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THE SUBTERAHERTZ SUN

EQUATORIAL AND POLAR RADII FROM SST AND ALMA



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U. P. MACKENZIE - CRAAM

THE SUBTERAHERTZ SUN

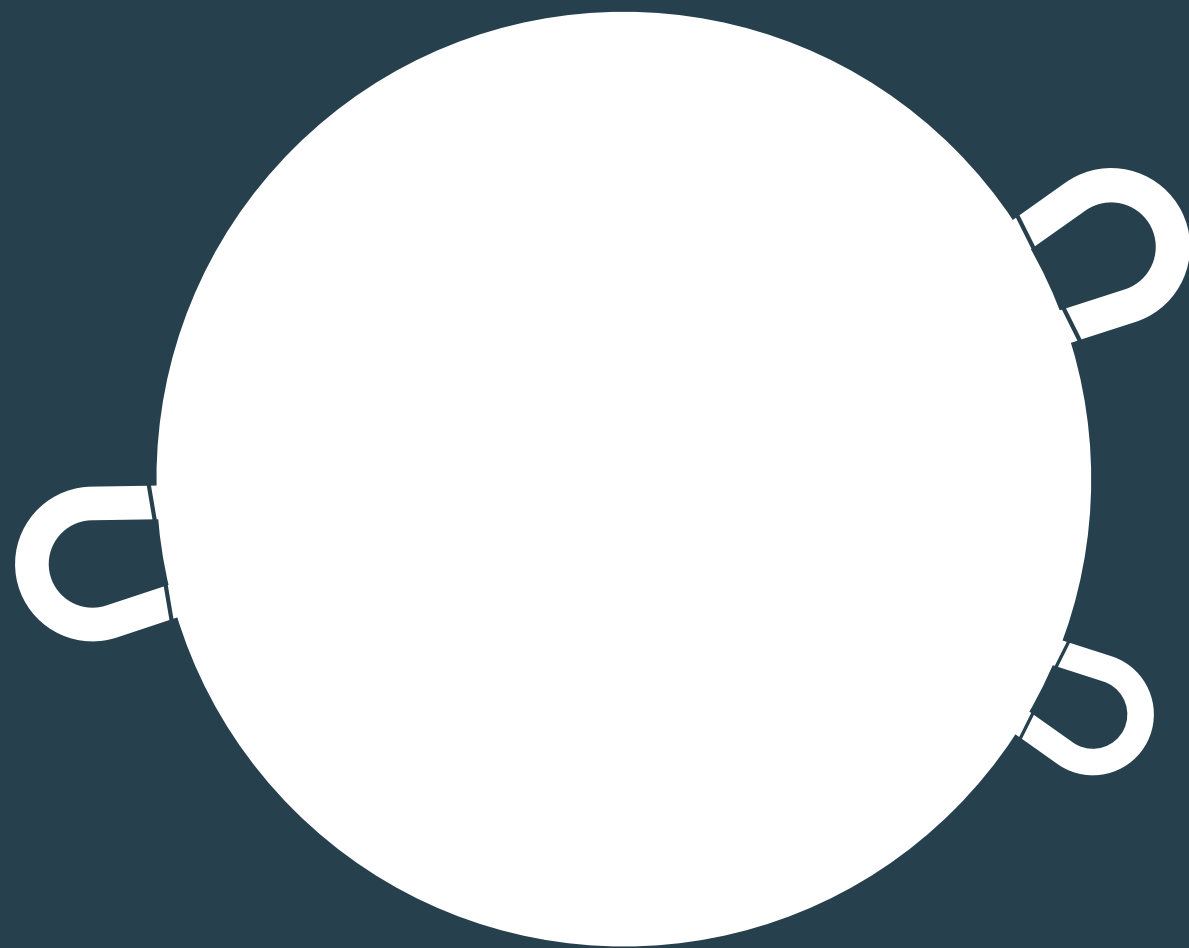
EQUATORIAL AND POLAR RADII FROM SST AND ALMA

Why?

- ▶ Related to solar activity
- ▶ Its variations over time indicate changes in the solar atmosphere
- ▶ Important parameter to improve solar atmosphere models
- ▶ Gap of subterahertz frequencies in the measurements of the solar radius and other parameters of the atmosphere

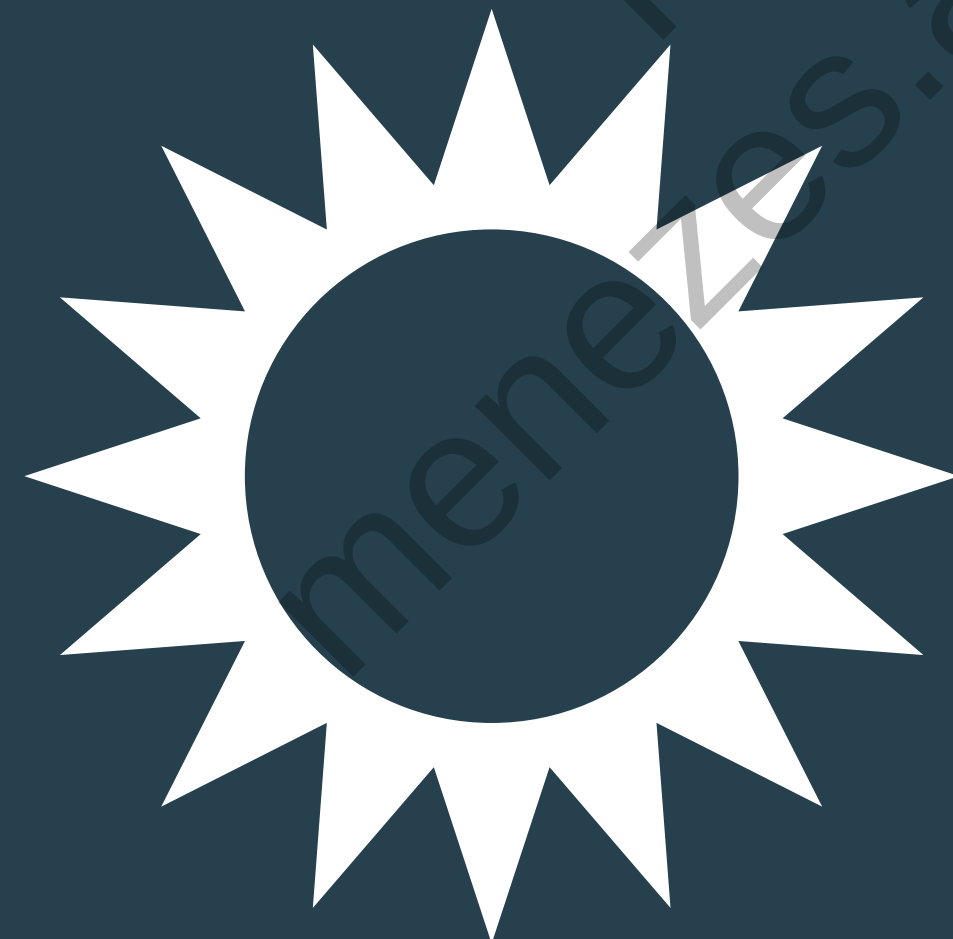
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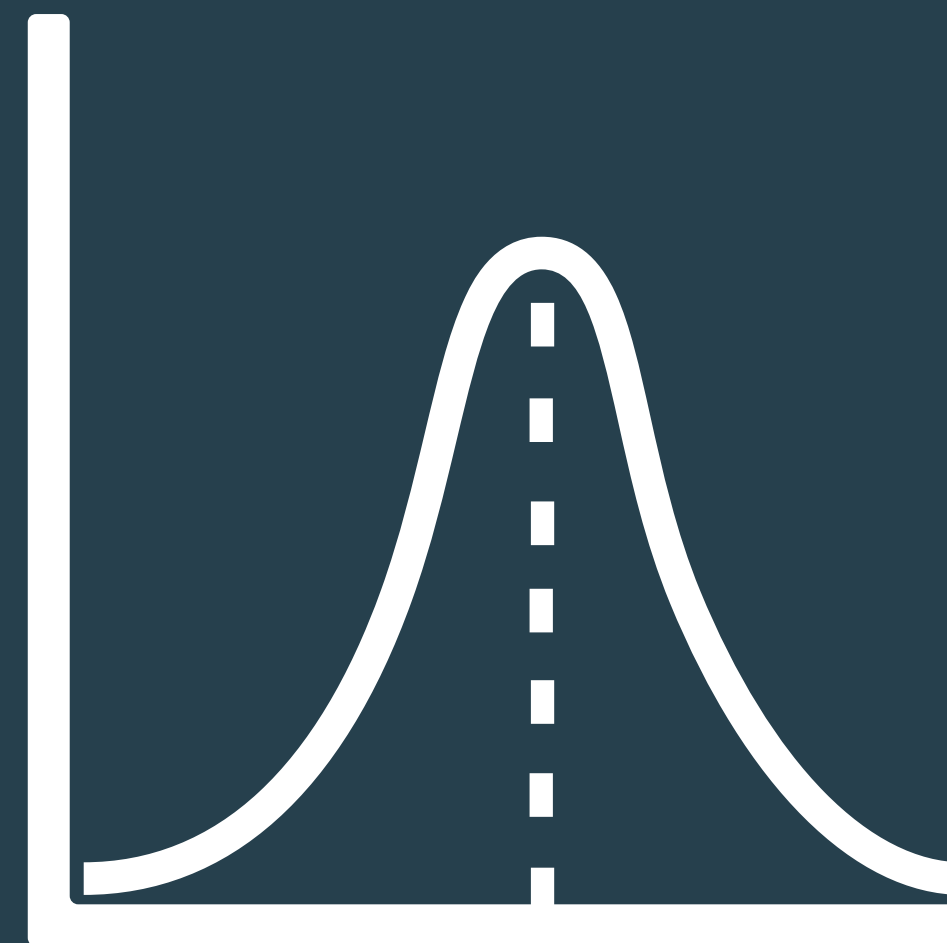
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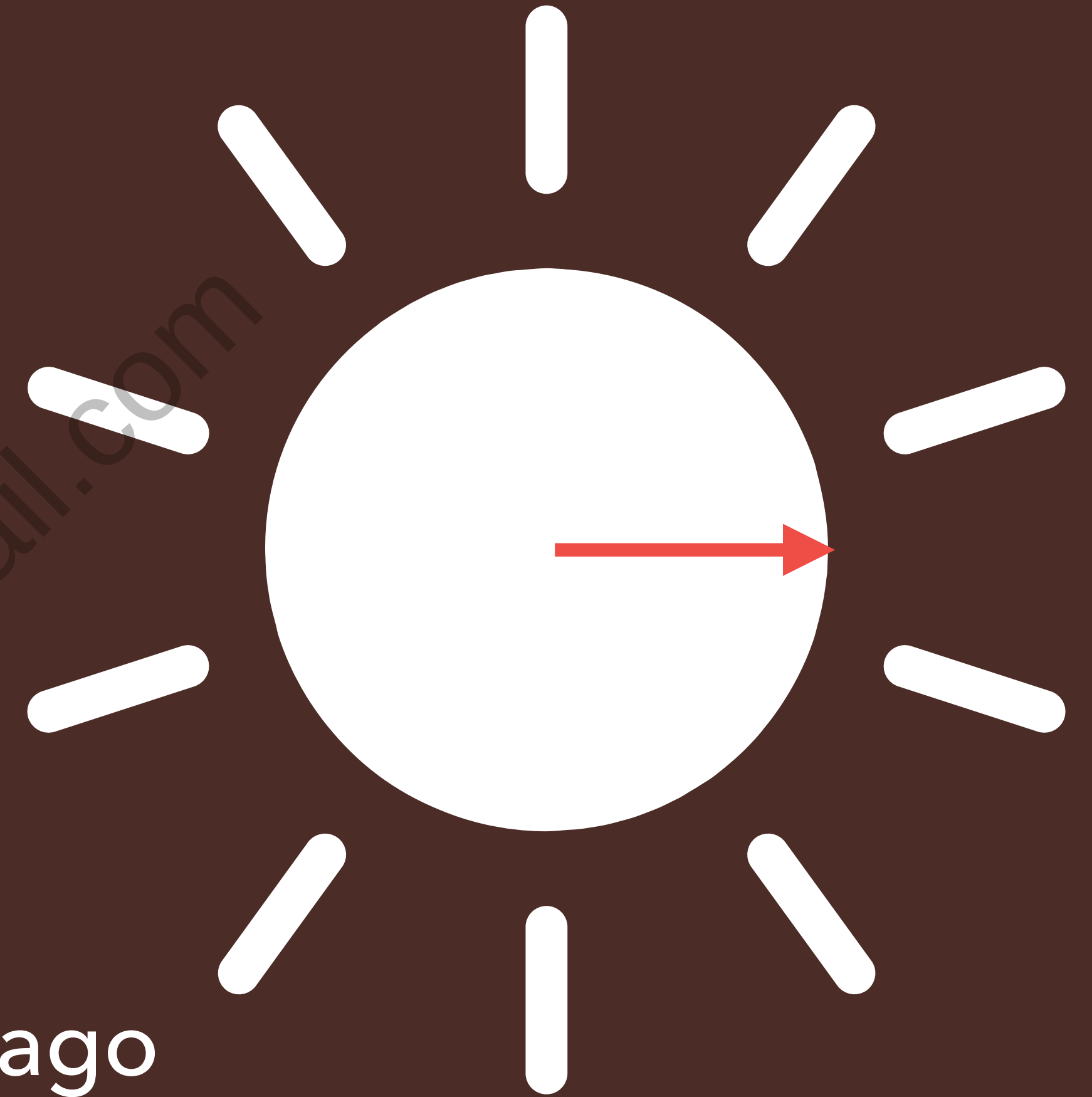
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Solar Radius

- ▶ $R_{\odot}^N = 6.957 \times 10^8 \text{ m}$
- ▶ Centuries of measures
 - ▶ Vaquero et al. (2016) – 233 years
 - ▶ Gilliland (1981) – 265 years
- ▶ Radio wavelengths started some decades ago



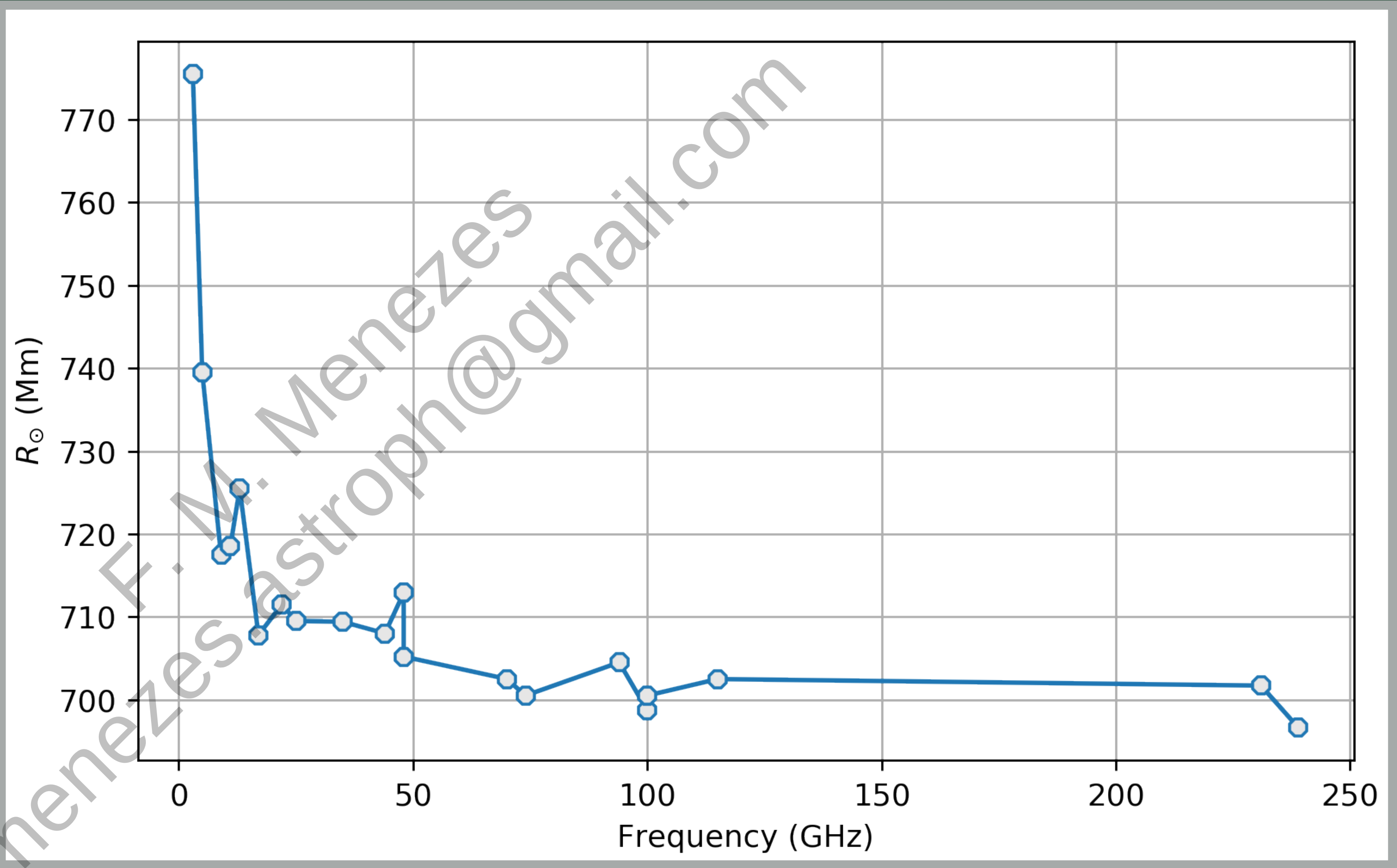
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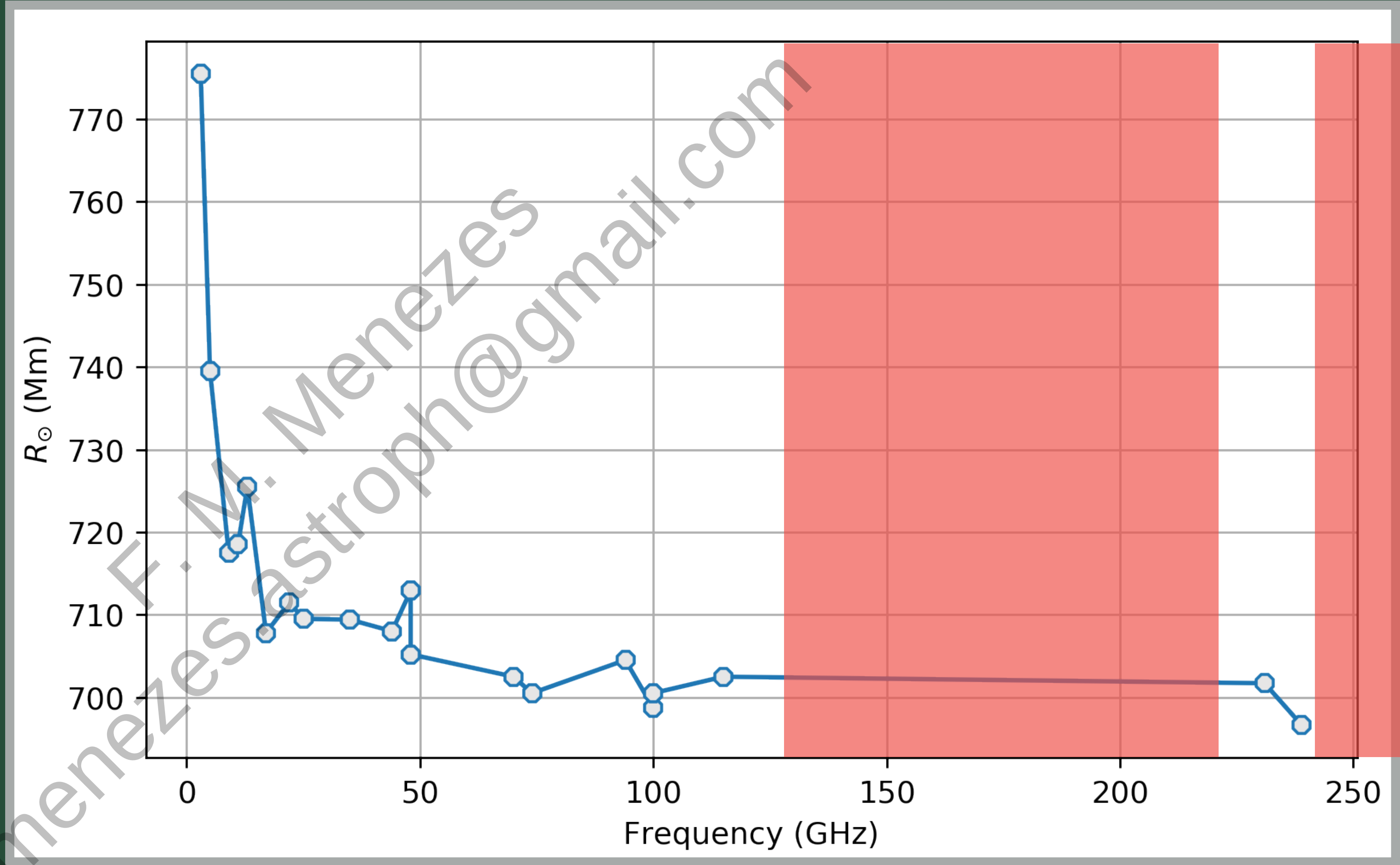
Solar Radius

