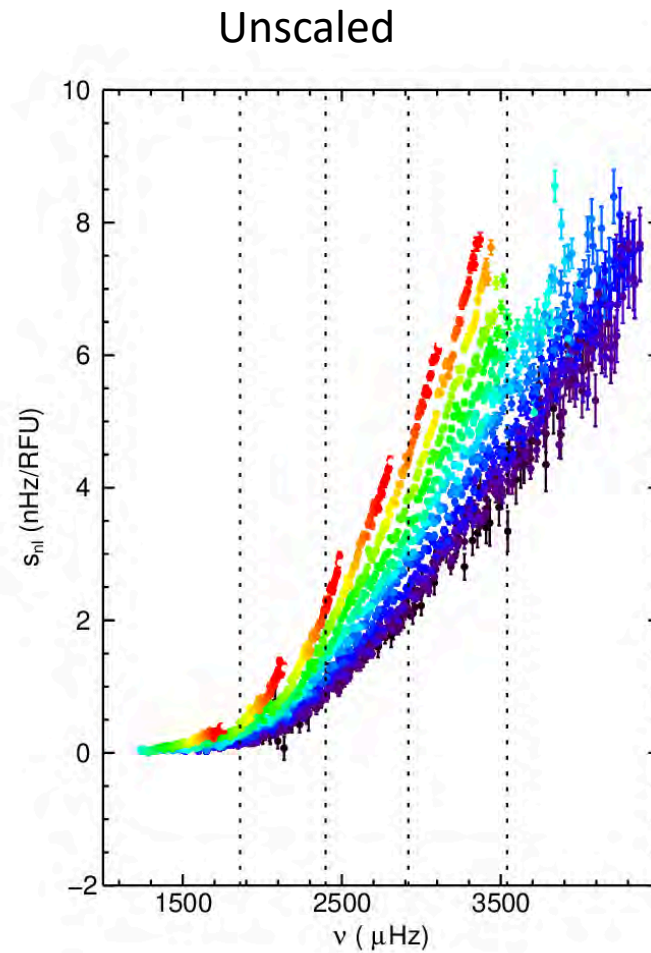
# Frequency shift vs activity proxy





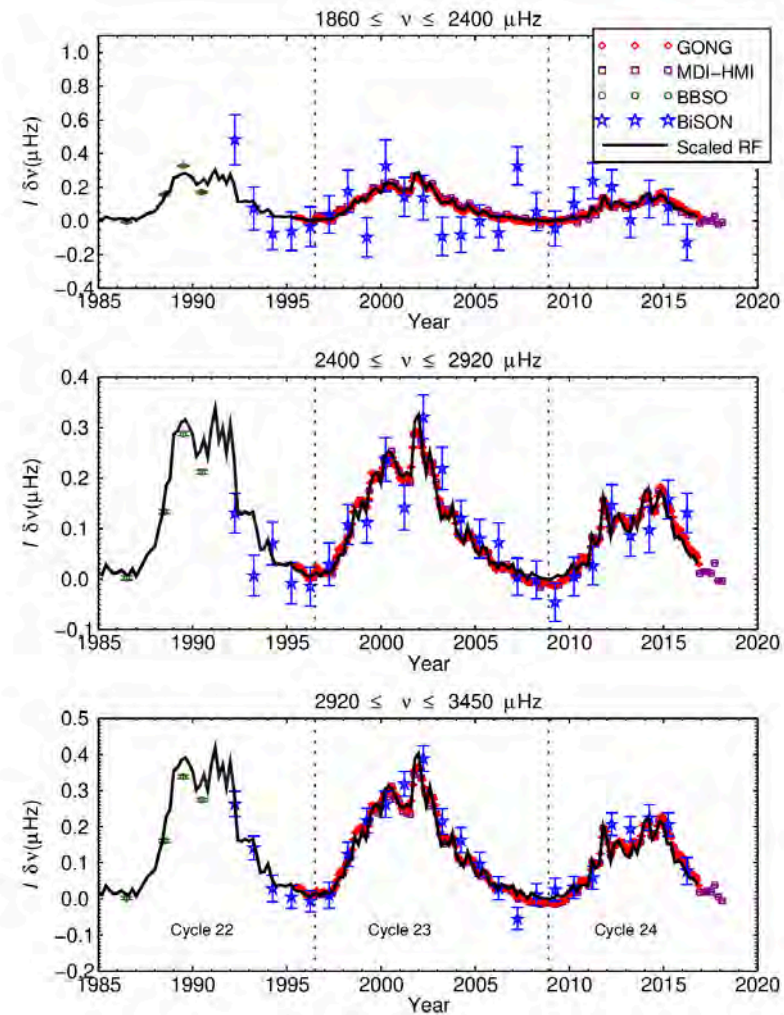
# Sensitivity vs frequency

Color-coded by degree  $l$

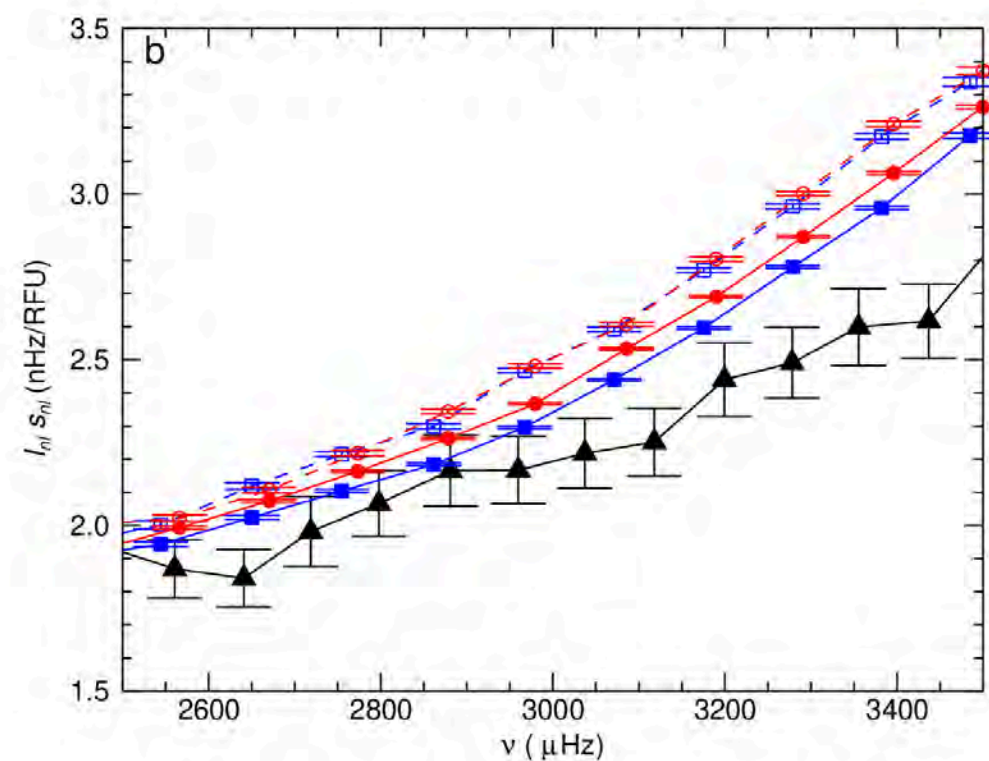
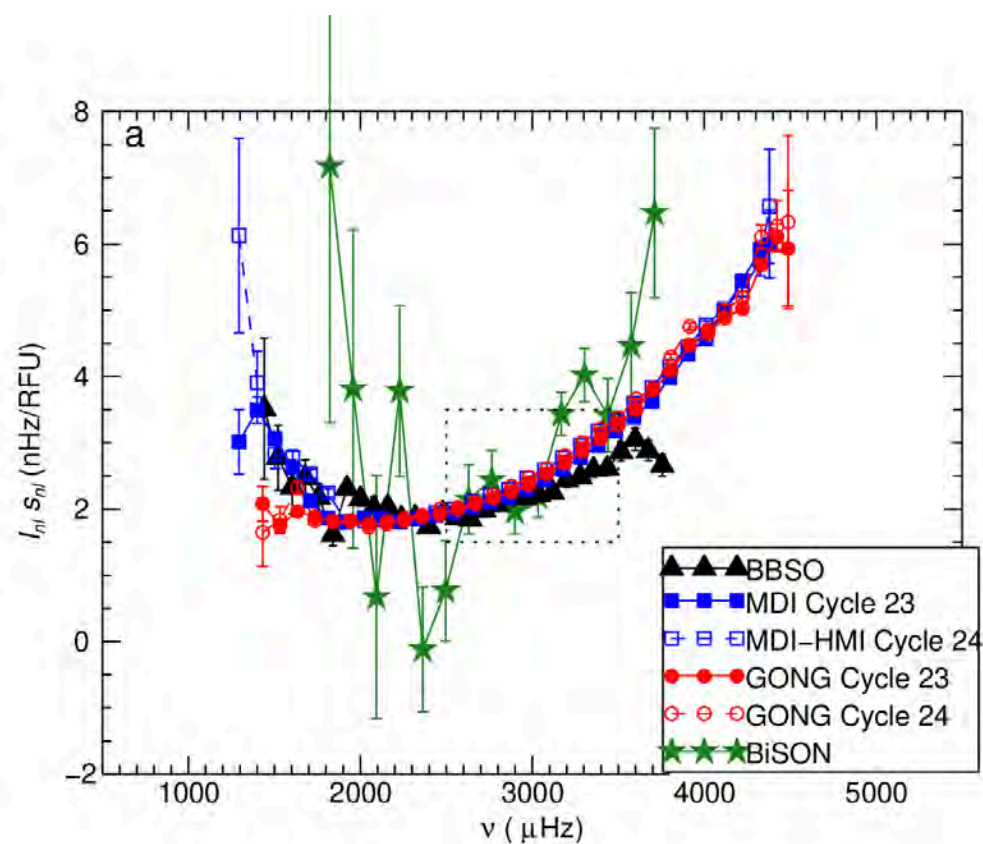