ν (GHz)	R (Mm)	ν (GHz)	R (Mm)
11	719	48 ^b	713.4
16	718	70	703
17	708.3	74	701
22	712	94	705
25	710	100 ^c	699.2
30	710.0	100 ^d	701
35	709.9	231	702.2
44	708.5	239	697.1
48 ^a	705.7		



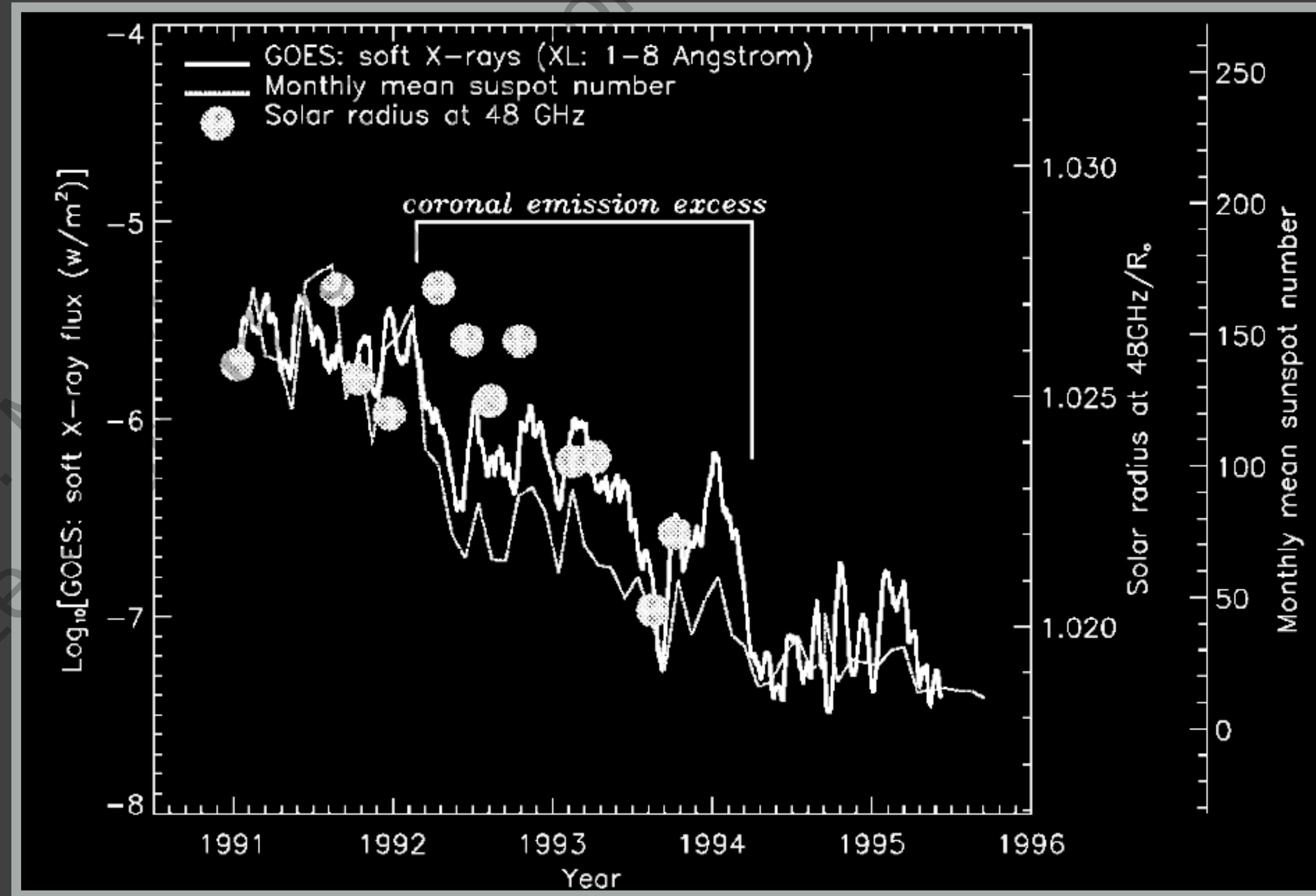
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Correlation to solar activity

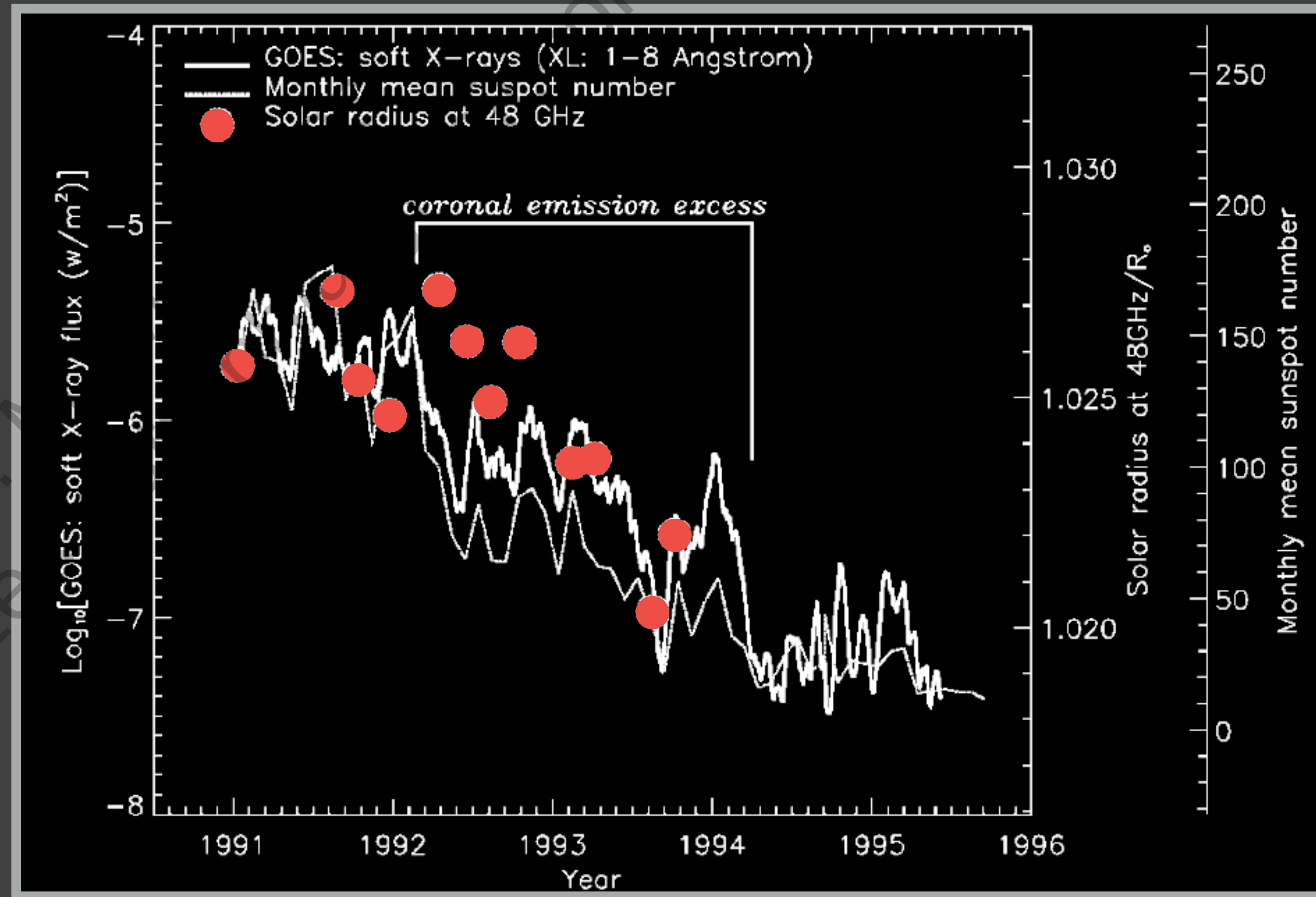
- ▶ **Costa et al. (1999)**
- ▶ Rádio Observatório de Itapetinga (**ROI**)
- ▶ 48 GHz single-dish
- ▶ 1991 – 1993 (3 yrs)



Correlation to solar activity

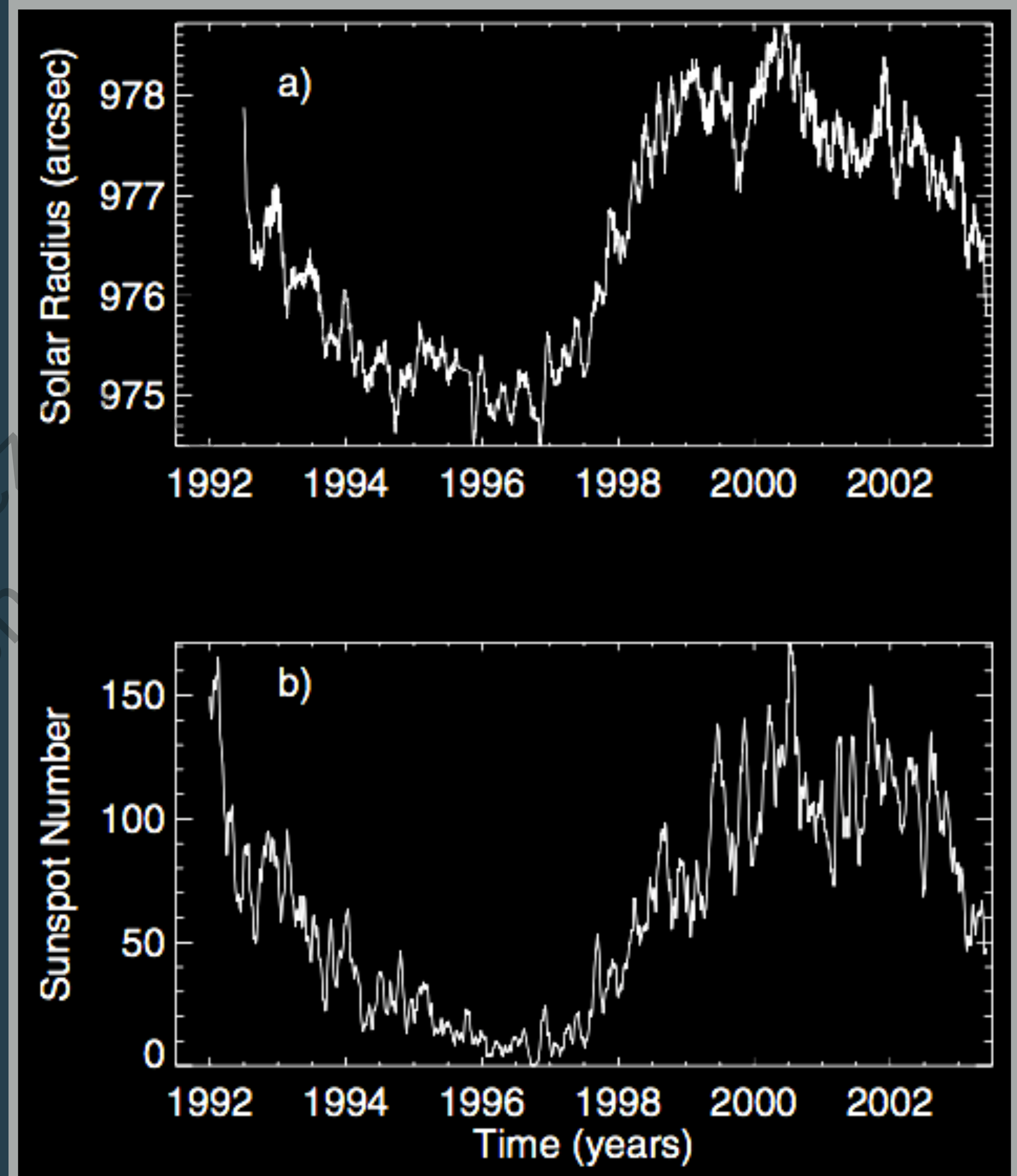
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$$\frac{R_{48}}{R_{\odot}} = 1.029 - 0.0015(\text{year} - 1990)$$



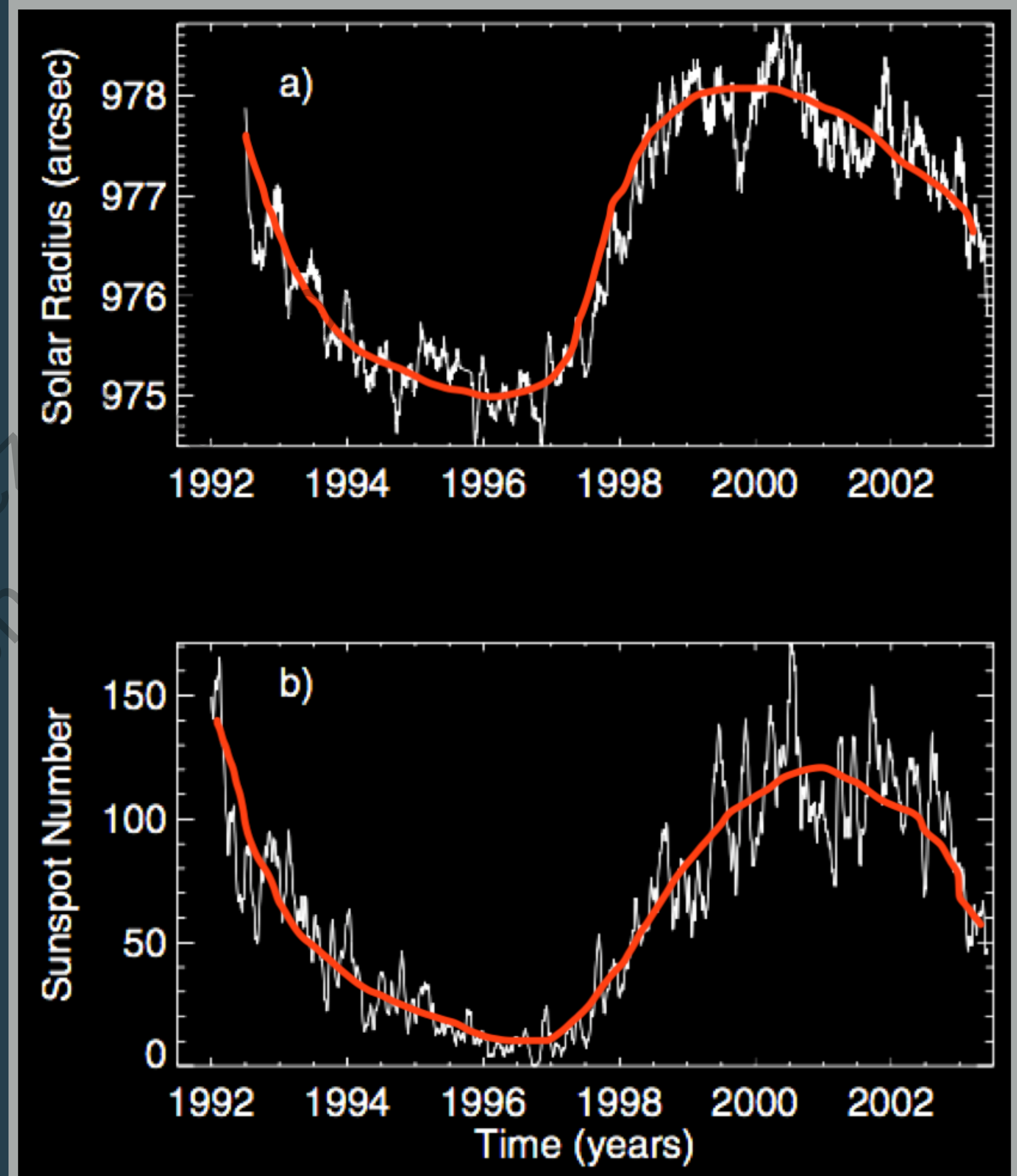
Correlation to solar activity

- ▶ **Selhorst et al. (2004)**
 - ▶ 17 GHz from NoRH
 - ▶ 1992 – 2003 (1 cycle)
 - ▶ Correlation coefficient: 0.88
 - ▶ Polar radius: corr. coef.: -0.64