# Comparing medium- and low-degree frequency shifts

GONG, MDI/HMI, BBSO, and BiSON  $l=0$



# Frequency-binned, scaled shifts



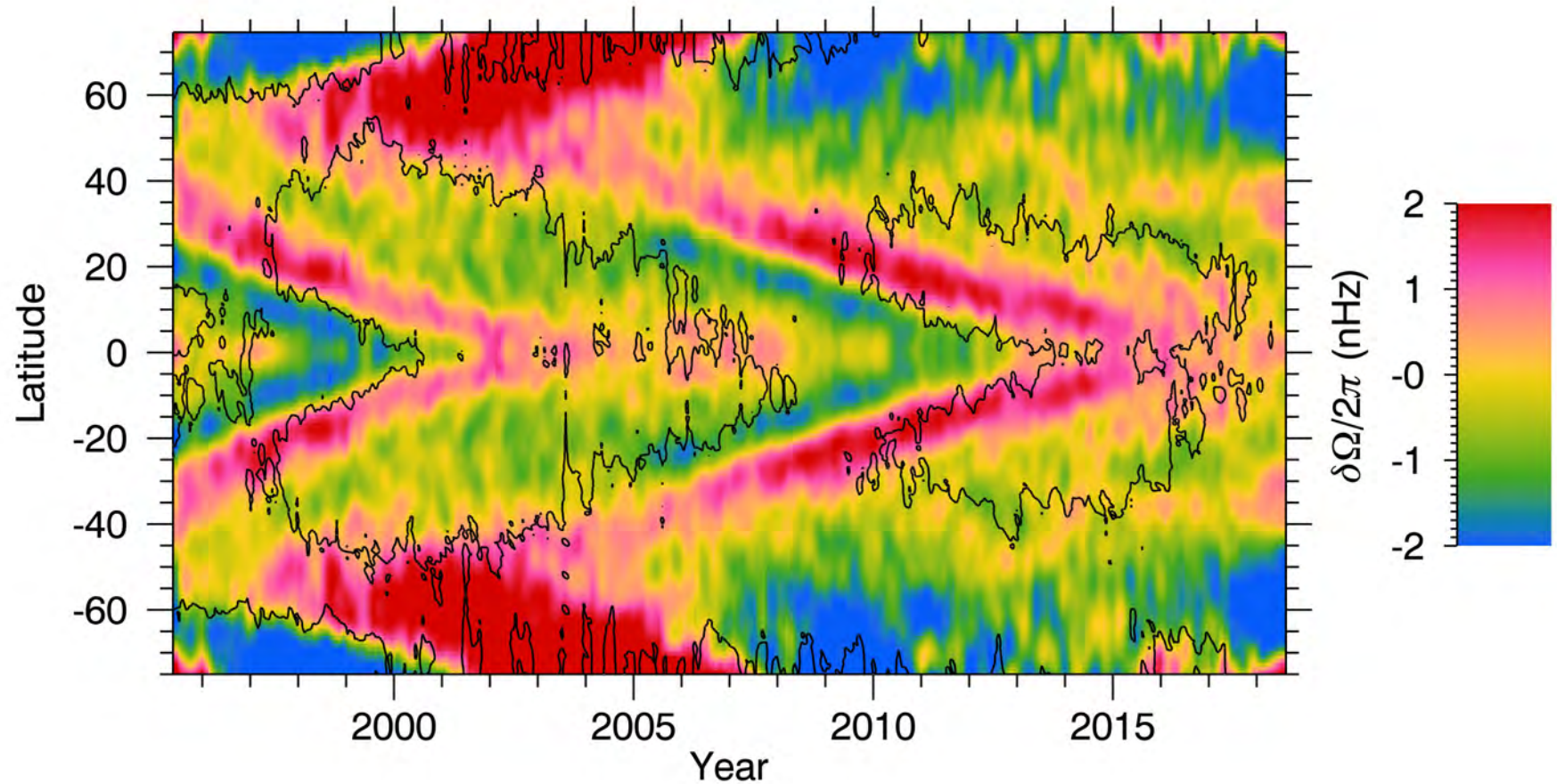
# Frequency Shifts - Summary

- Frequencies follow the activity level
- Low-degree, low frequency modes barely correlated with activity in the most recent two cycles – thinning magnetic layer?
- Medium-degree data does show correlation even at low frequencies.
- Both medium and low-degree data show slightly higher sensitivity to activity in Cycle 24 than Cycle 23 --- again, this could point to a thinning magnetic layer.



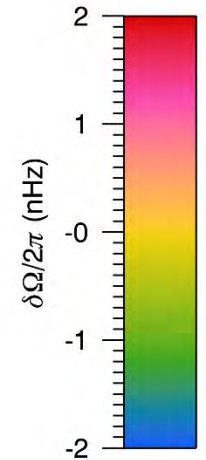
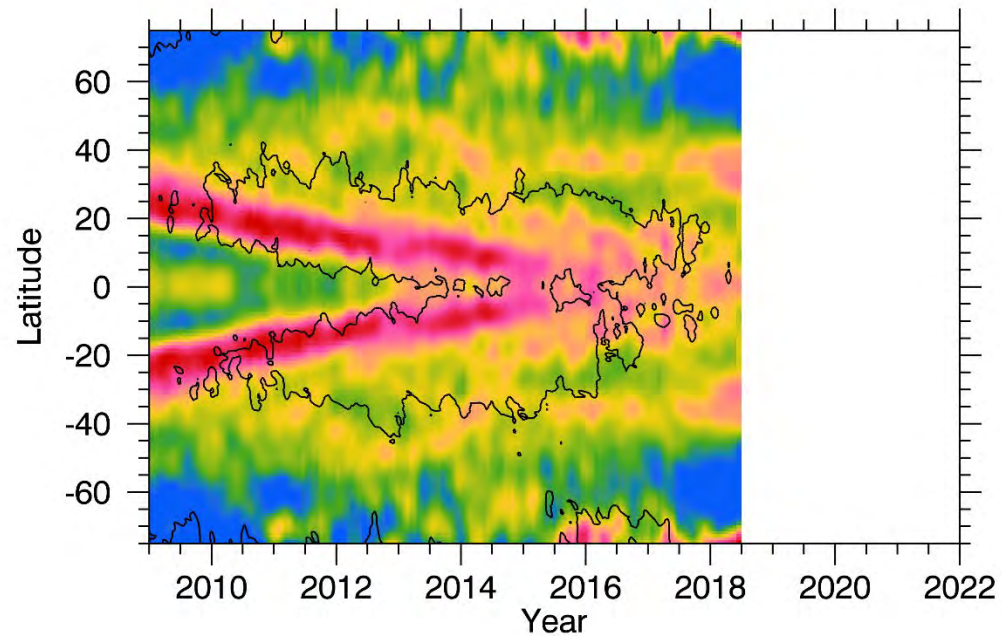
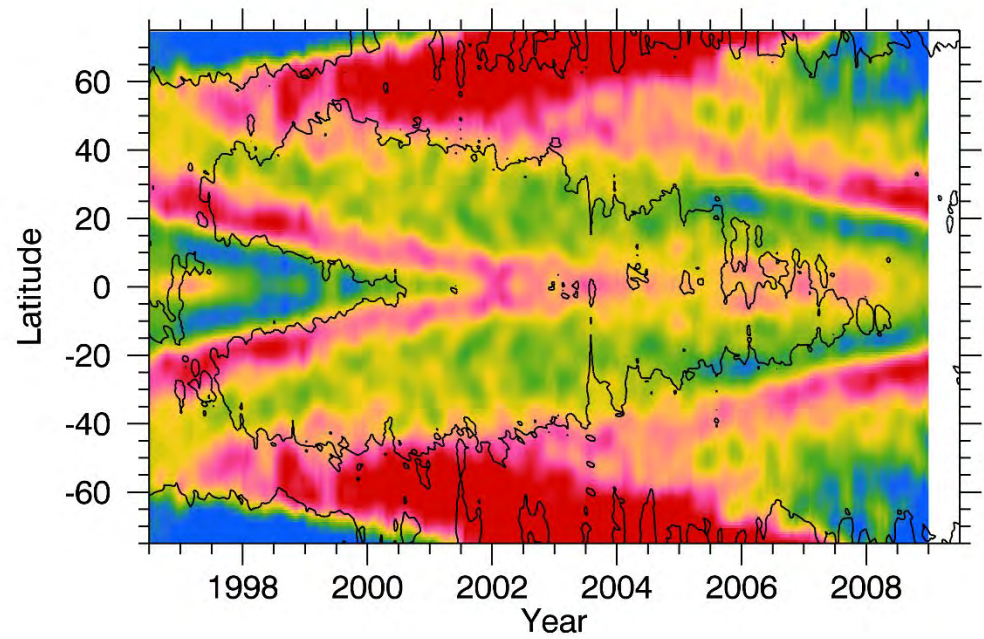
# The torsion oscillation

- 23 years of observations from GONG, MDI, HMI.
- Mean subtracted over whole time series of 2d rotation inversions at each location.