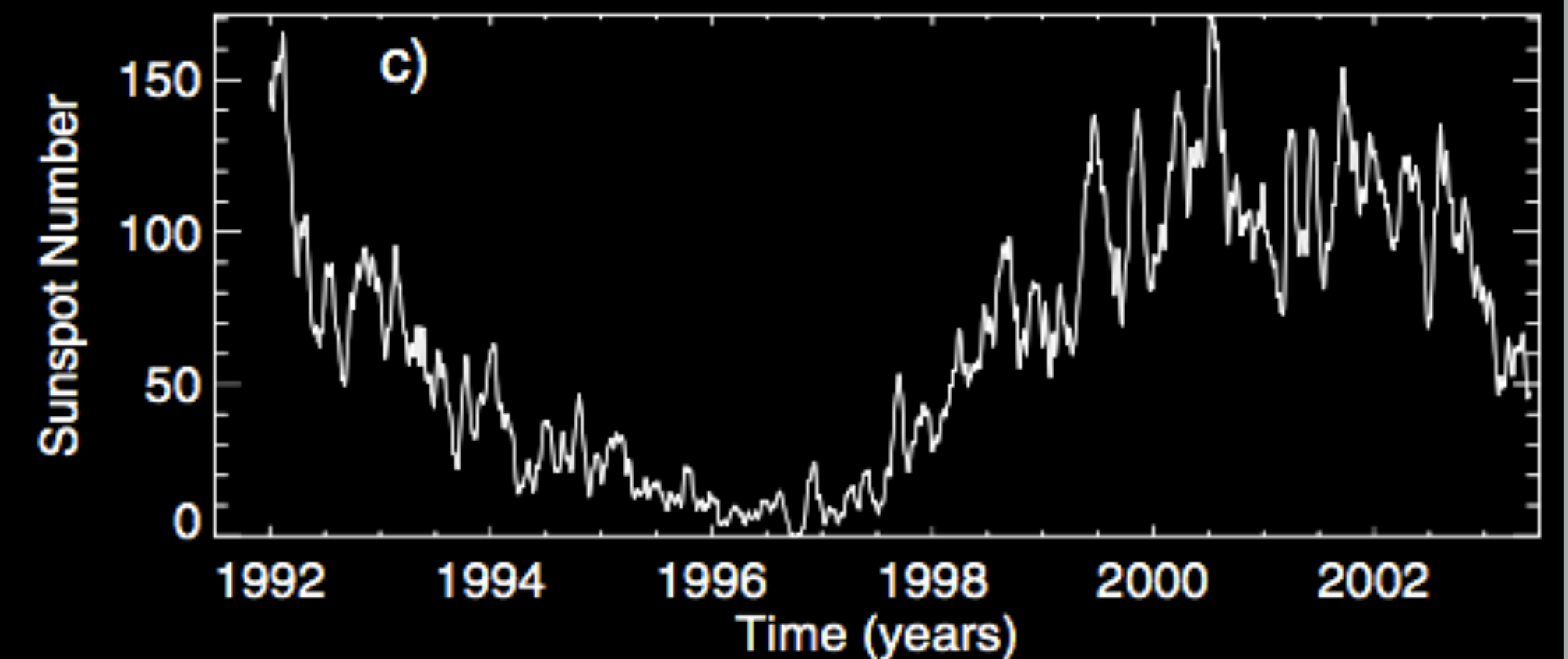
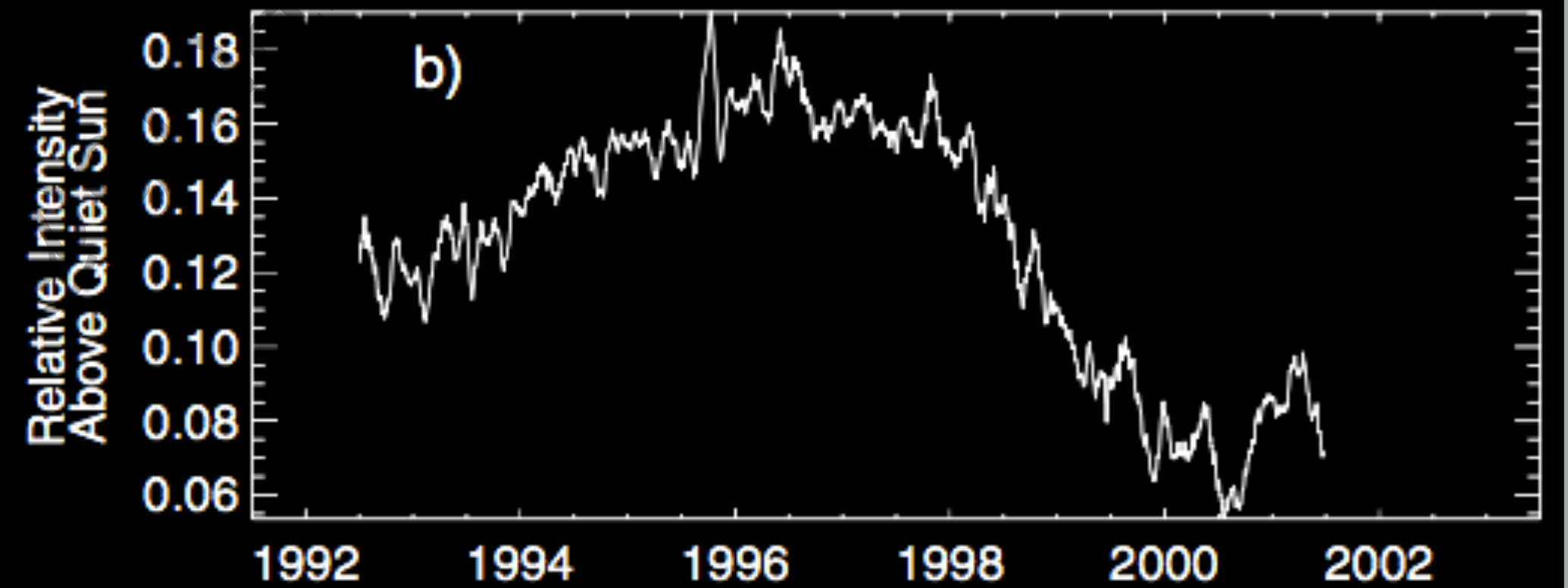
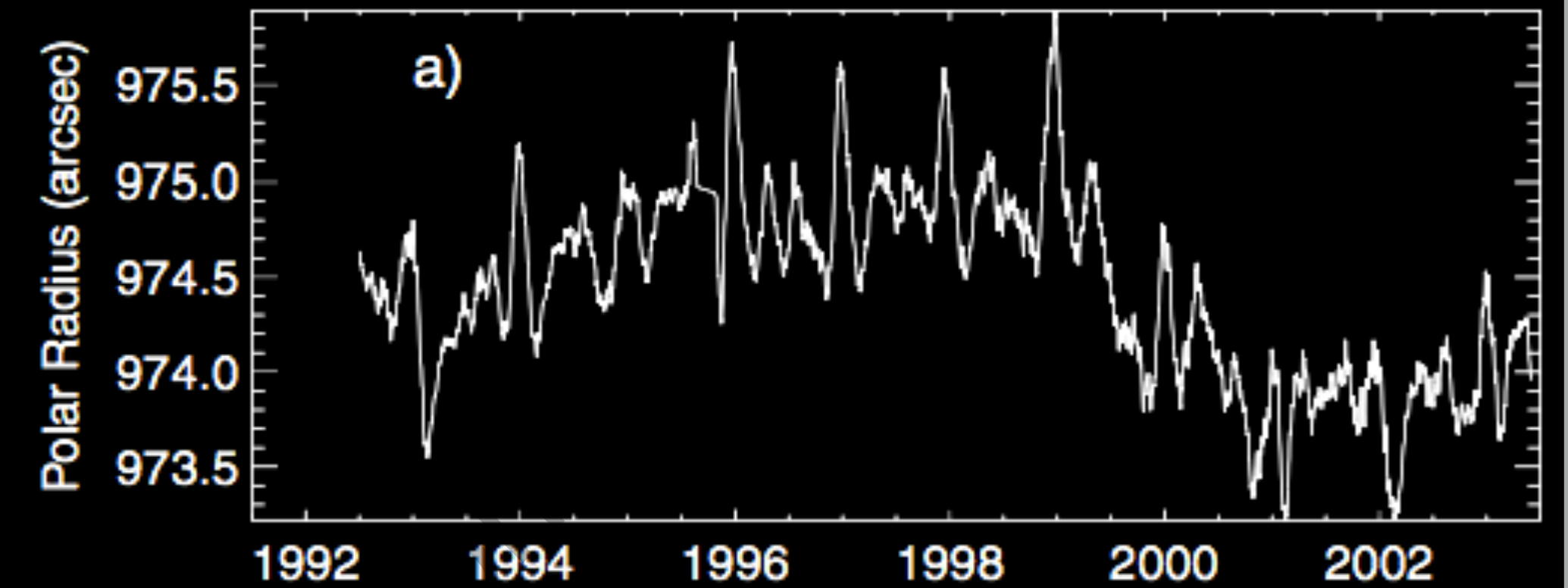
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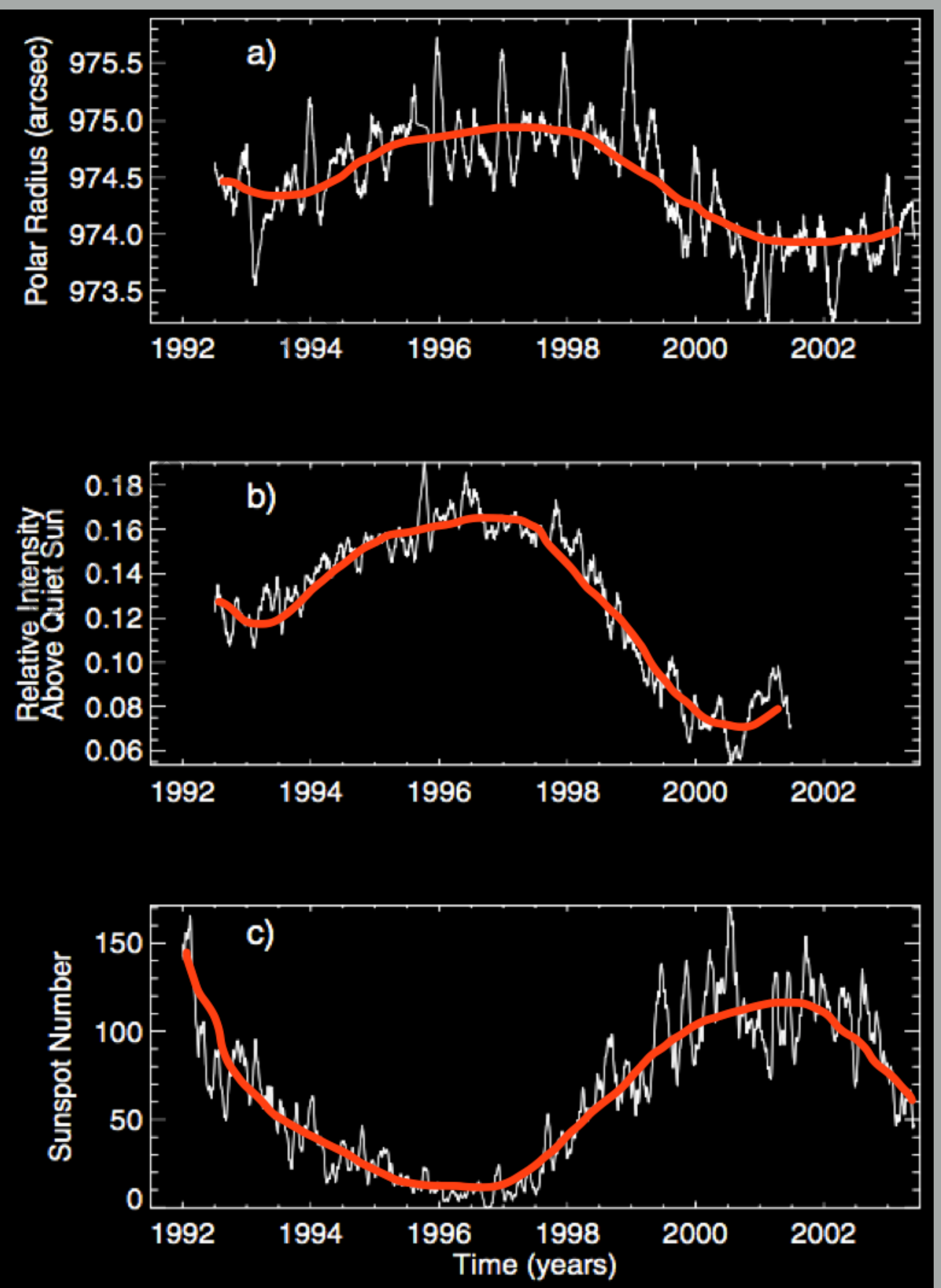
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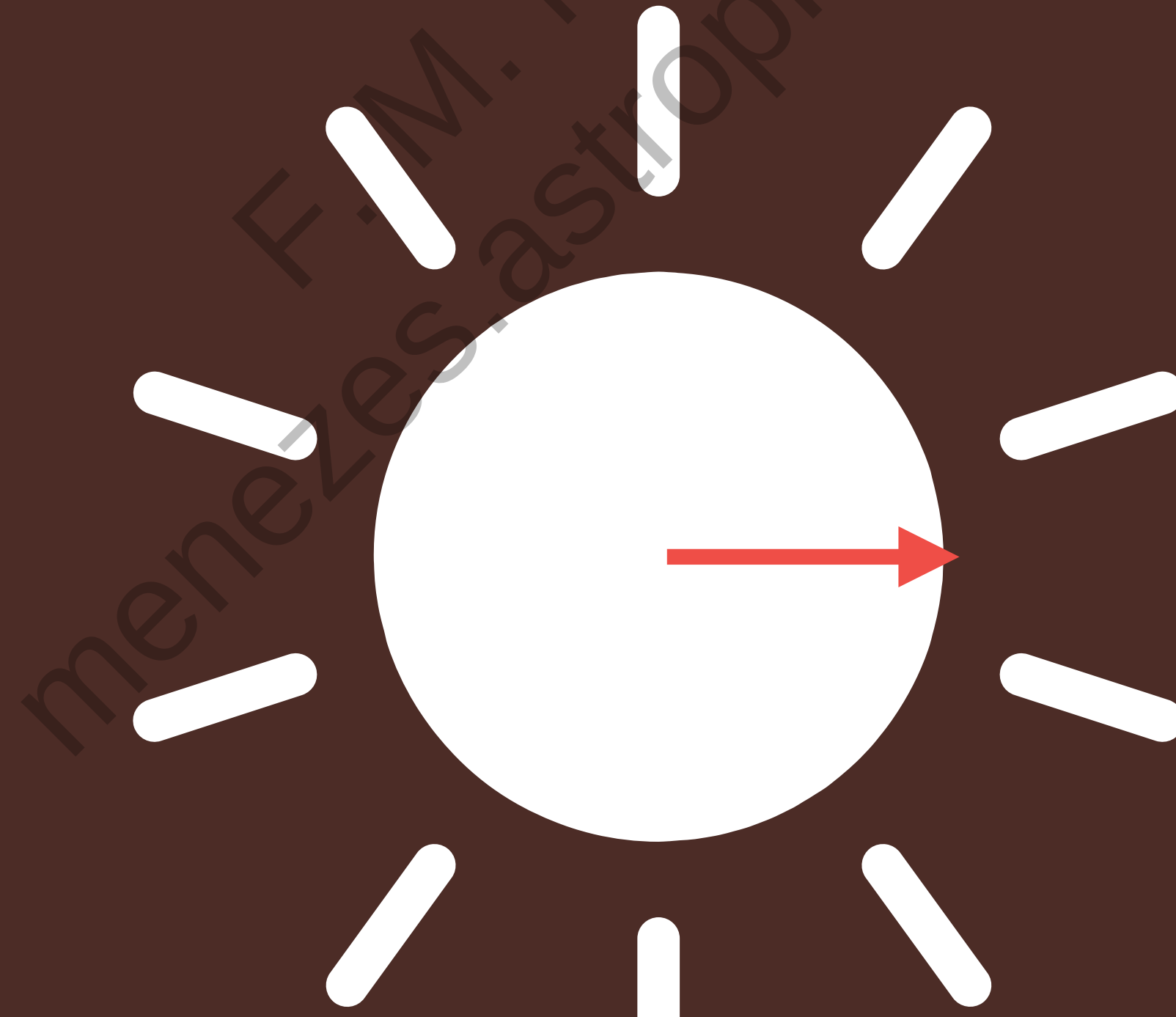


Goals

- ▶ Average solar radius at 0,212 e 0,405 THz
- ▶ Height above photosphere where these emissions are being mainly created
- ▶ Variation of the solar radius in relation to the solar cycle

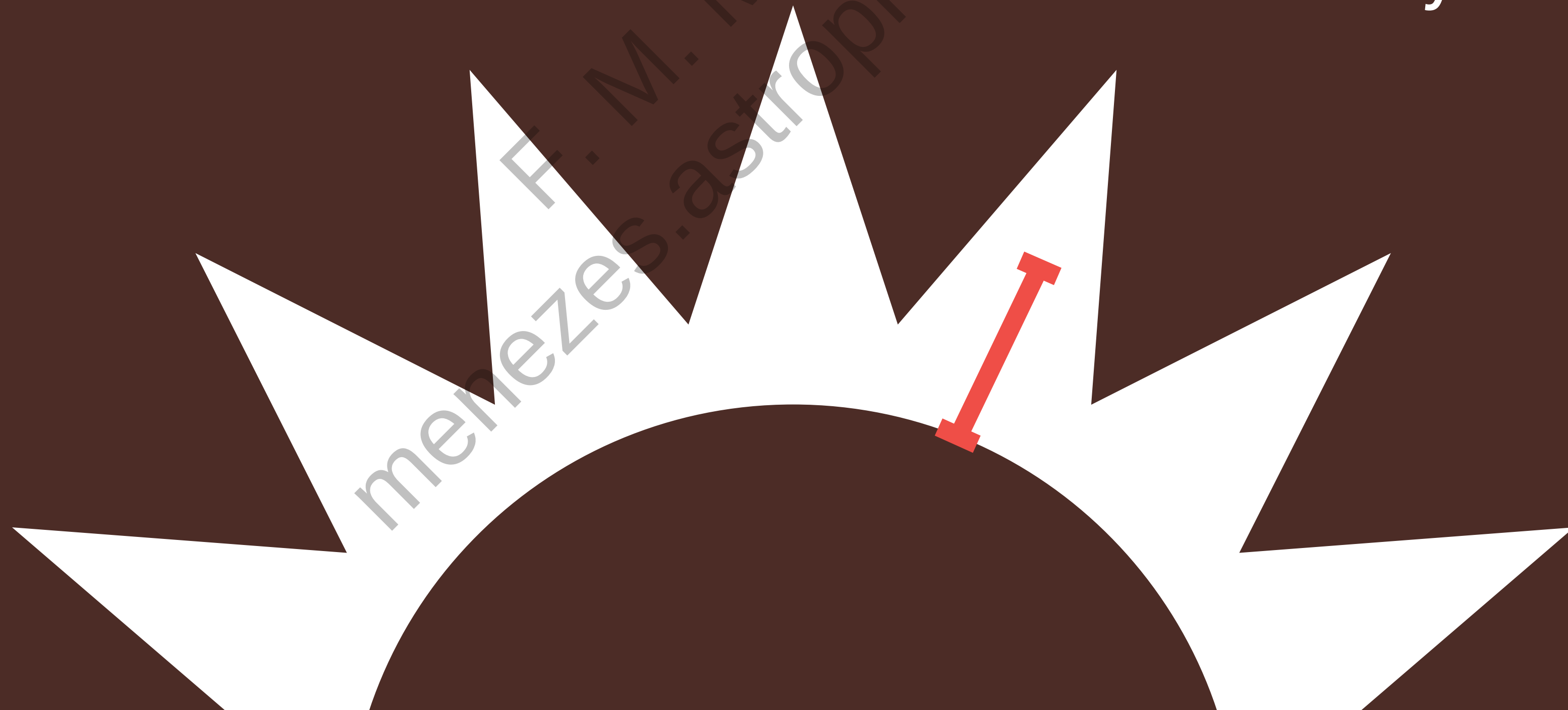
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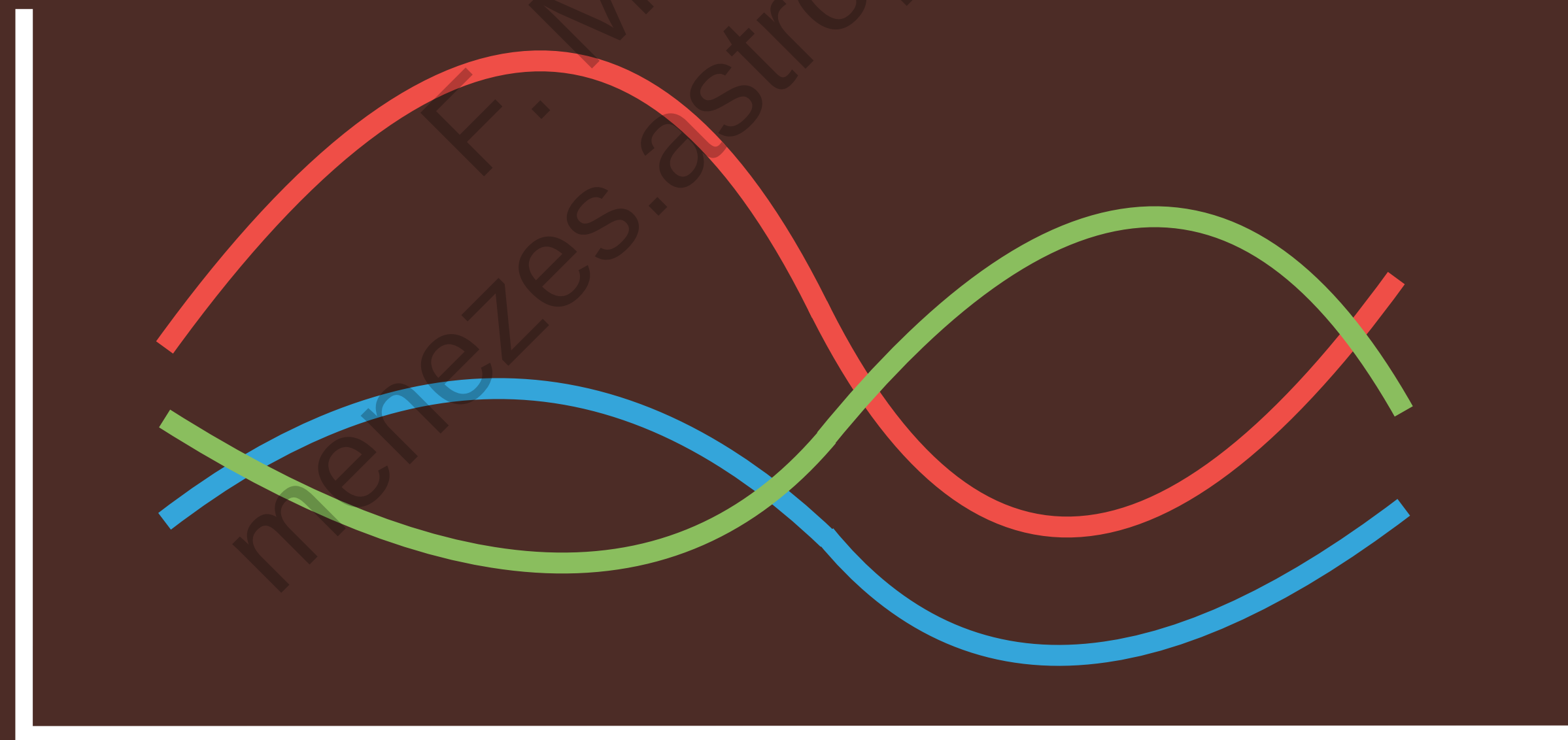
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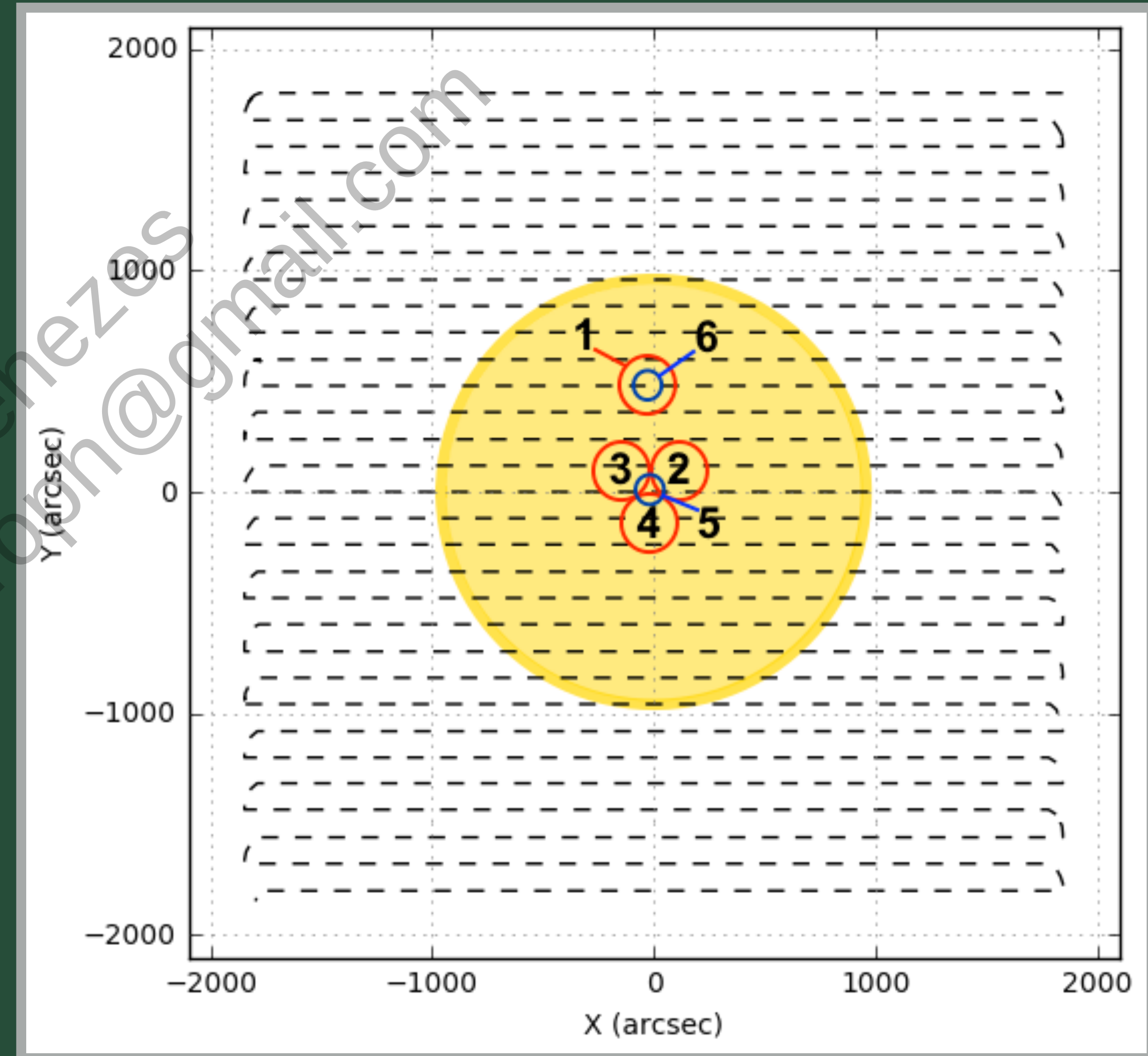
Solar Submillimeter Telescope

- ▶ CASLEO - San Juan, Argentina
(Andes, 2550 m)
- ▶ from 1999 to 2017
- ▶ 0.212 THz: 4 beams, HPBW = 4'
- ▶ 0.405 THz: 2 beams, HPBW = 2'
- ▶ Time resolution: 5 millisec.

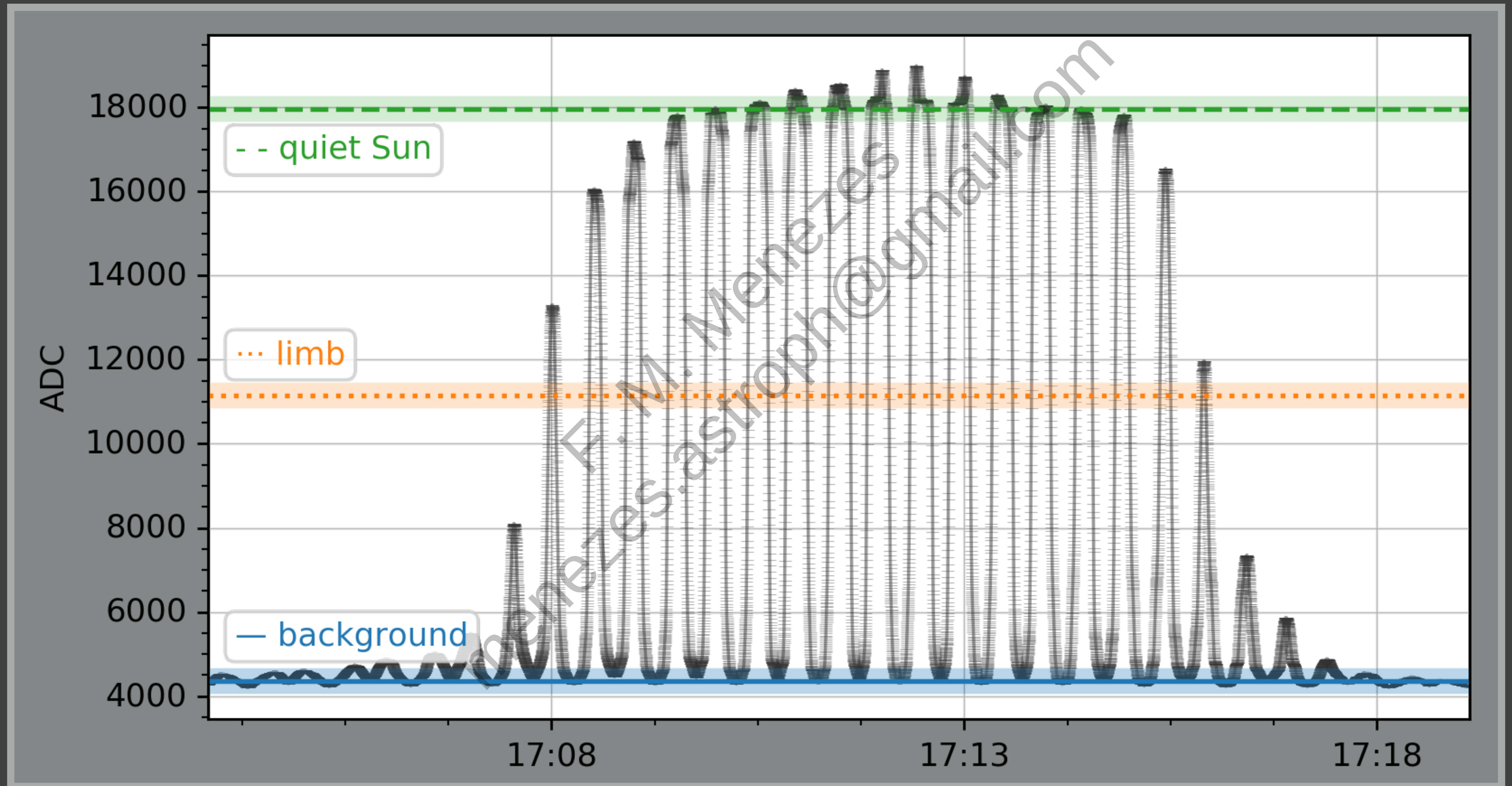


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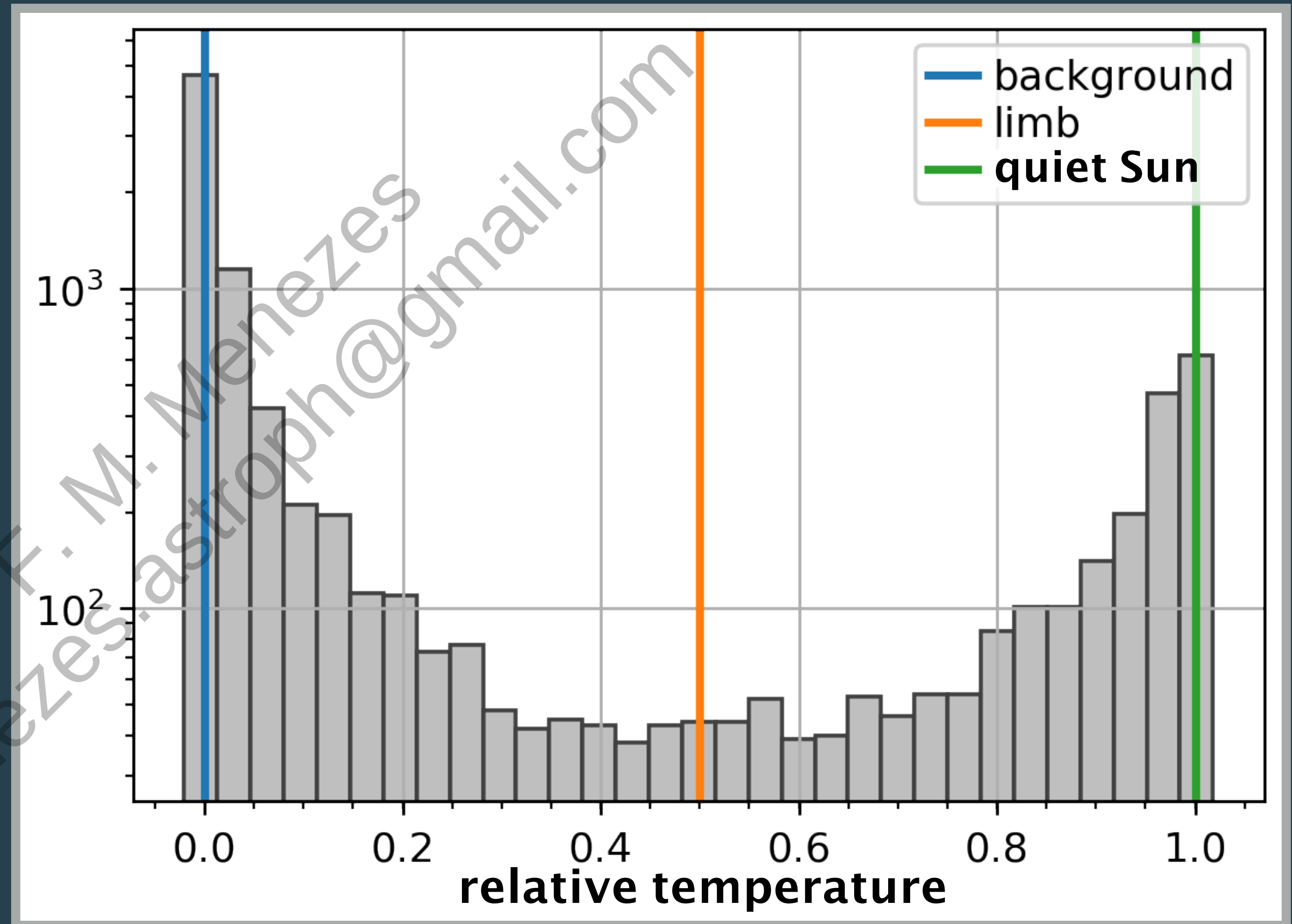


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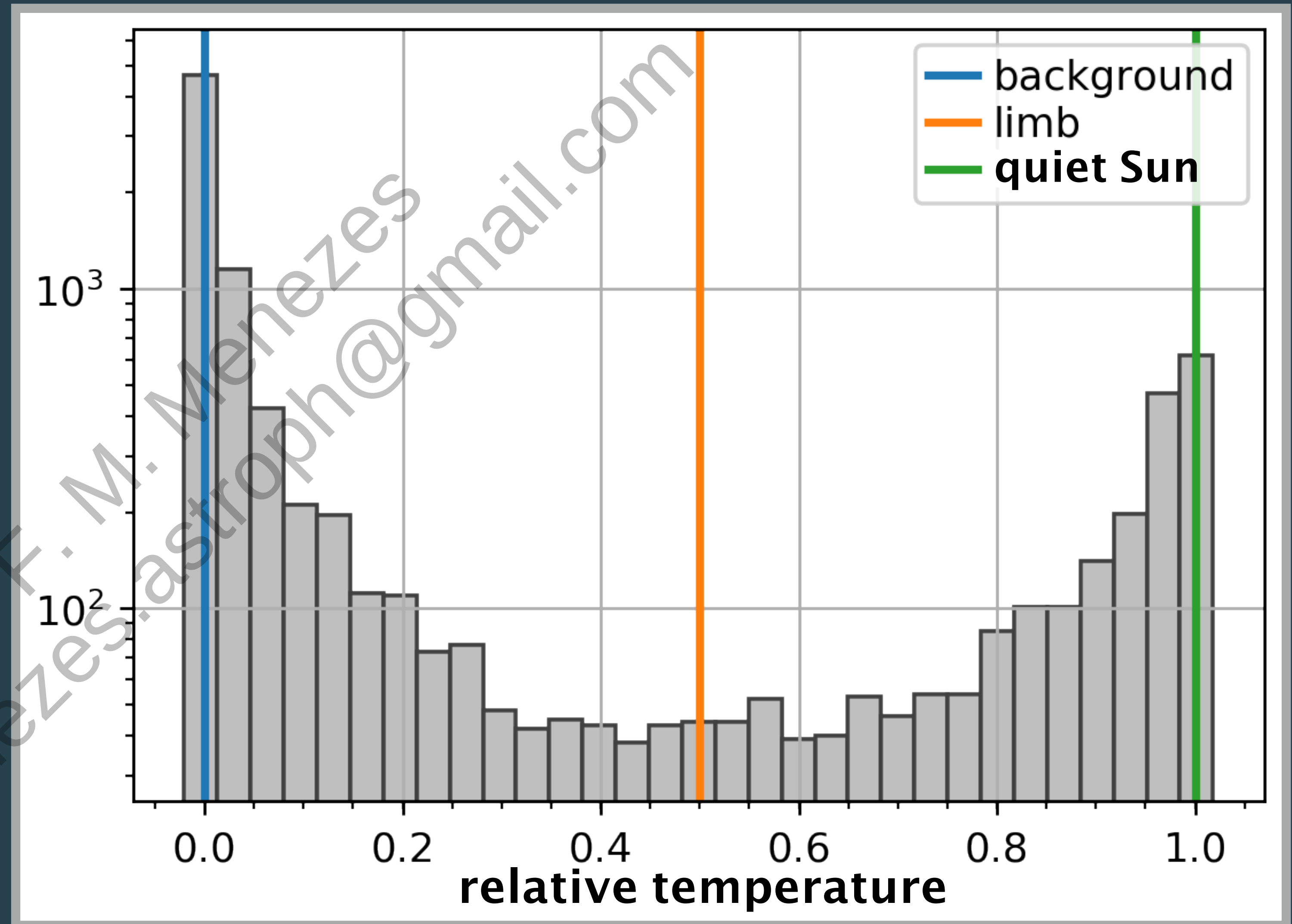
Background, quiet sun & limb levels

- ▶ Background is the most common value (mode)
- ▶ Quiet Sun level is the most common value within the solar disk (second peak in the histogram)
- ▶ Limb level is the half of the quiet Sun level