# Comparing Cycles 23 and 24

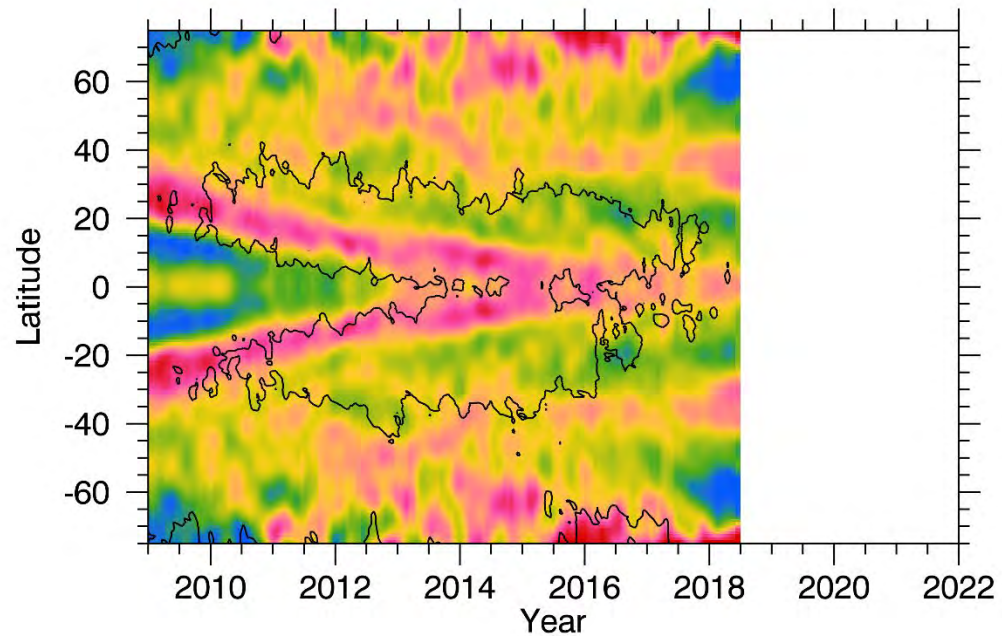
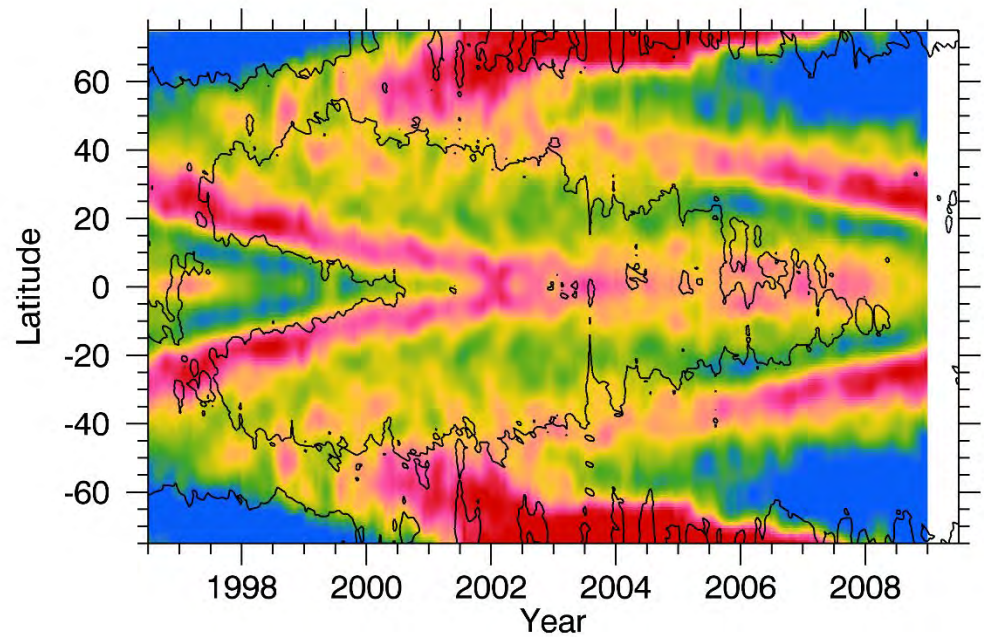
Mean subtracted over whole time series