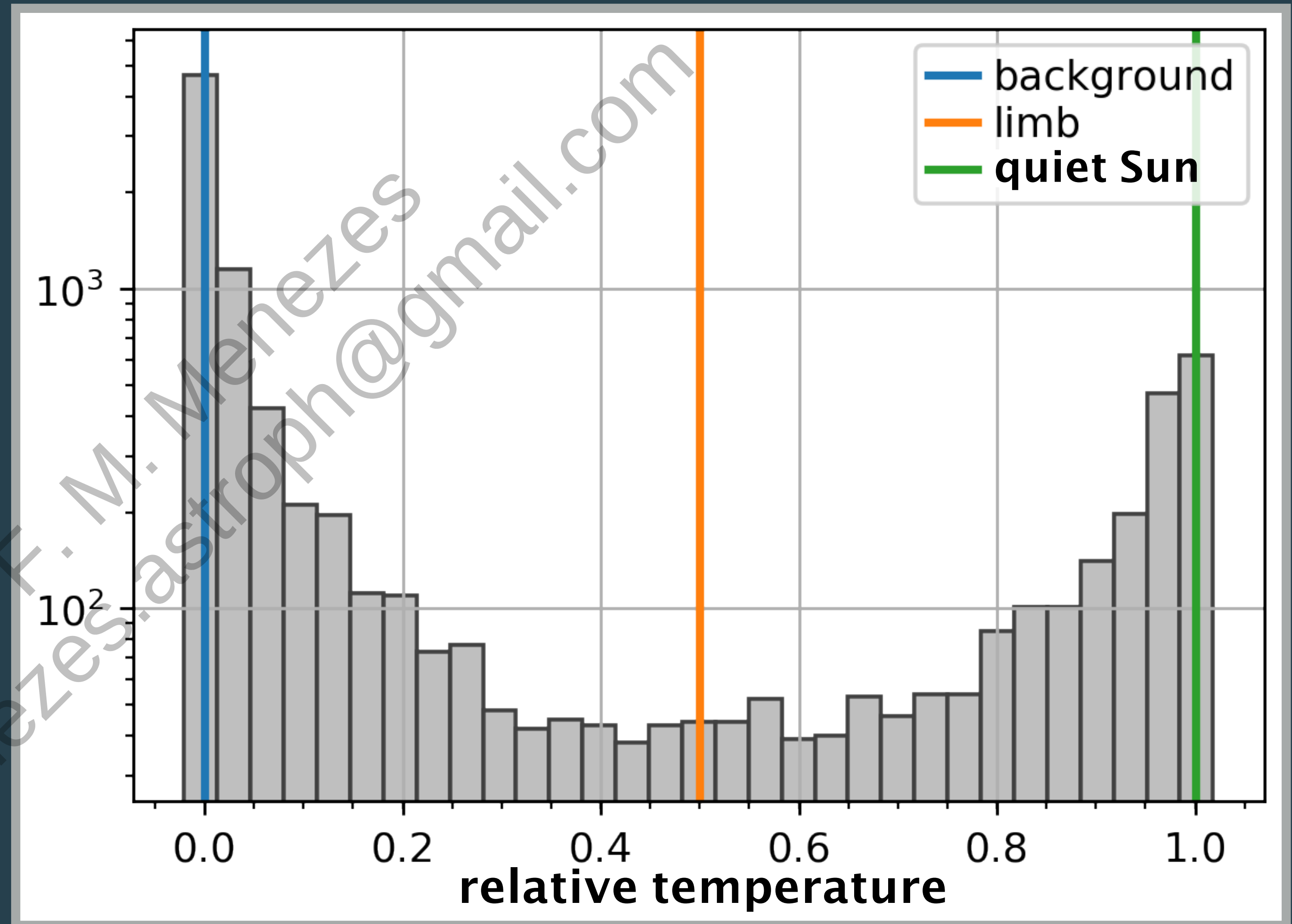
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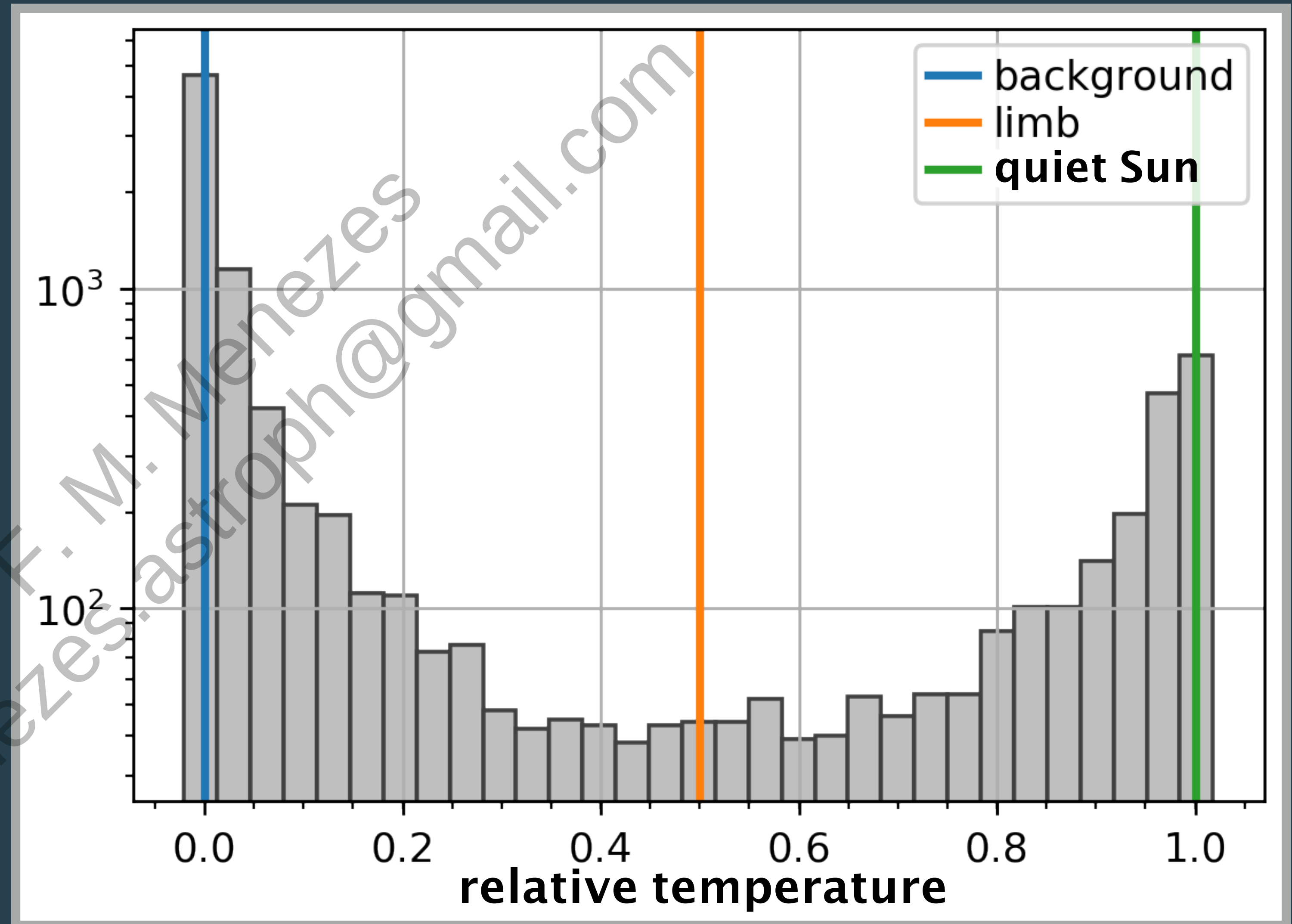
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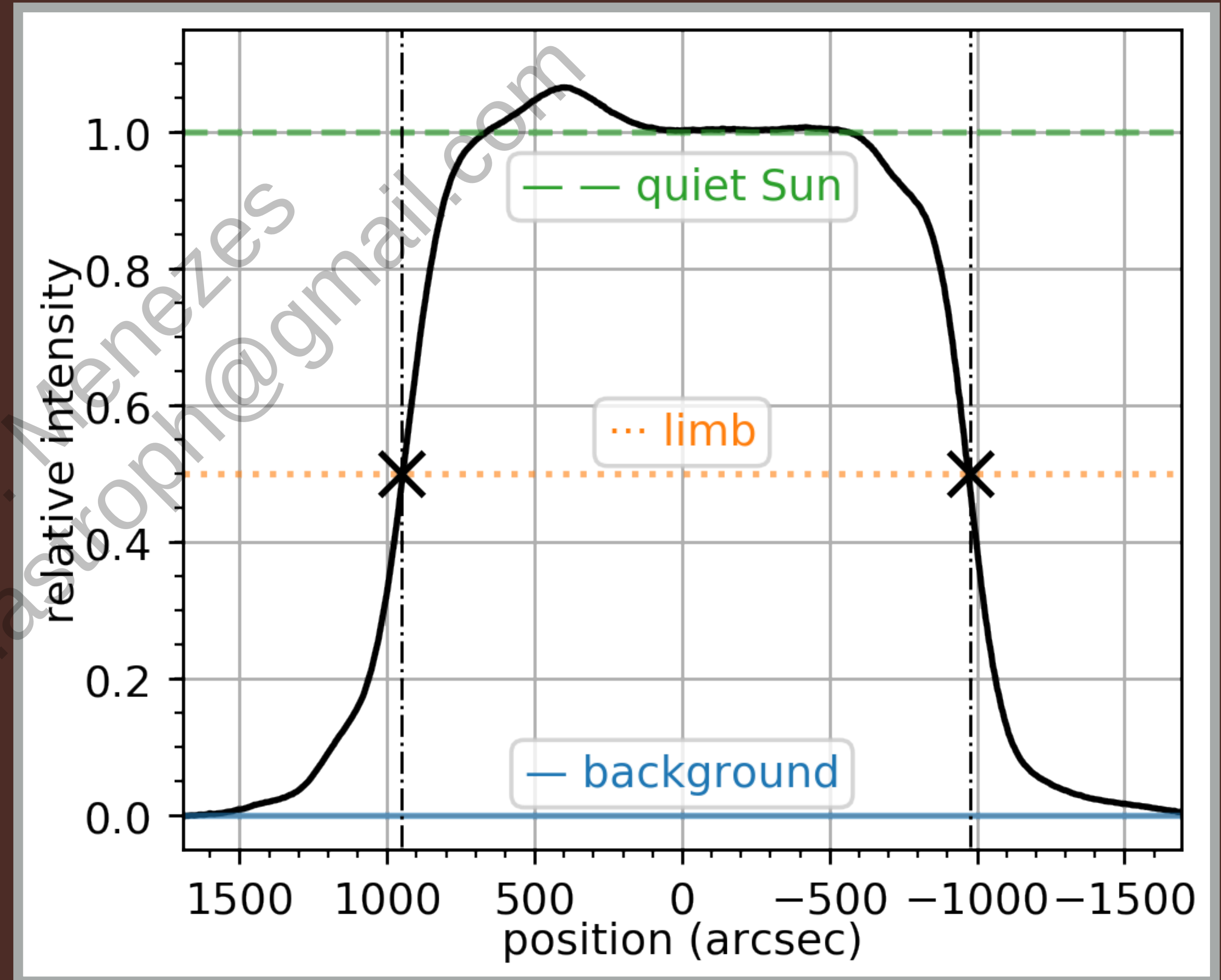
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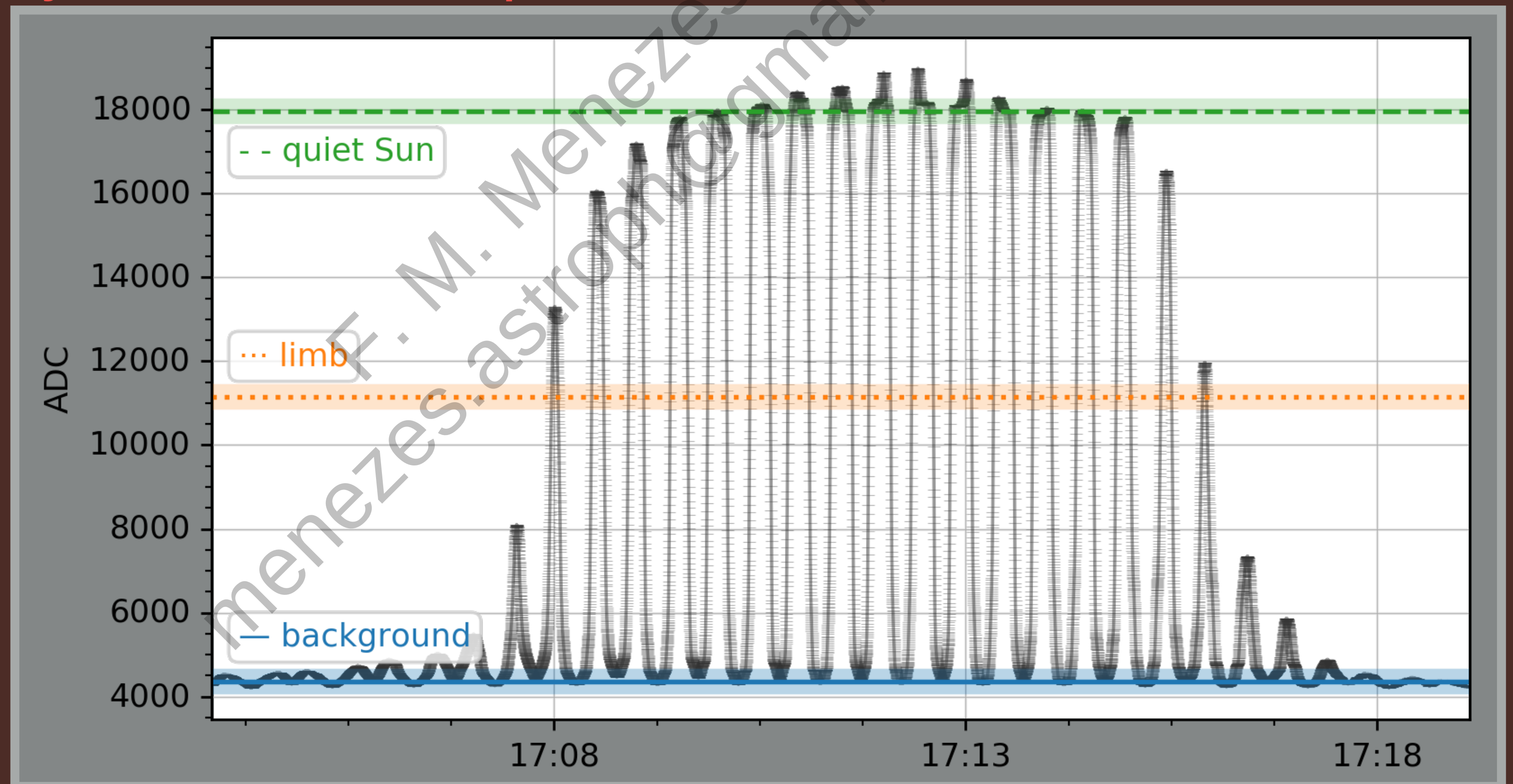
Background, quiet sun & limb levels

- ▶ Get the coordinates corresponding to the limb level
- ▶ Made for every scan of the map



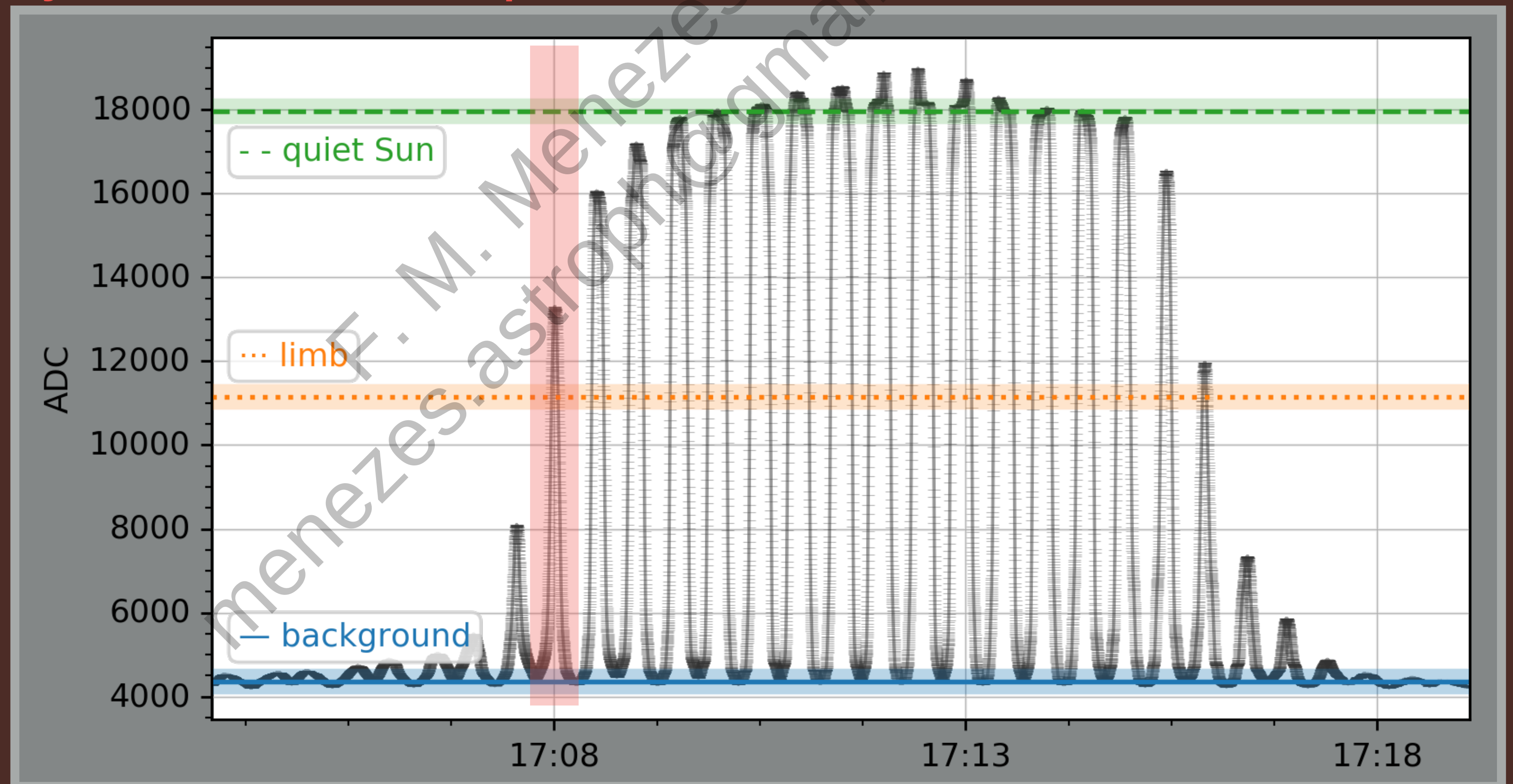
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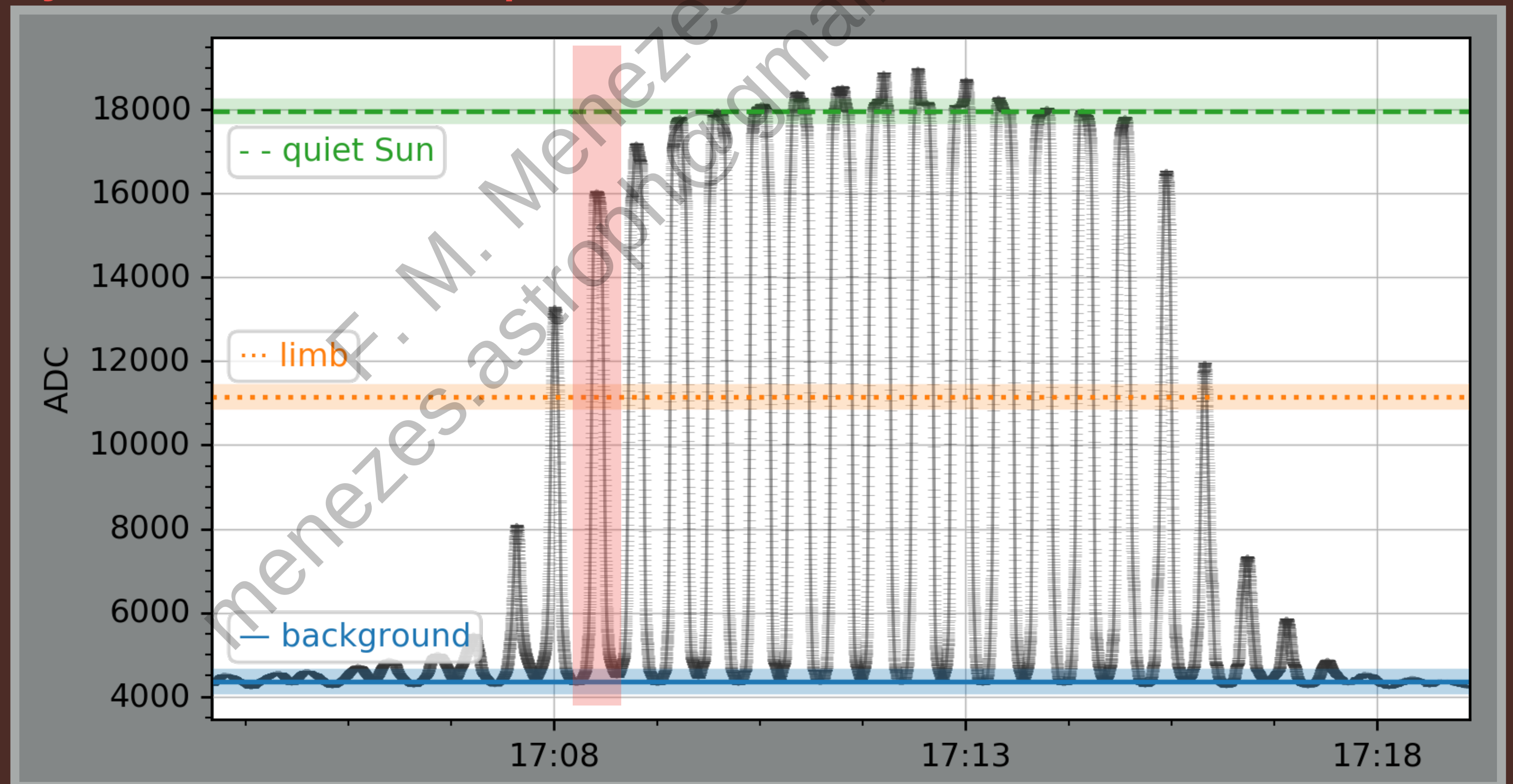
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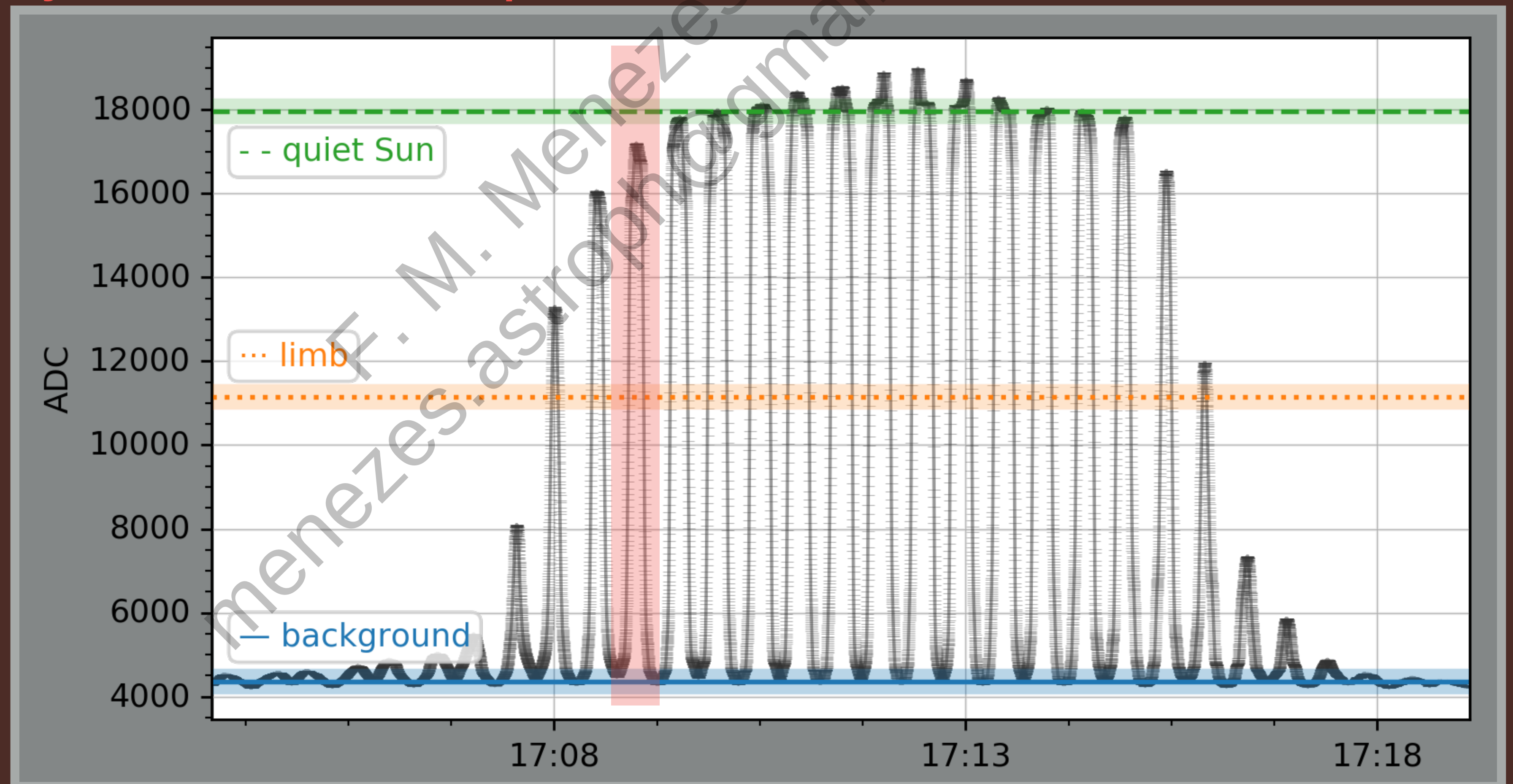
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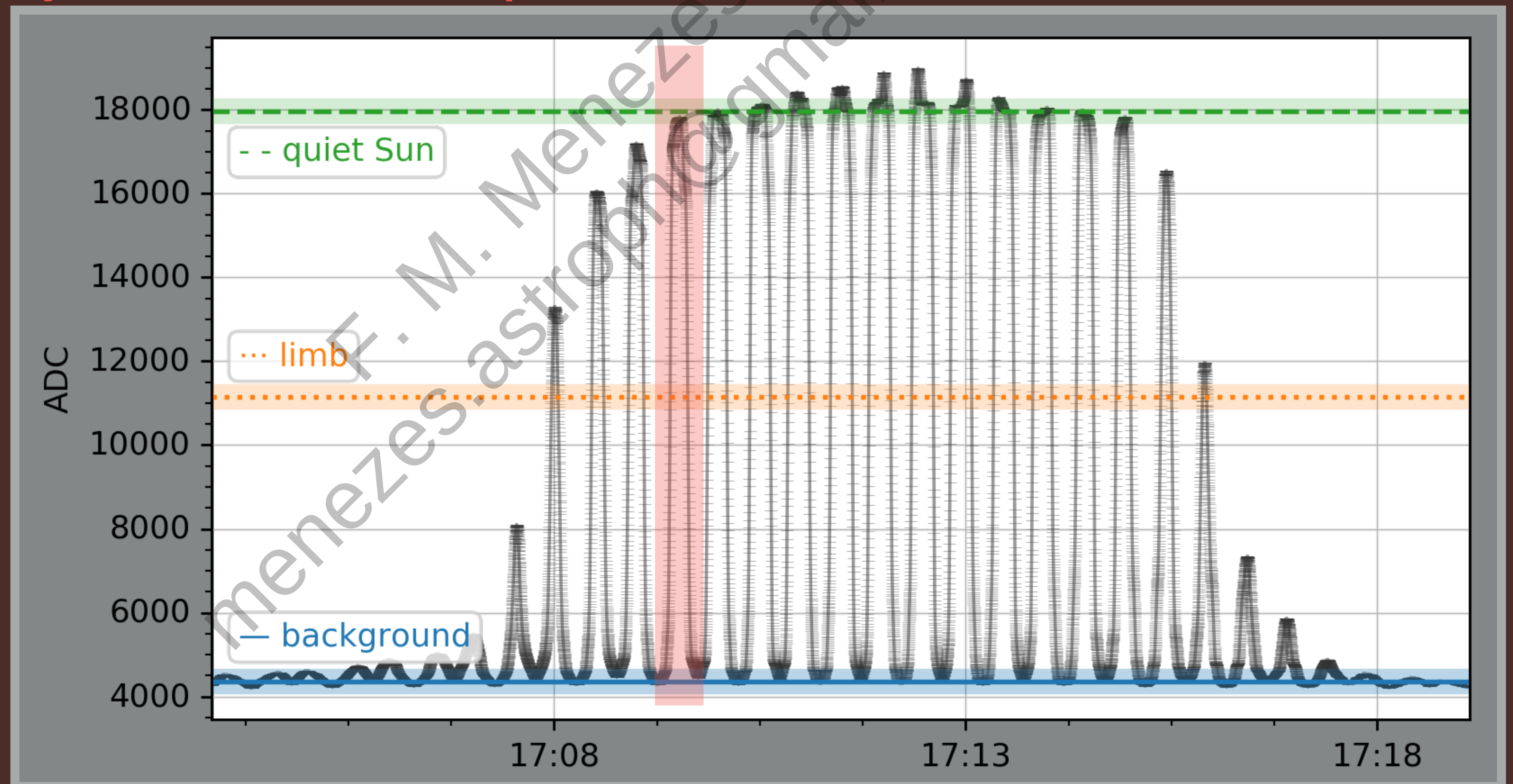
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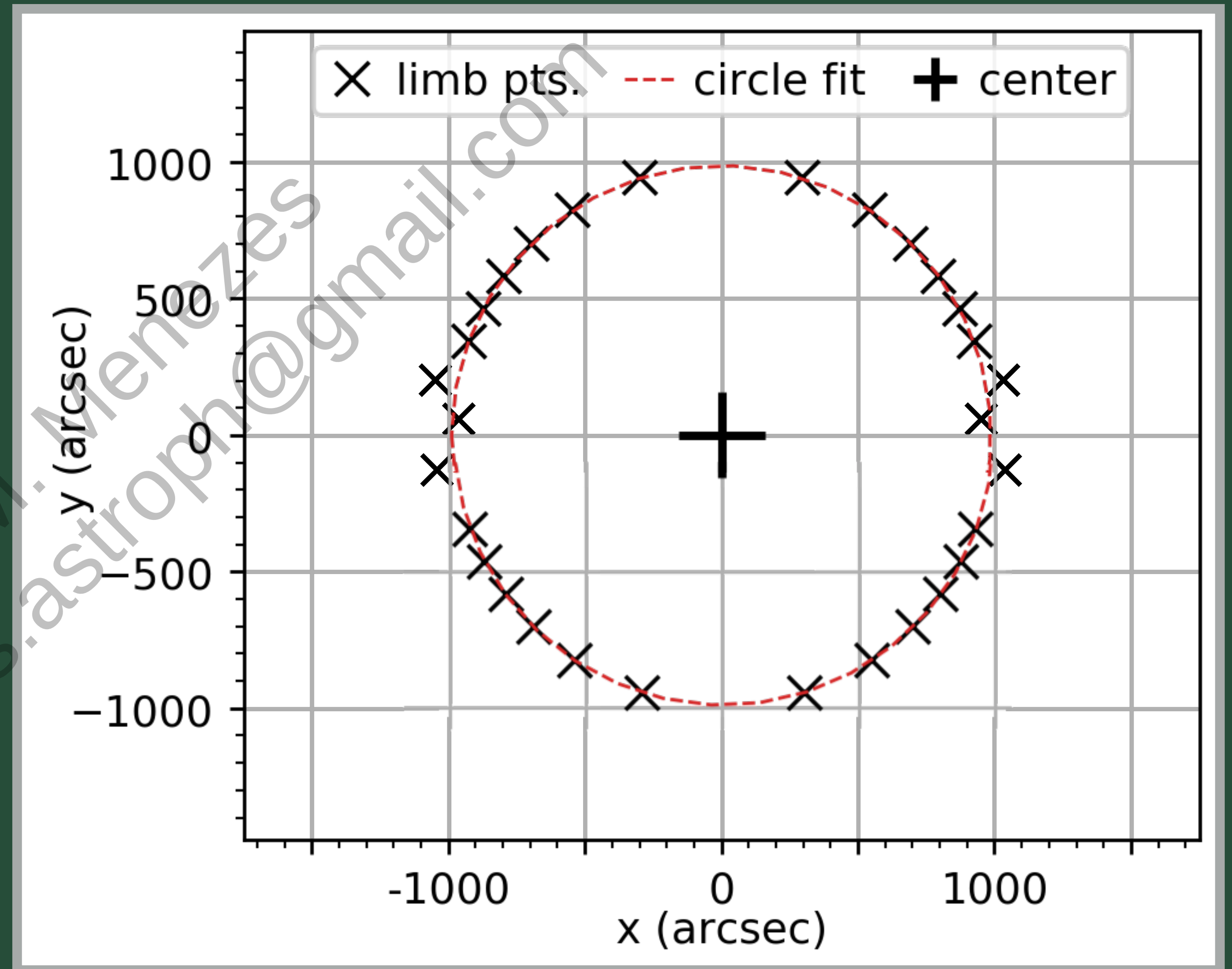
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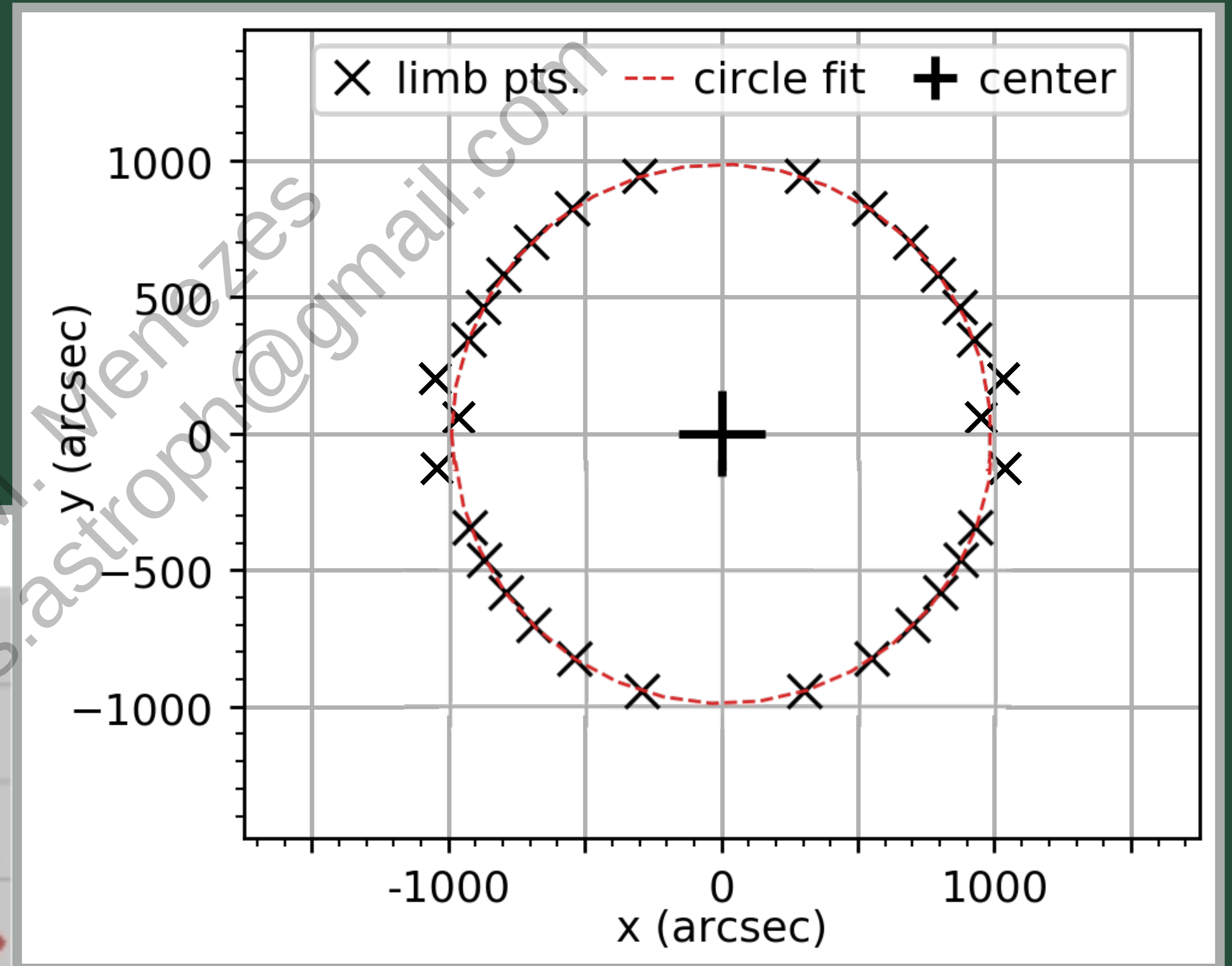
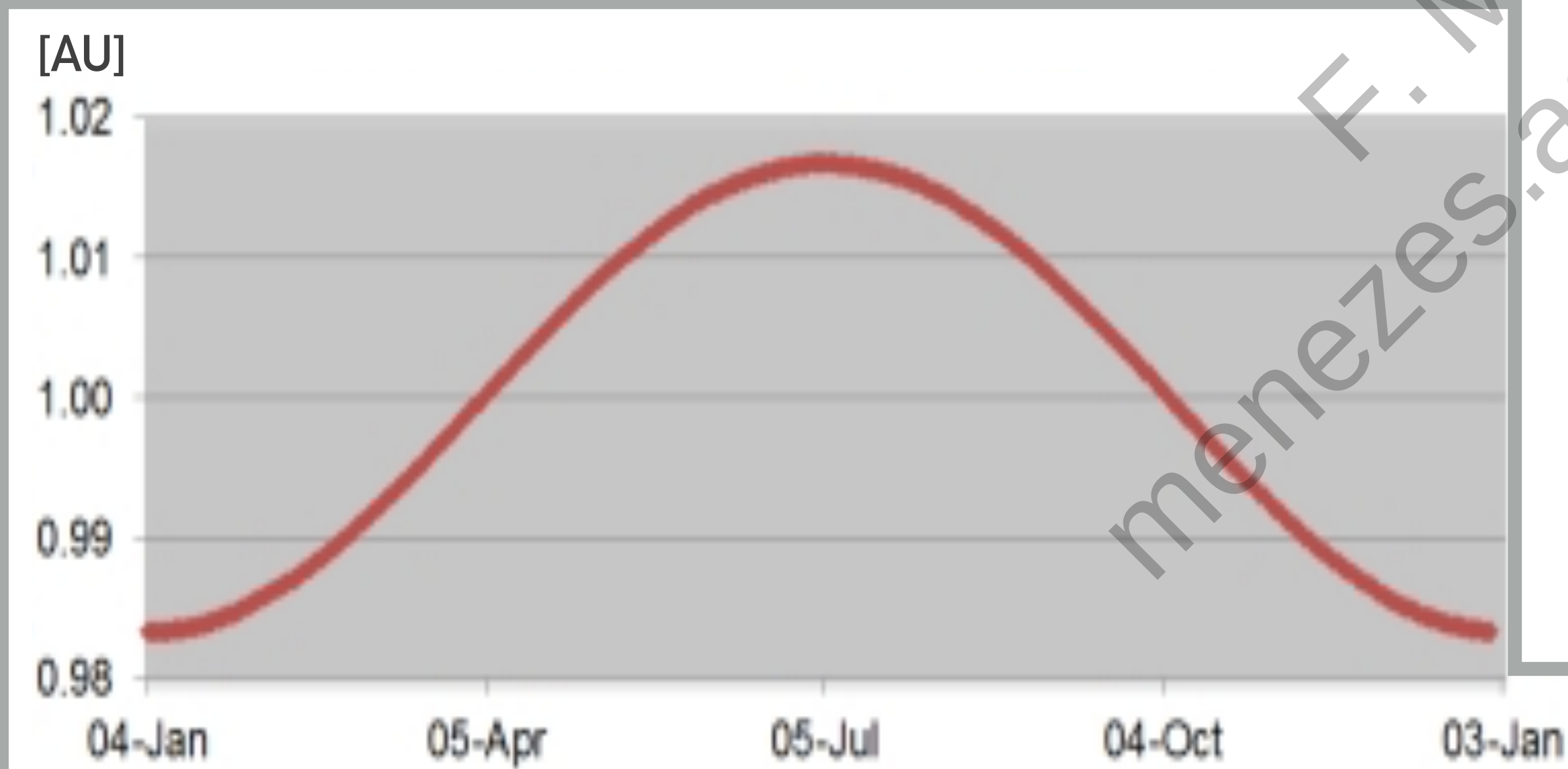
Background, quiet sun & limb levels