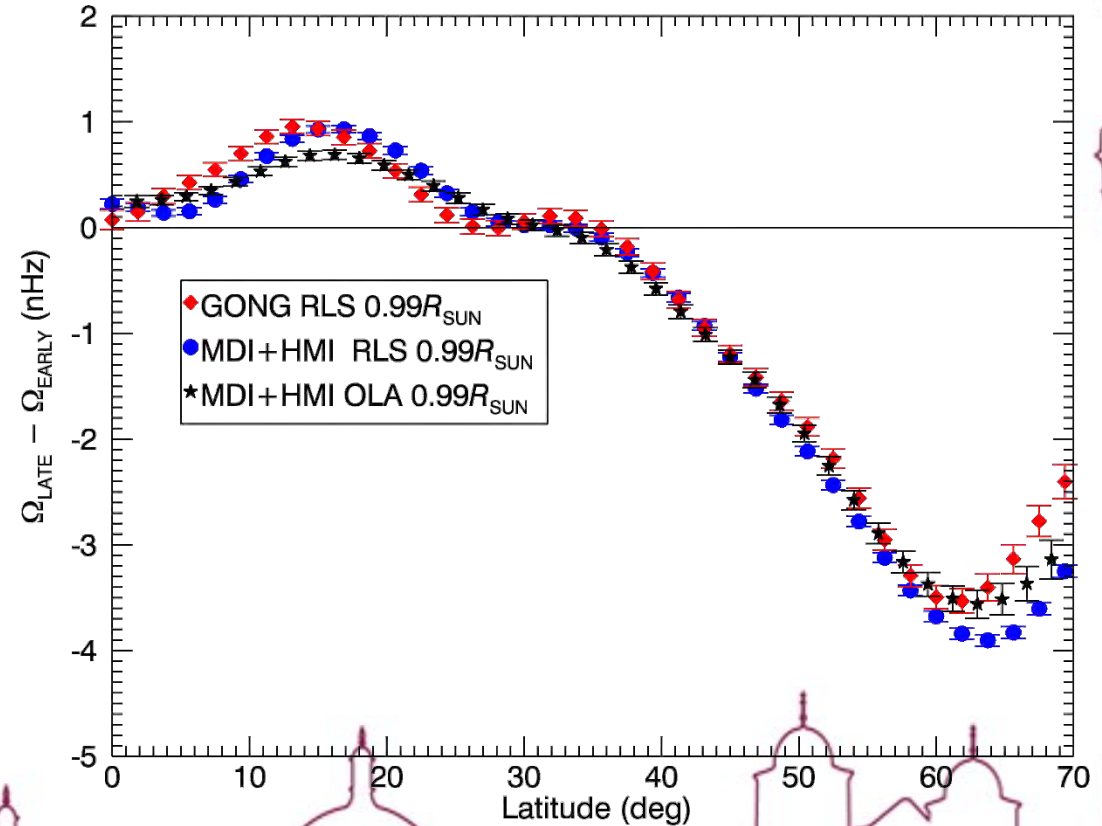
# Comparing Cycles 23 and 24

Mean subtracted over 8.5 years starting a year before solar minimum



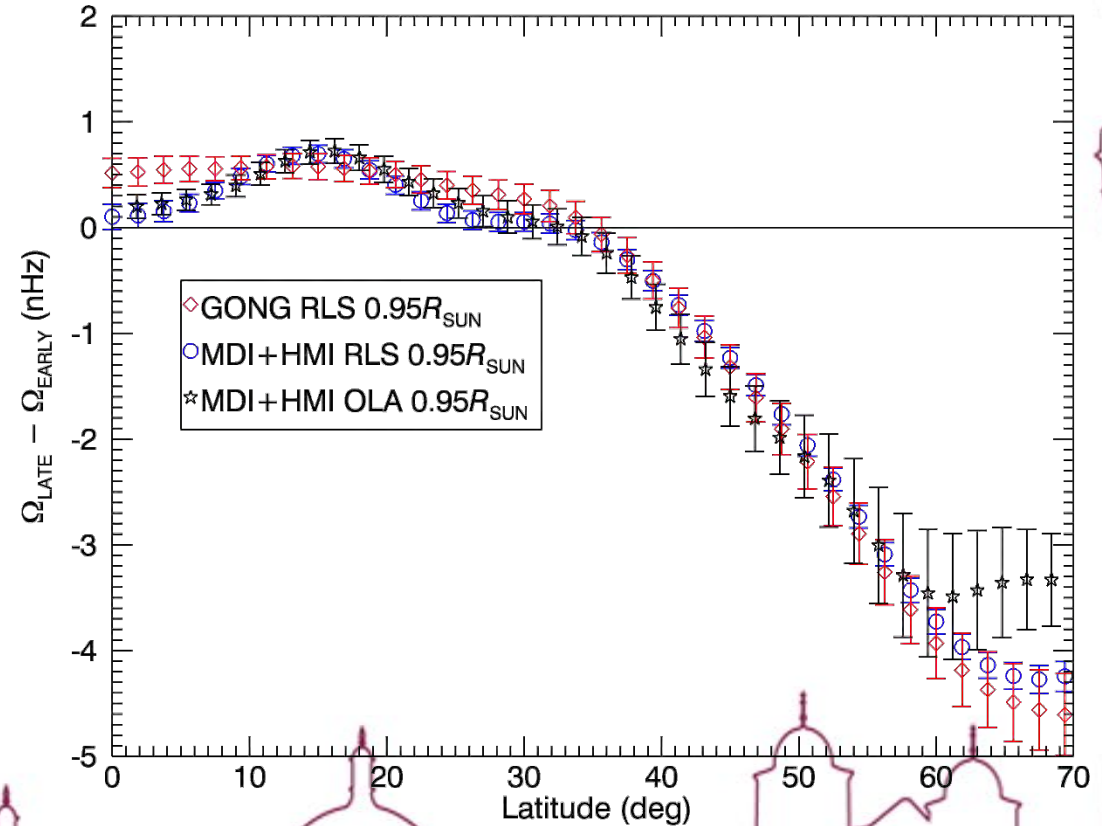
# Difference between Cycle 24 and Cycle 23 mean rotation rate

Mean taken over 8.5 years starting  
one year before minimum

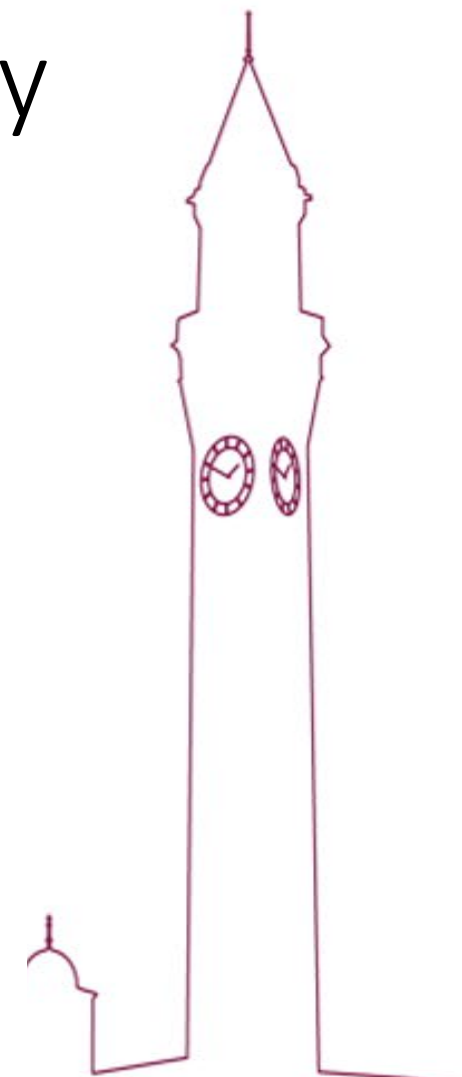
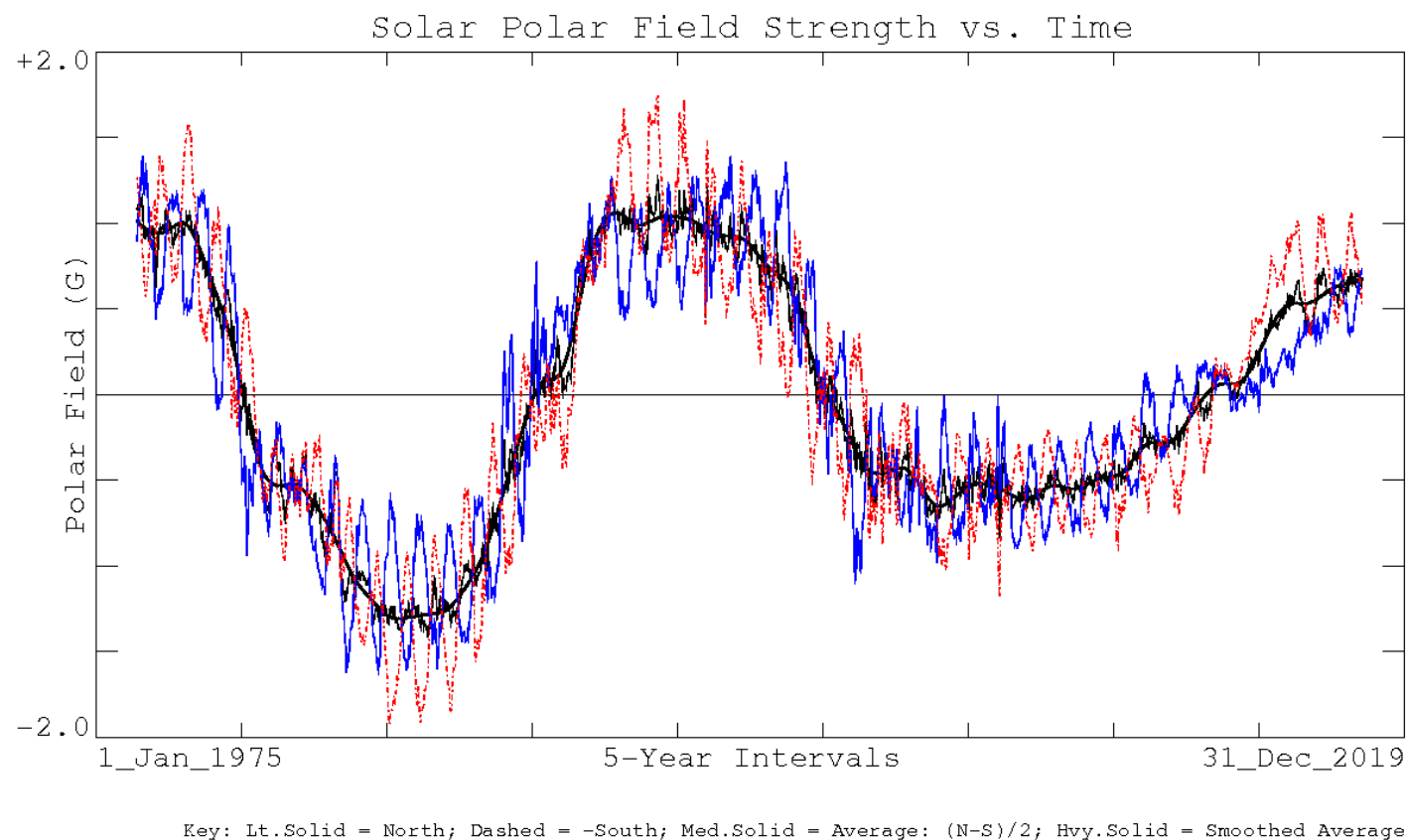


# Difference between Cycle 24 and Cycle 23 mean rotation rate

Mean taken over 8.5 years starting  
one year before minimum

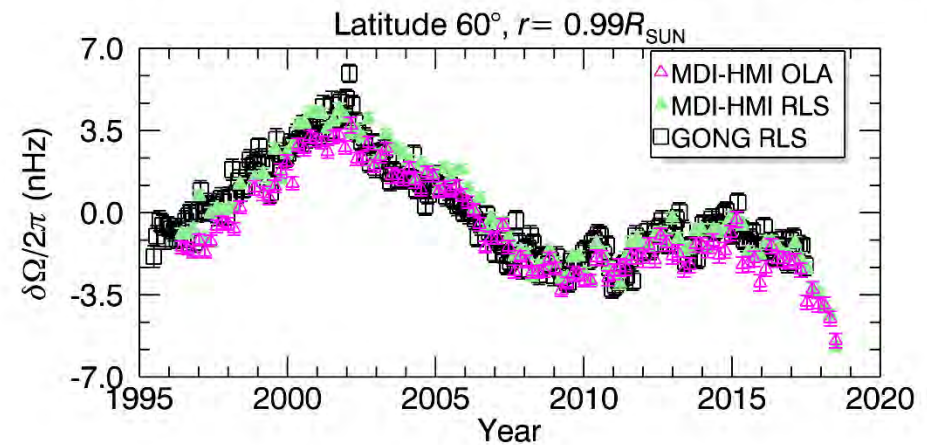
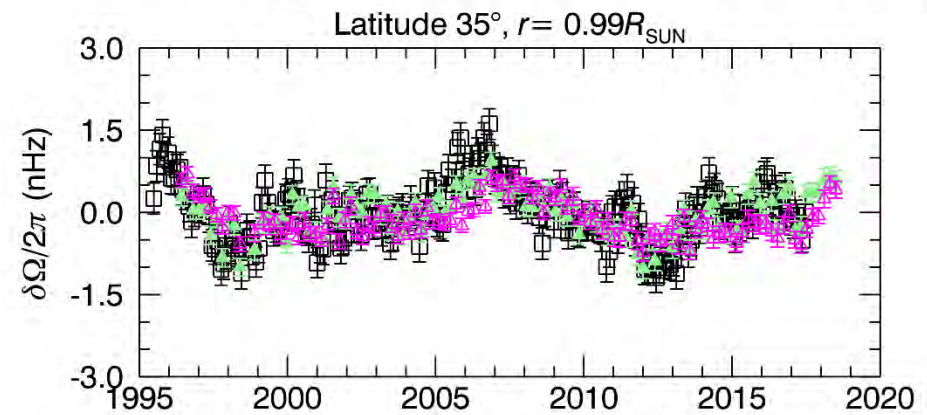
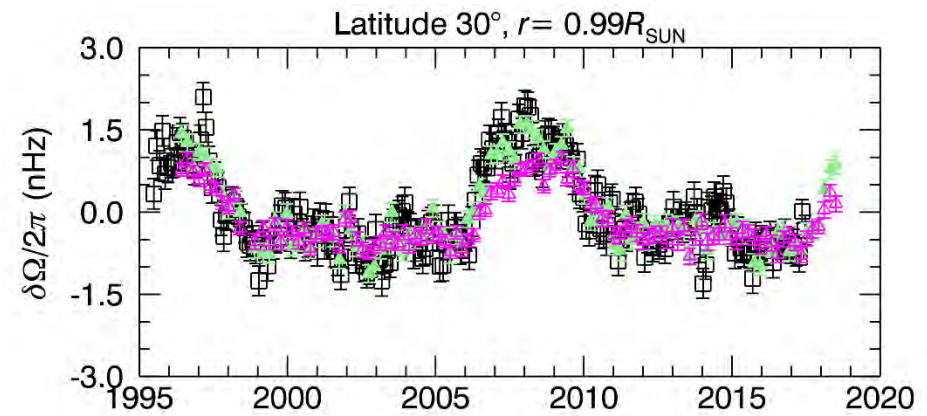


# Polar Field from Wilcox Solar Observatory



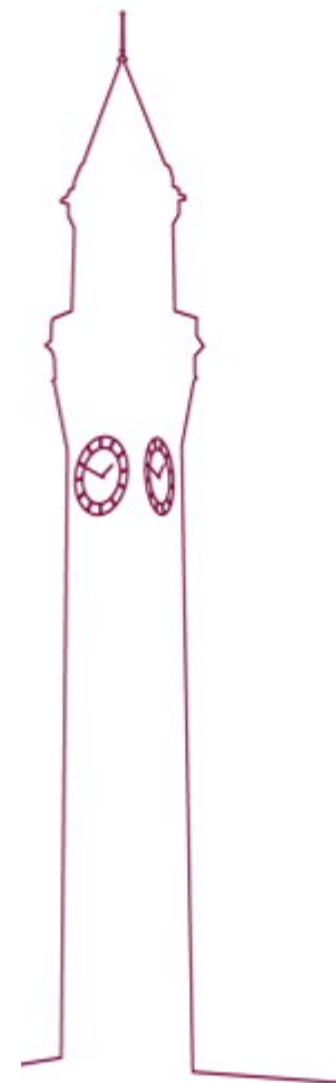
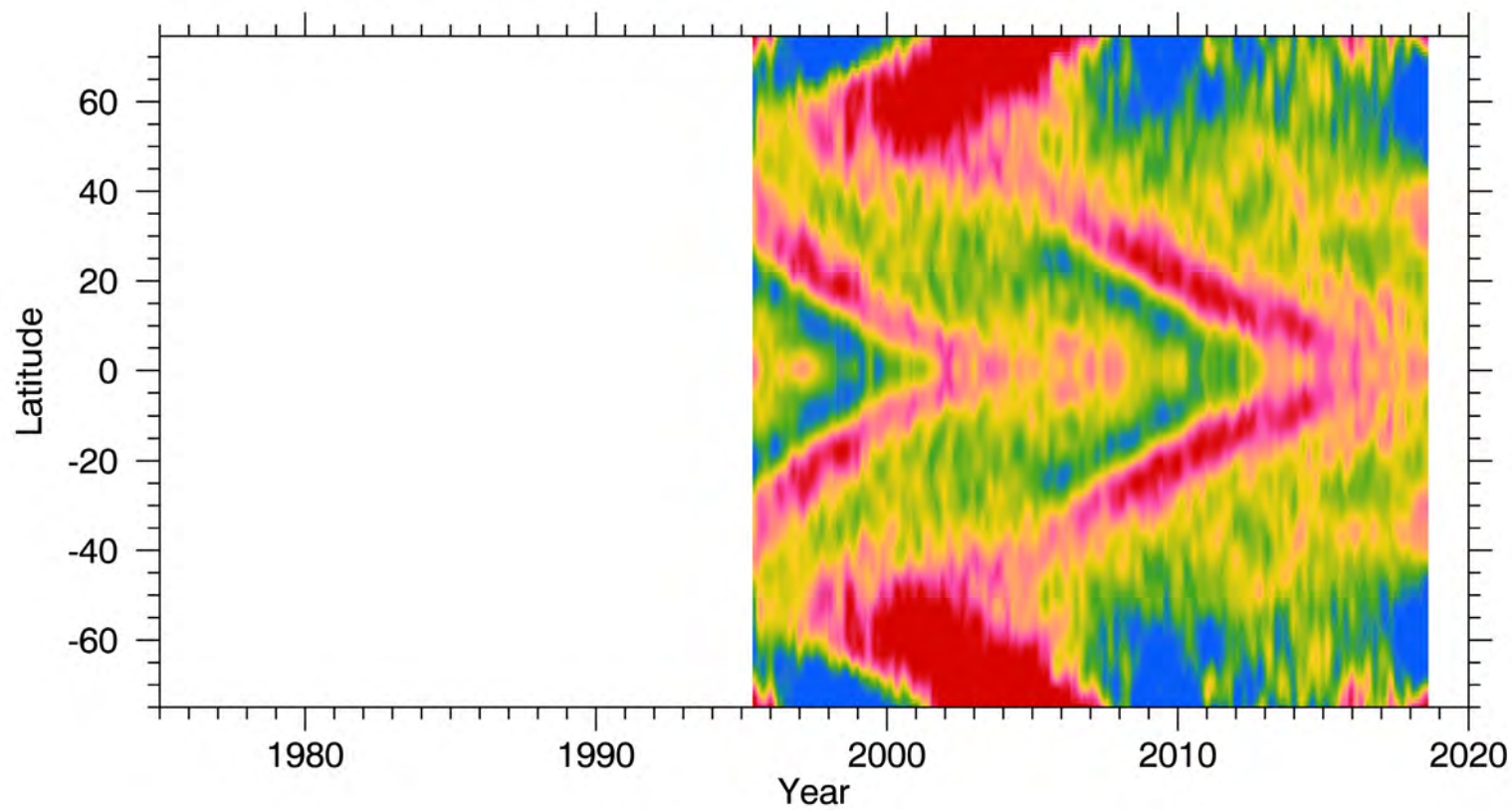
## Near-surface residuals at selected latitudes

- Acceleration at 30-35 degrees
- Deceleration at 60 degrees.