- ▶ A circle (red line) is fitted using the limb points
- ▶ Correction due to Earth's orbit eccentricity

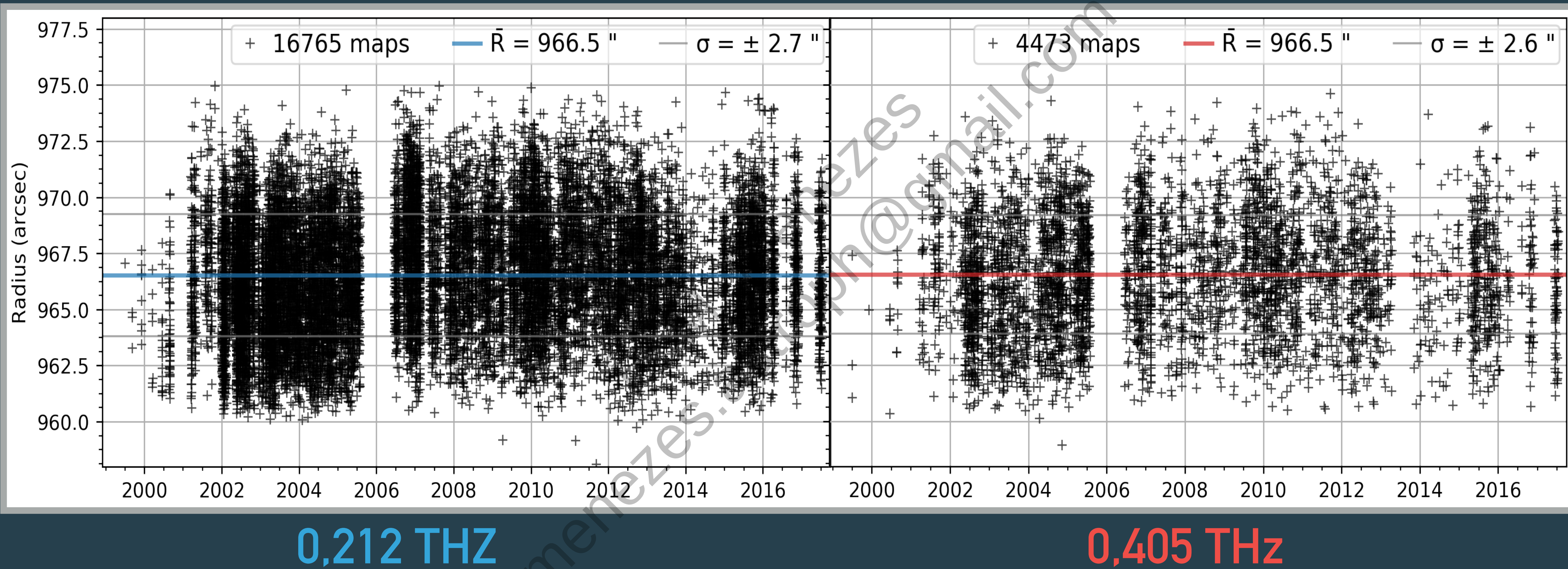


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Results



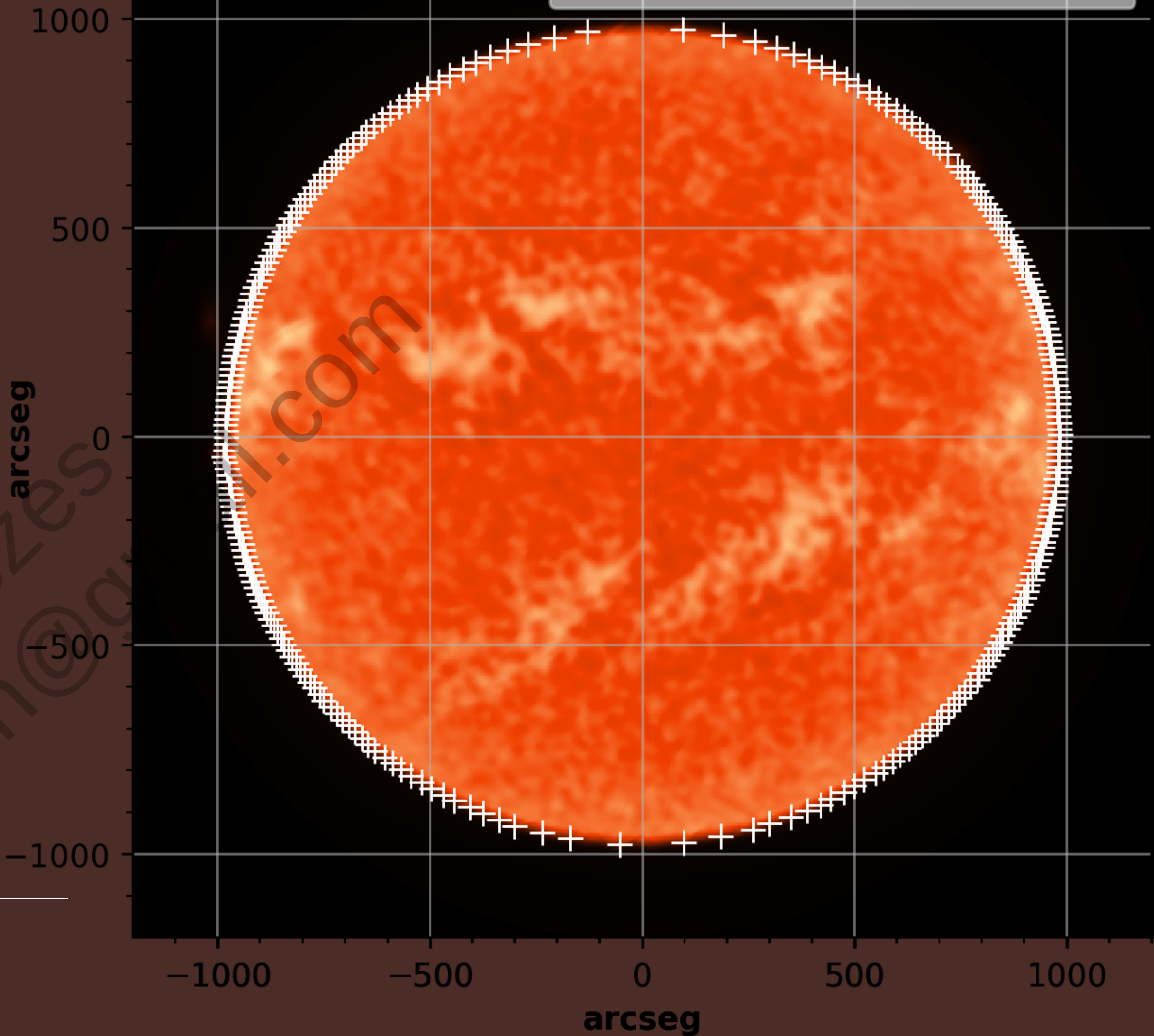
Results

Frequency	Radius (arcsec)	Radius ($R_{\odot}$)	Radius (10^8 m)	Altitude (10^6 m)
212 GHz	966 $\pm$ 2.7	1.007 $\pm$ 0.003	7.01 $\pm$ 0.02	5.0 $\pm$ 2.0
405 GHz	966 $\pm$ 2.6	1.007 $\pm$ 0.003	7.01 $\pm$ 0.02	5.0 $\pm$ 2.0
Optical	959.63	1	6.957	0

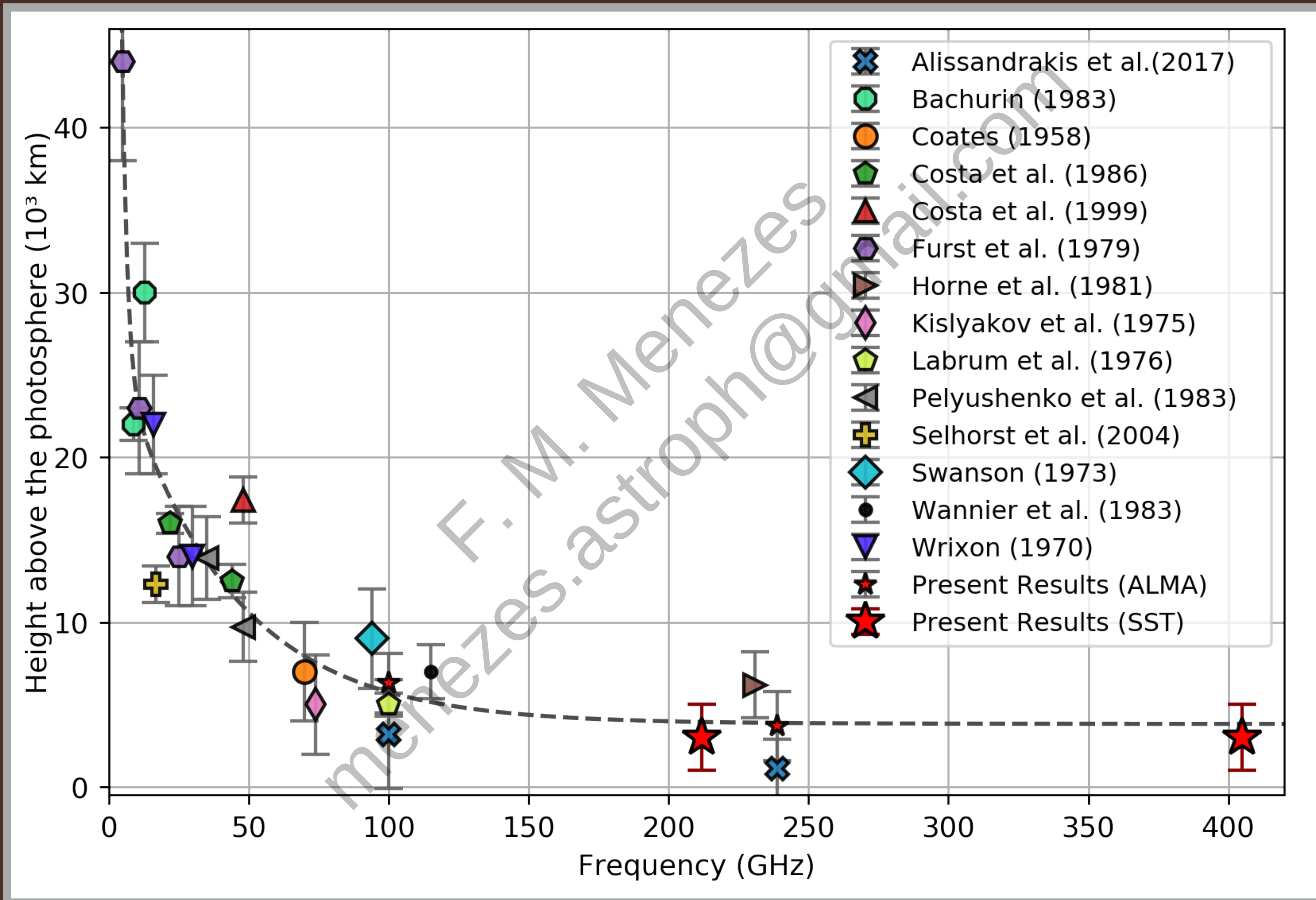
ALMA

- ▶ 0.239 THz
- ▶ 2015 Dec 17

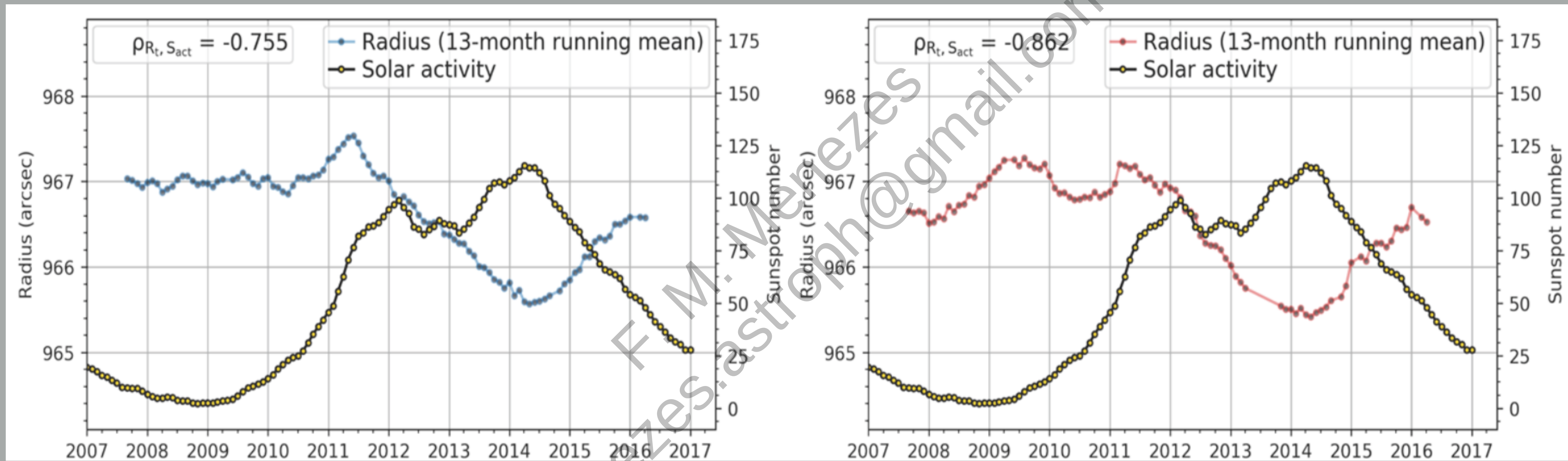
0.2 THz	R (arcsec)	R (10^8 m)	Height (10^6 m)
SST	966 ± 2.7	7.01 ± 0.02	5.0 ± 2.0
ALMA	962.9 ± 1.9	6.980 ± 1.3	2.3 ± 1.3