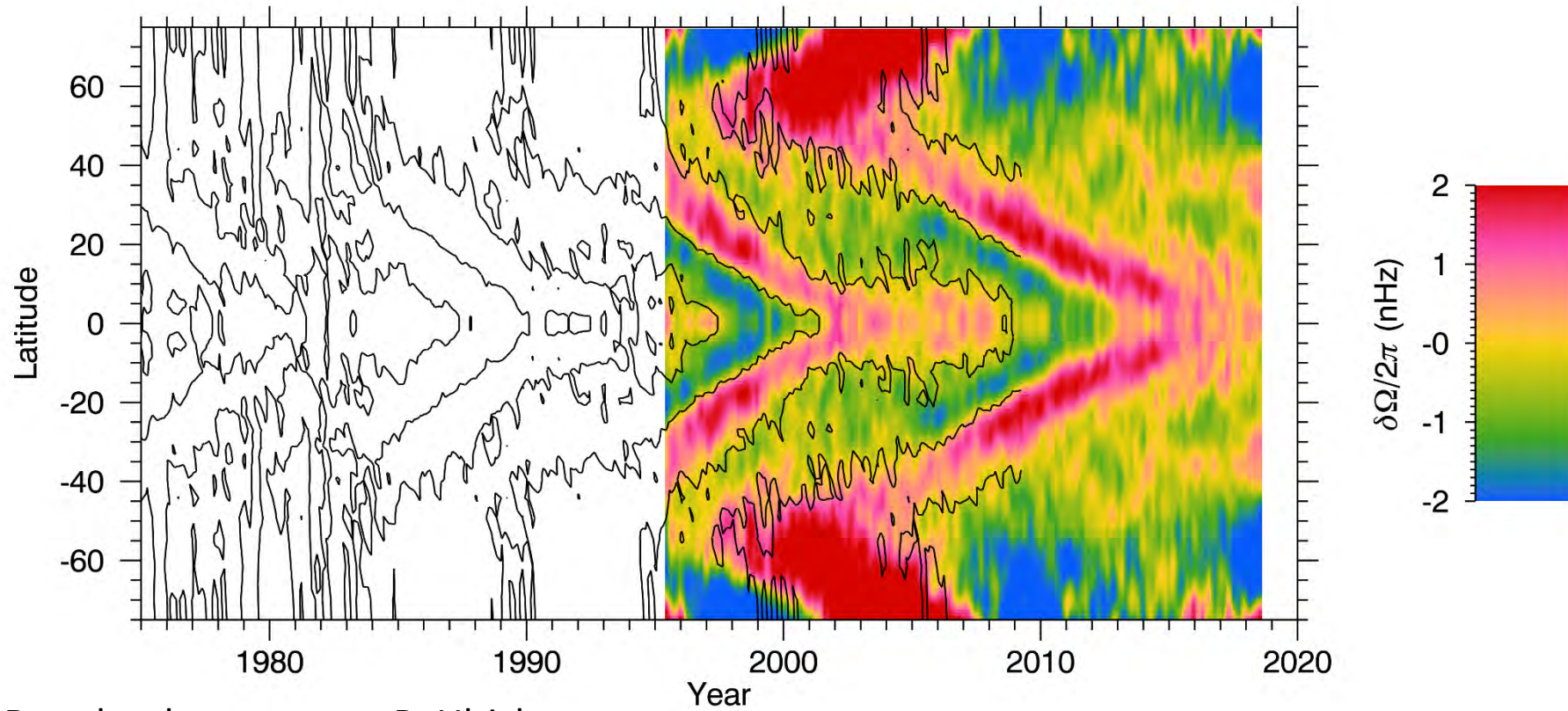


# Historical Context

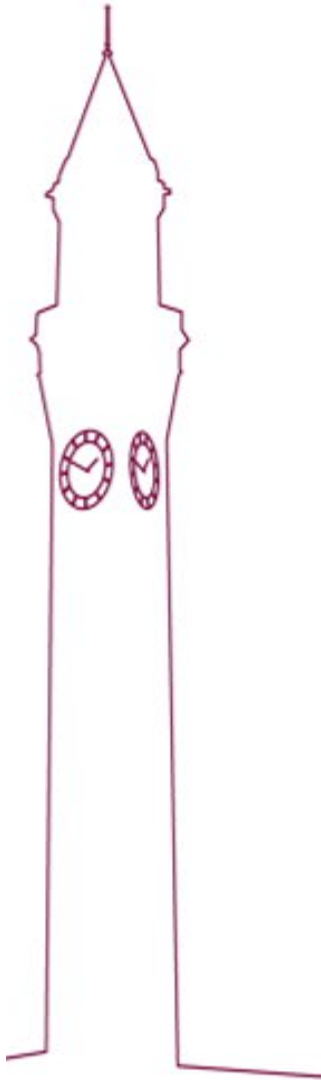




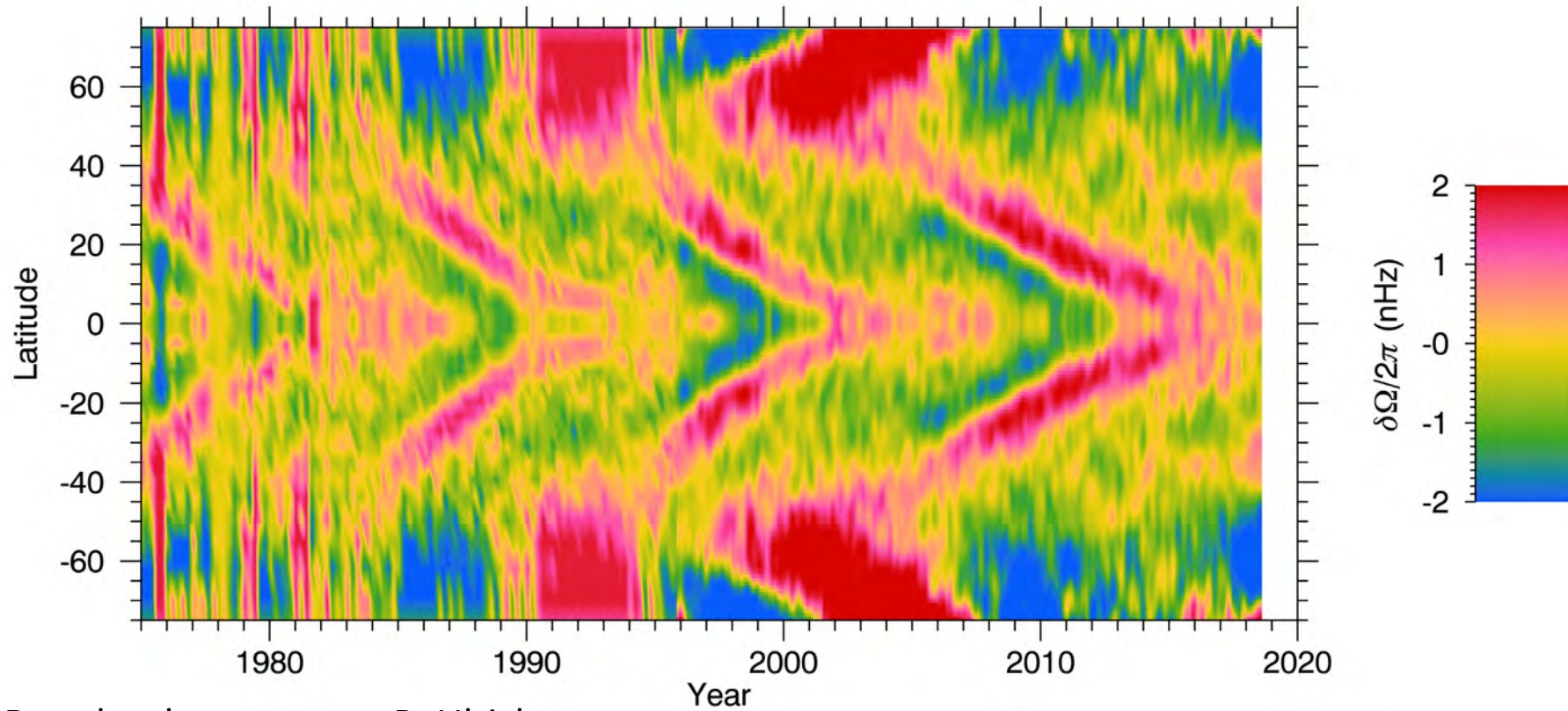
# Historical Context



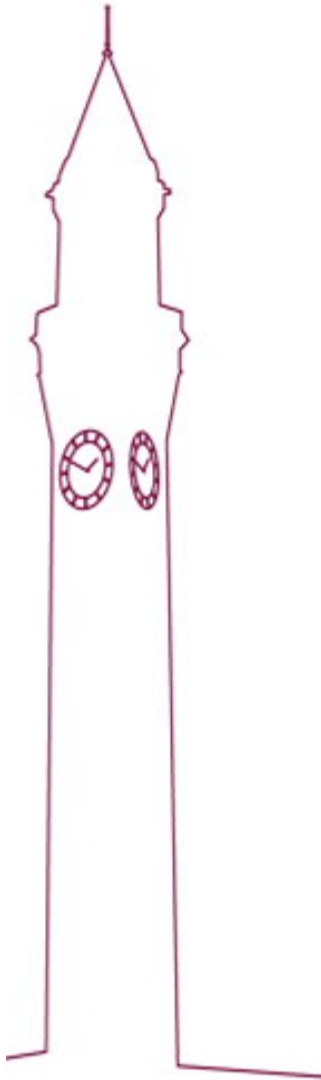
Mount Wilson Doppler data courtesy R. Ulrich