Results



Correlation to the sunspot cycle

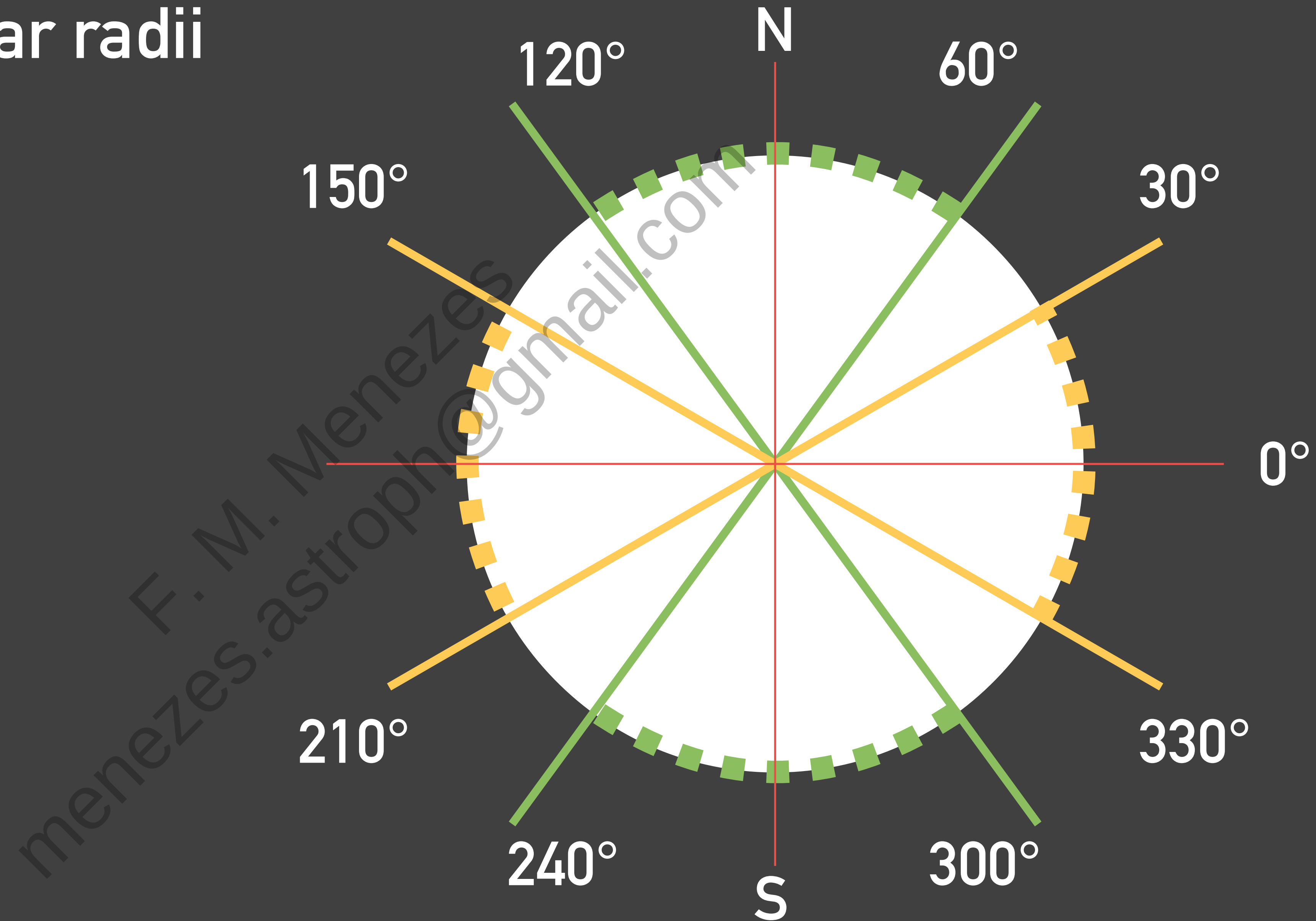


0,212 THZ

0,405 THz

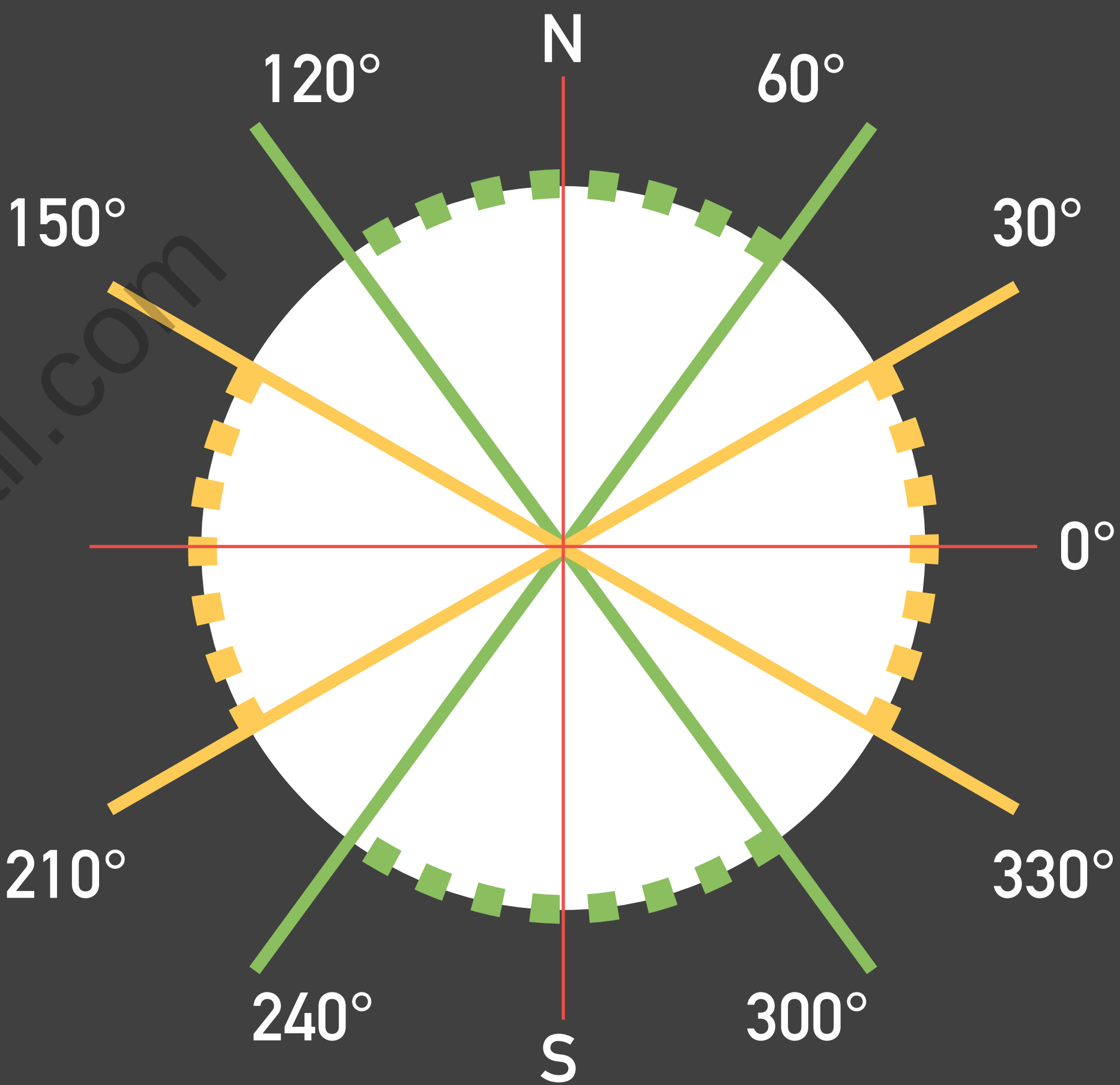
Equatorial and polar radii

- ▶ Equatorial radius
- ▶ Polar radius



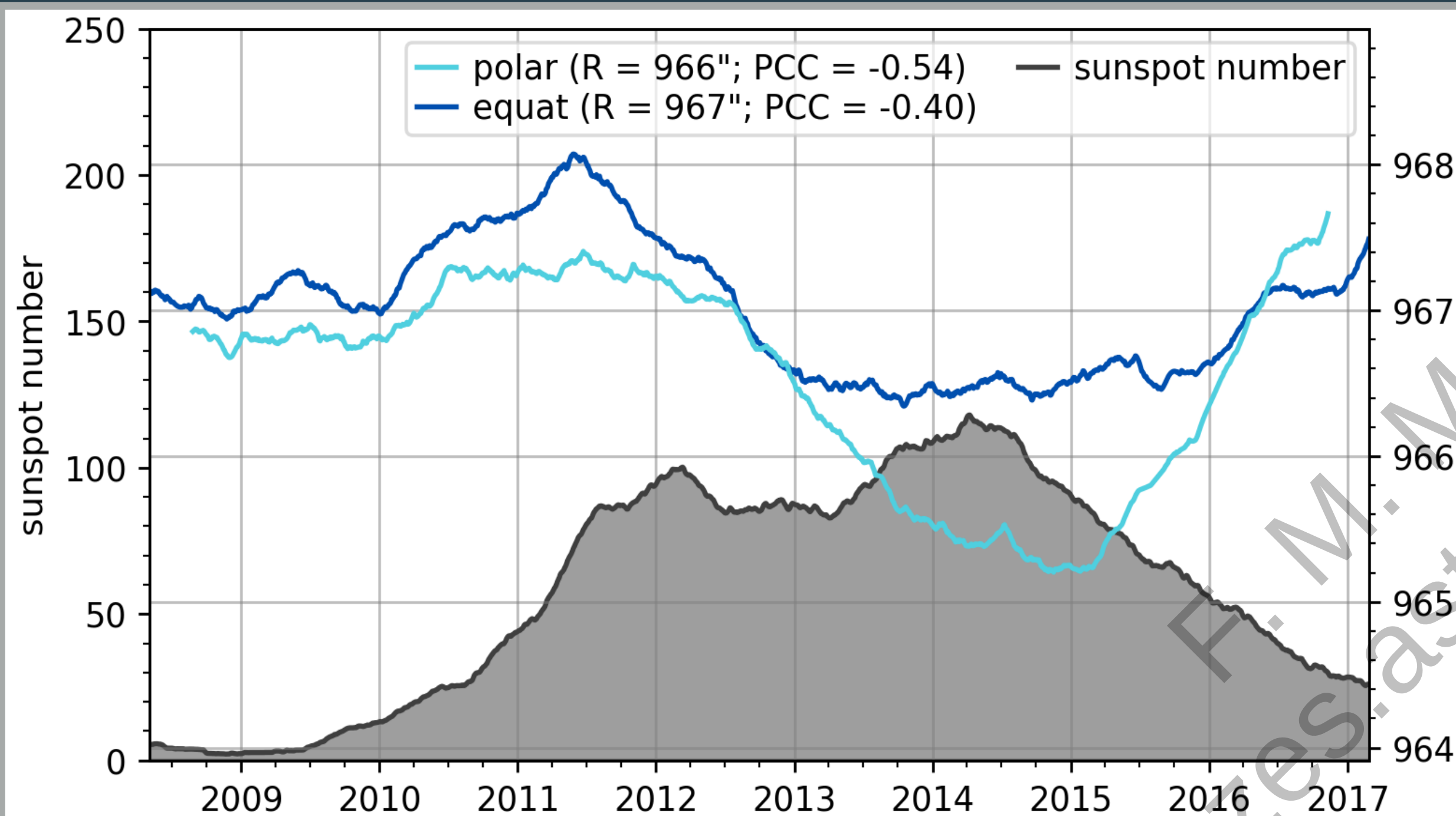
Equatorial and polar radii

R (arcsec)	All pts.	Equat.	Polar
ALMA (0.2 THz)	962.9 ± 1.9	964.2 ± 1.9	962.1 ± 1.1
SST (0.2 THz)	966 ± 2.7	966 ± 3.3	966 ± 3.5
SST (0.4 THz)	966 ± 2.6	967 ± 3.2	966 ± 3.4

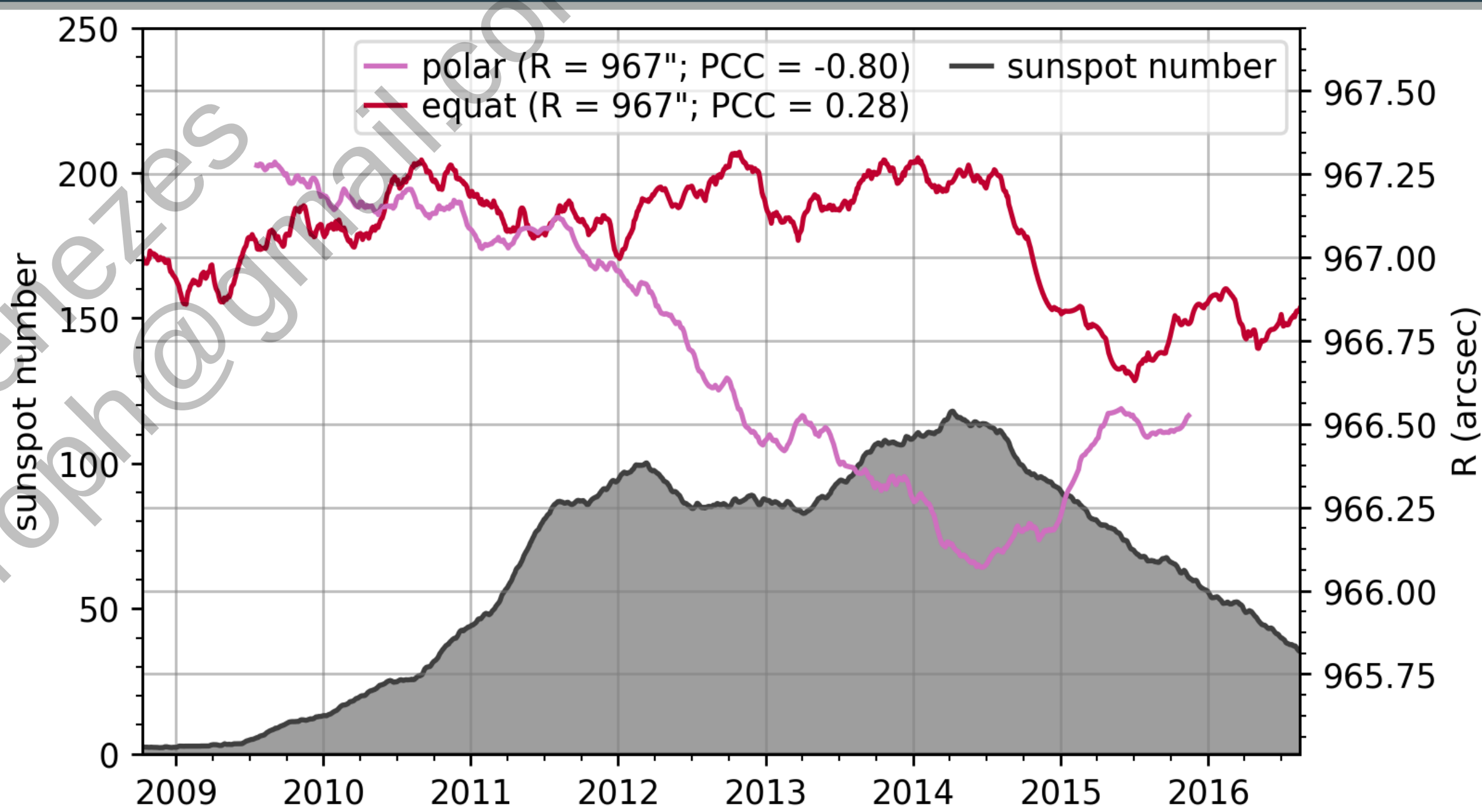


- ▶ Equatorial radius
- ▶ Polar radius

Correlation to the sunspot cycle



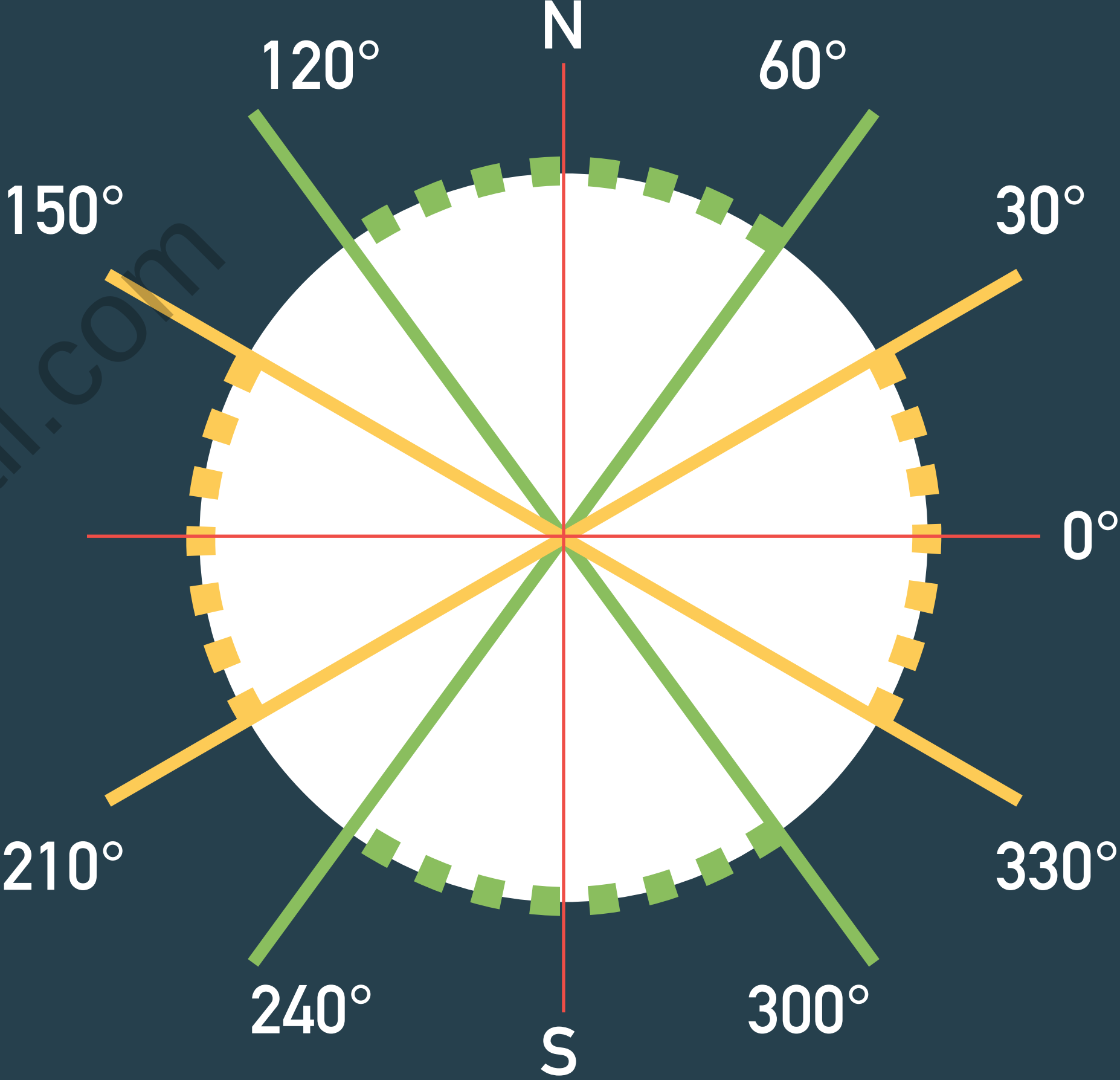
0,212 THz



0,405 THz

Correlation to the sunspot cycle

SST	All pts.	Equat.	Polar
0.2 THz	-0.75	-0.40	-0.54
0.4 THz	-0.86	0.28	-0.80



- ▶ Equatorial radius
- ▶ Polar radius

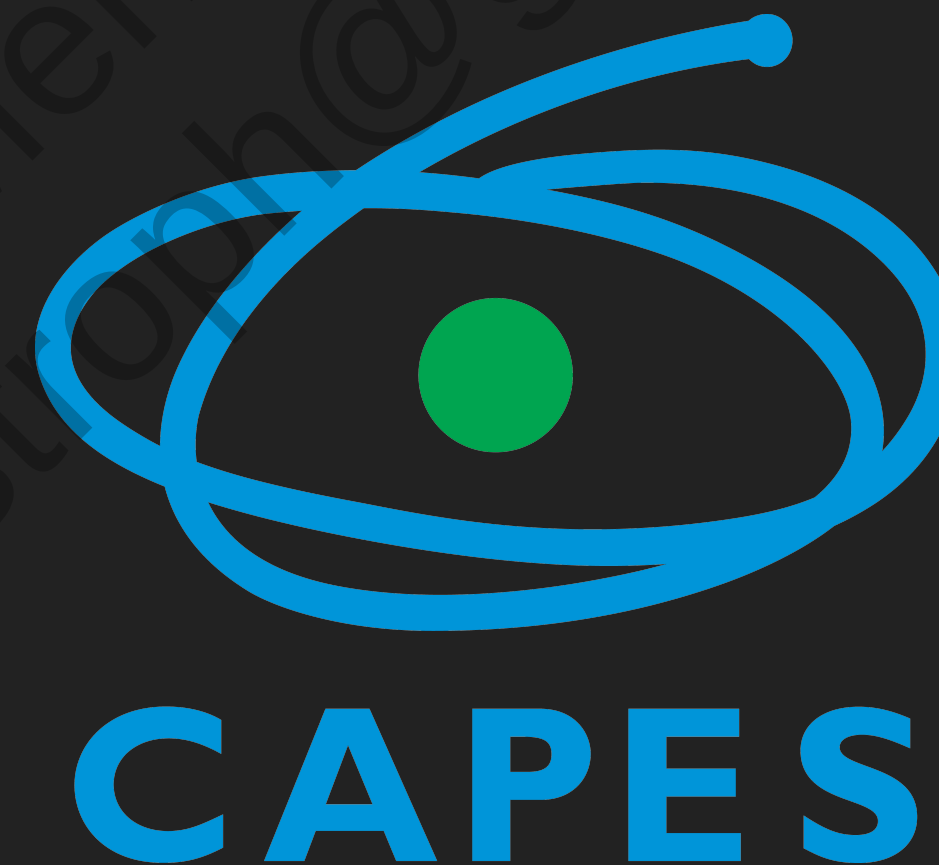
Final remarks

- ▶ Mean solar radii at both frequencies, $R = 966'' \pm 3''$
 - ▶ Between 2500 - 7500 km above photosphere (chromosphere - corona)
 - ▶ Solar atm. models predict these freq. to be at the chromosphere
- ▶ Polar radii at both freq. are anticorr. to sunspot cycle
 - ▶ At 17 GHz there are a significant increase of polar limb brightening during solar minimum (Selhorst, 2004)
 - ▶ High solar activity is associated with increased magnetic fields, which would lead to a reduced energy flux transported by convection (Gilliland, 1981).
- ▶ 0.405 THz equat. radius weakly correlated to sunspot cycle
- ▶ 0.212 THz equat. radius anticorrelated to sunspot cycle

Aknowledgements



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