# Historical Context

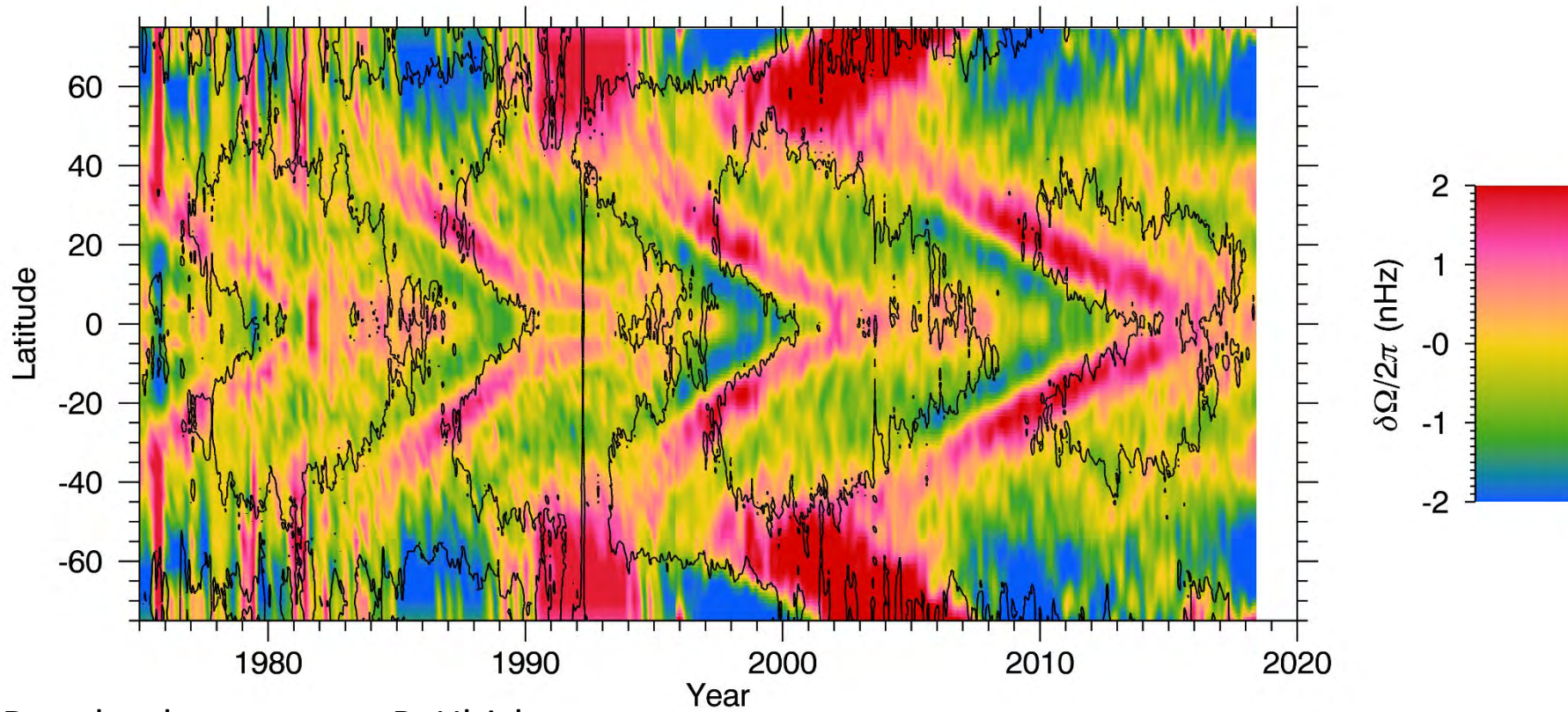


Mount Wilson Doppler data courtesy R. Ulrich

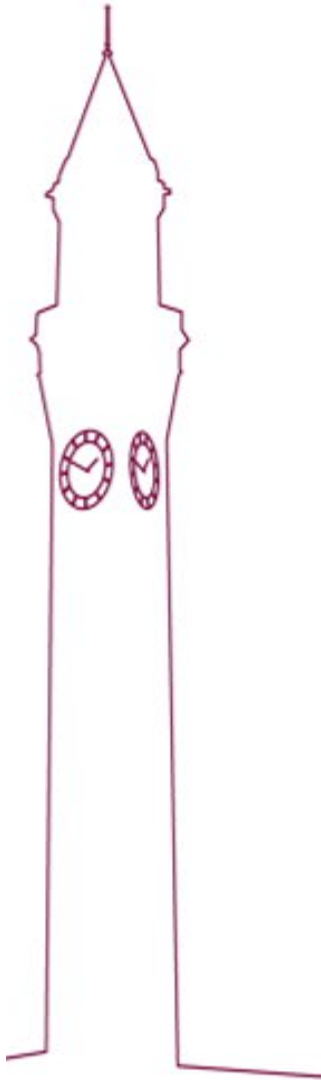




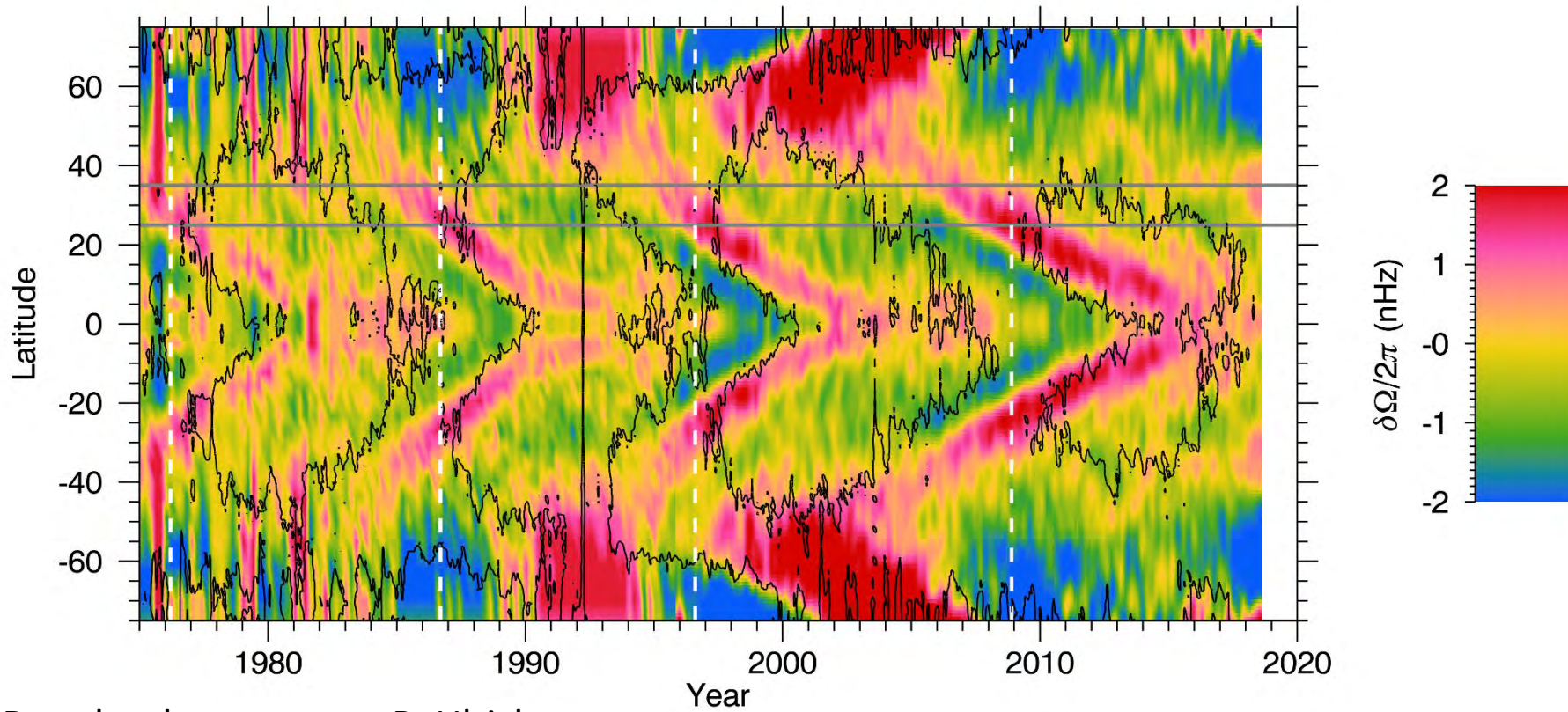
# Historical Context



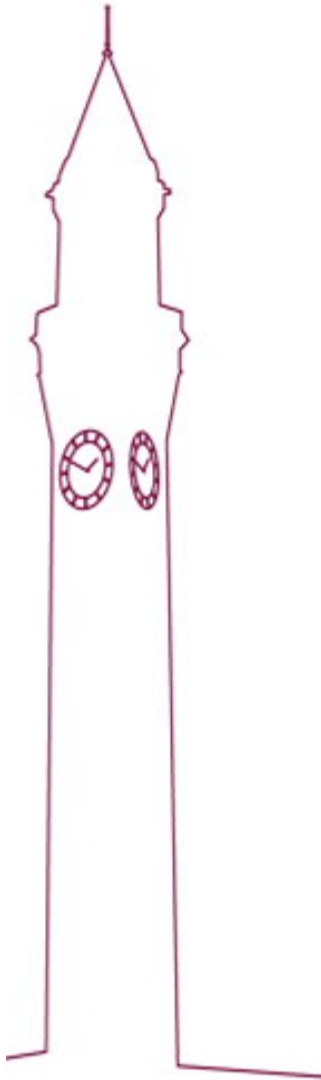
Mount Wilson Doppler data courtesy R. Ulrich



# Historical Context



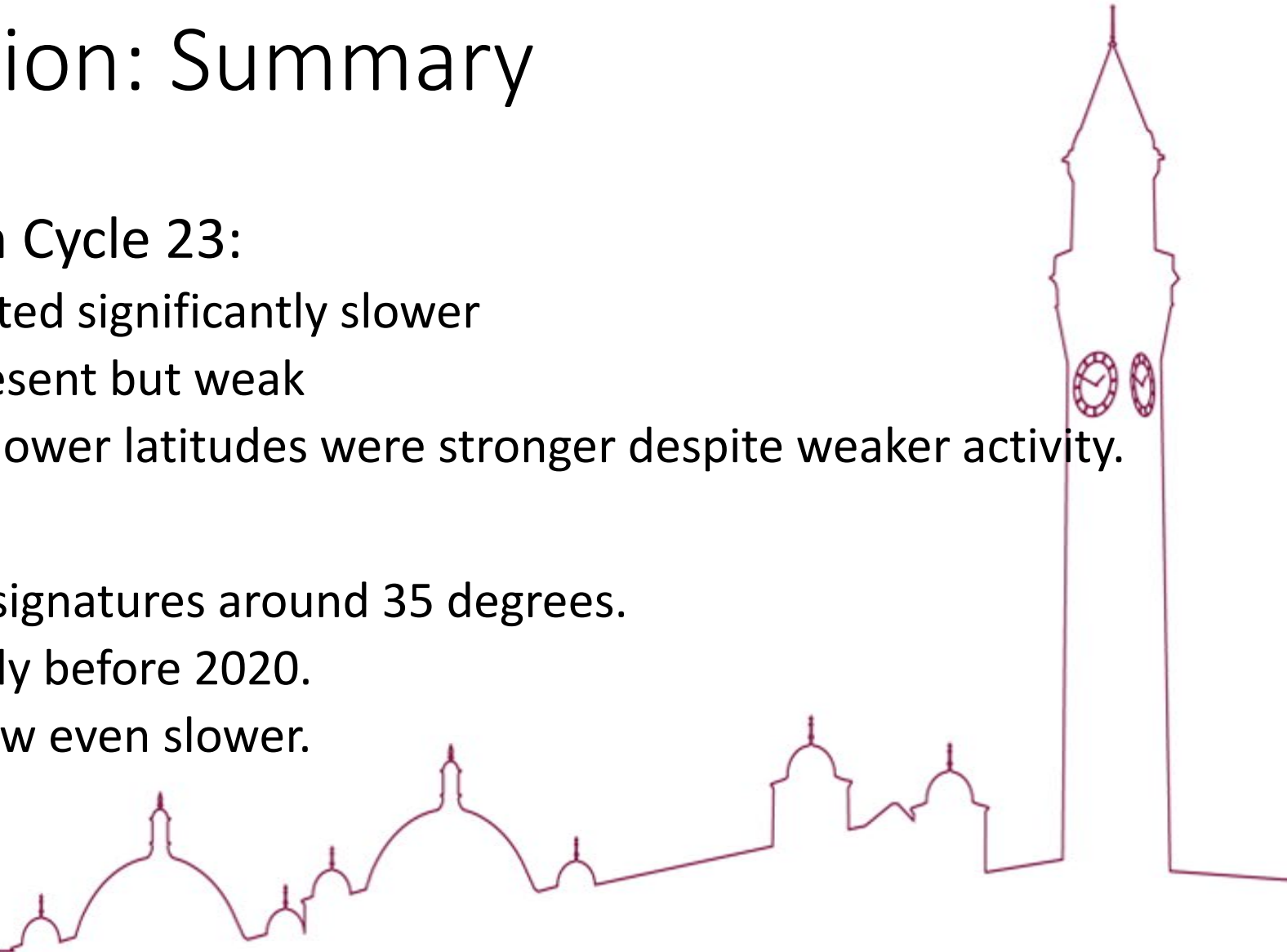
Mount Wilson Doppler data courtesy R. Ulrich





# Torsional Oscillation: Summary

- Comparing Cycle 24 with Cycle 23:
  - The higher latitudes rotated significantly slower
  - Poleward branch was present but weak
  - Faster-rotating bands at lower latitudes were stronger despite weaker activity.
- What about Cycle 25?
  - We are currently seeing signatures around 35 degrees.
  - New-cycle activity unlikely before 2020.
  - High-latitude rotation now even slower.



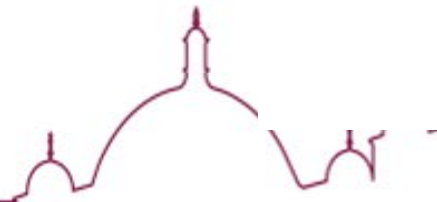
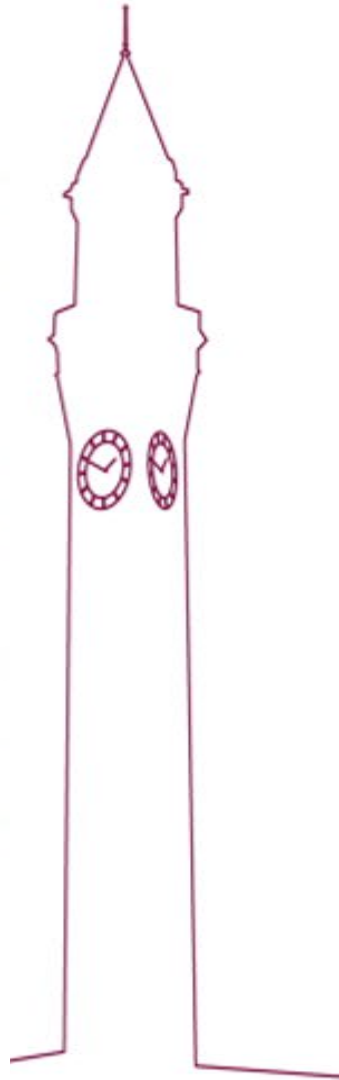
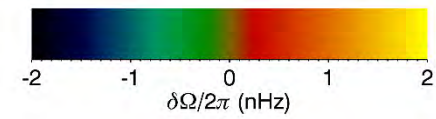
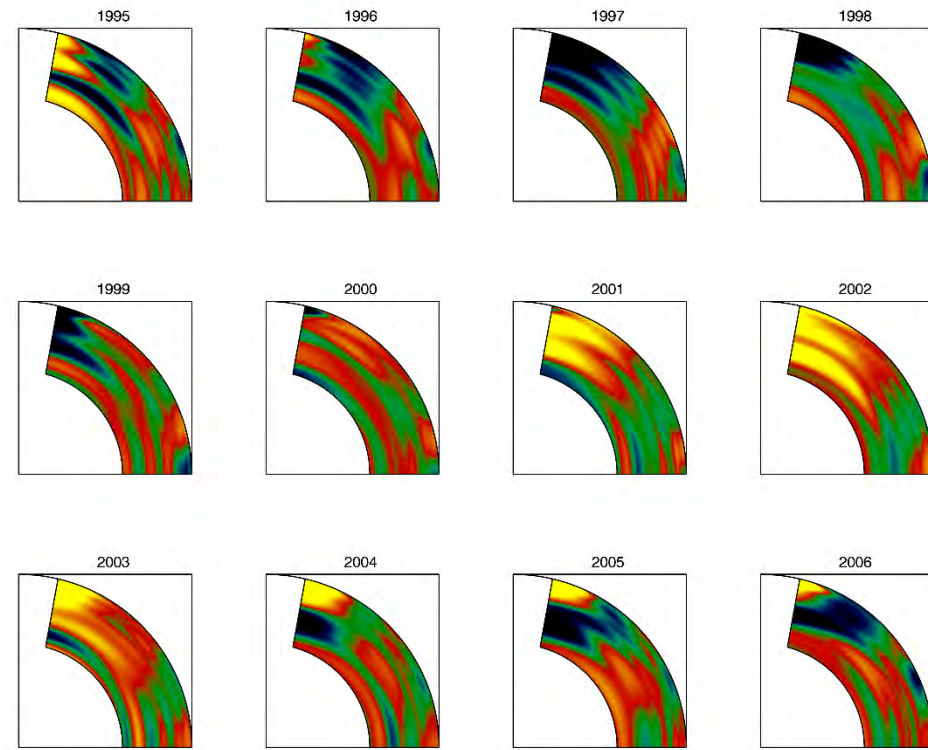
# Conclusions

- Changes in the sensitivity of mode frequencies to activity hint at a thinner near-surface magnetic layer.
- Polar fields are weaker and this has been linked to slower rotation at high latitudes.
- Could these be signs that the solar dynamo is about to change to the weaker mode associated with aging stars? Maybe?
- Or it could 'just' be a grand minimum (or at least the end of a grand maximum?)



# Yearly average residuals 1995-2006

Mean for Cycle 23



# Yearly average residuals 2007-2018

Mean for Cycle 24

